



# A New Charged Lepton Flavor Violation Experiment: Muon-Electron Conversion at Fermilab

R. Bernstein  
Fermilab  
for the Mu2e Collaboration



Boston University

Brookhaven National Laboratory

University of California, Berkeley

University of California, Irvine

California Institute of Technology

City University of New York

Duke University

Fermilab

University of Houston

University of Illinois, Urbana-Champaign

Lawrence Berkeley National Laboratory

Los Alamos National Laboratory

Northwestern University

Rice University

Syracuse University

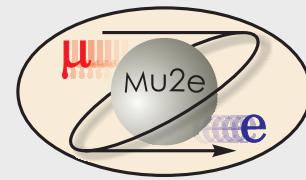
University of Virginia

College of William and Mary

University of Washington, Seattle

R. Bernstein, FNAL

# Collaboration



Istituto G. Marconi Roma

Laboratori Nazionali di Frascati

Università di Pisa, Pisa

Università del Salento

Gruppo Collegato di Udine



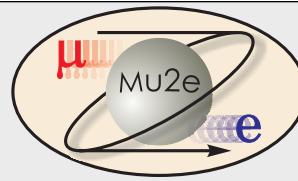
Institute for Nuclear Research, Moscow, Russia

JINR, Dubna, Russia

~130 collaborators



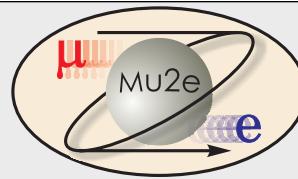
# Outline



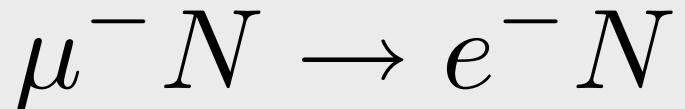
- The search for muon-electron conversion
- Experimental Technique
- Fermilab Accelerator
- Project X Upgrades and Mu2e
- Cost and Schedule
- Conclusions



# What is $\mu e$ Conversion?



muon converts to electron in the presence of a nucleus



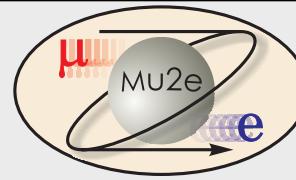
$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A,Z) \rightarrow e^- + N(A,Z))}{\Gamma(\mu^- + N(A,Z) \rightarrow \text{all muon captures})}$$

- Charged Lepton Flavor Violation (CLFV)
  - will measure  $R_{\mu e} < 6 \times 10^{-17}$  @ 90% CL
  - Related Processes:

$\mu$  or  $\tau \rightarrow e\gamma$ ,  $e^+e^-e$ ,  $K_L \rightarrow \mu e$ , and more

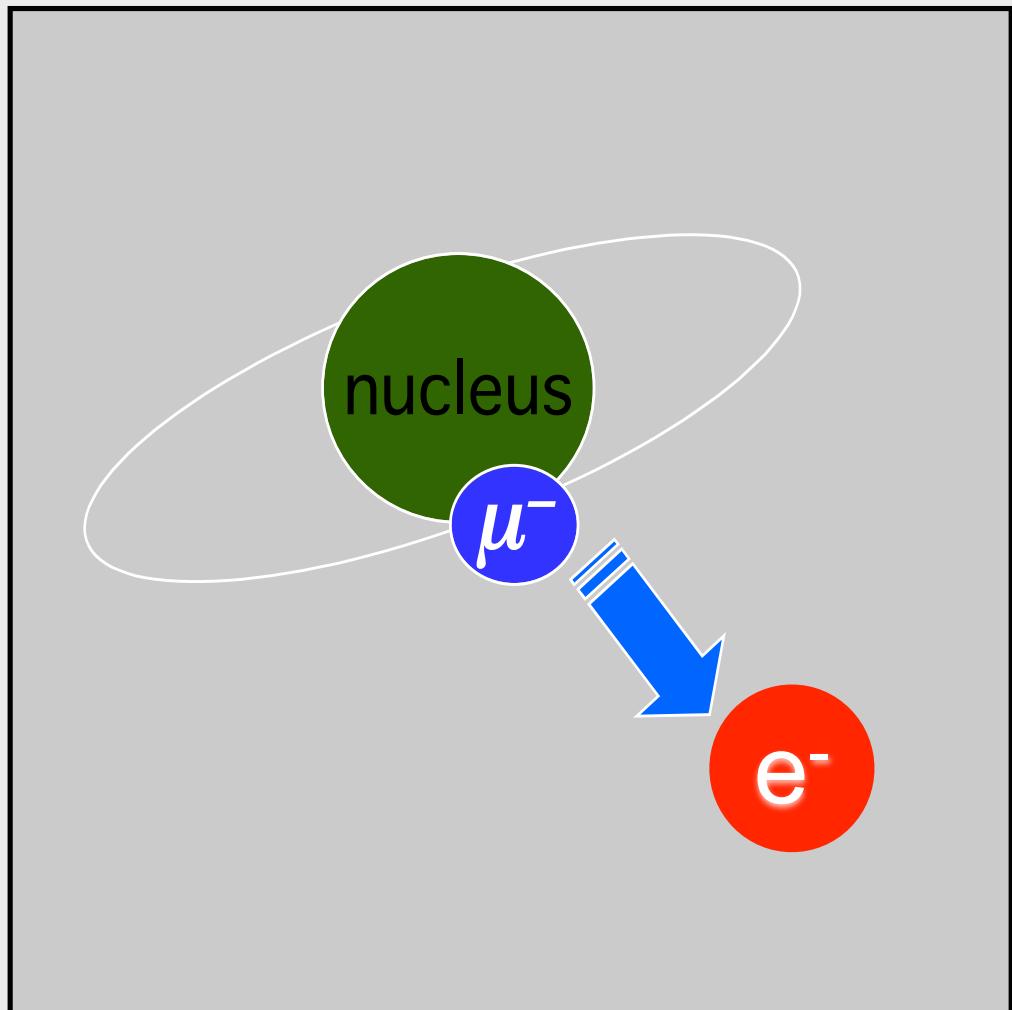


# Experimental Signal



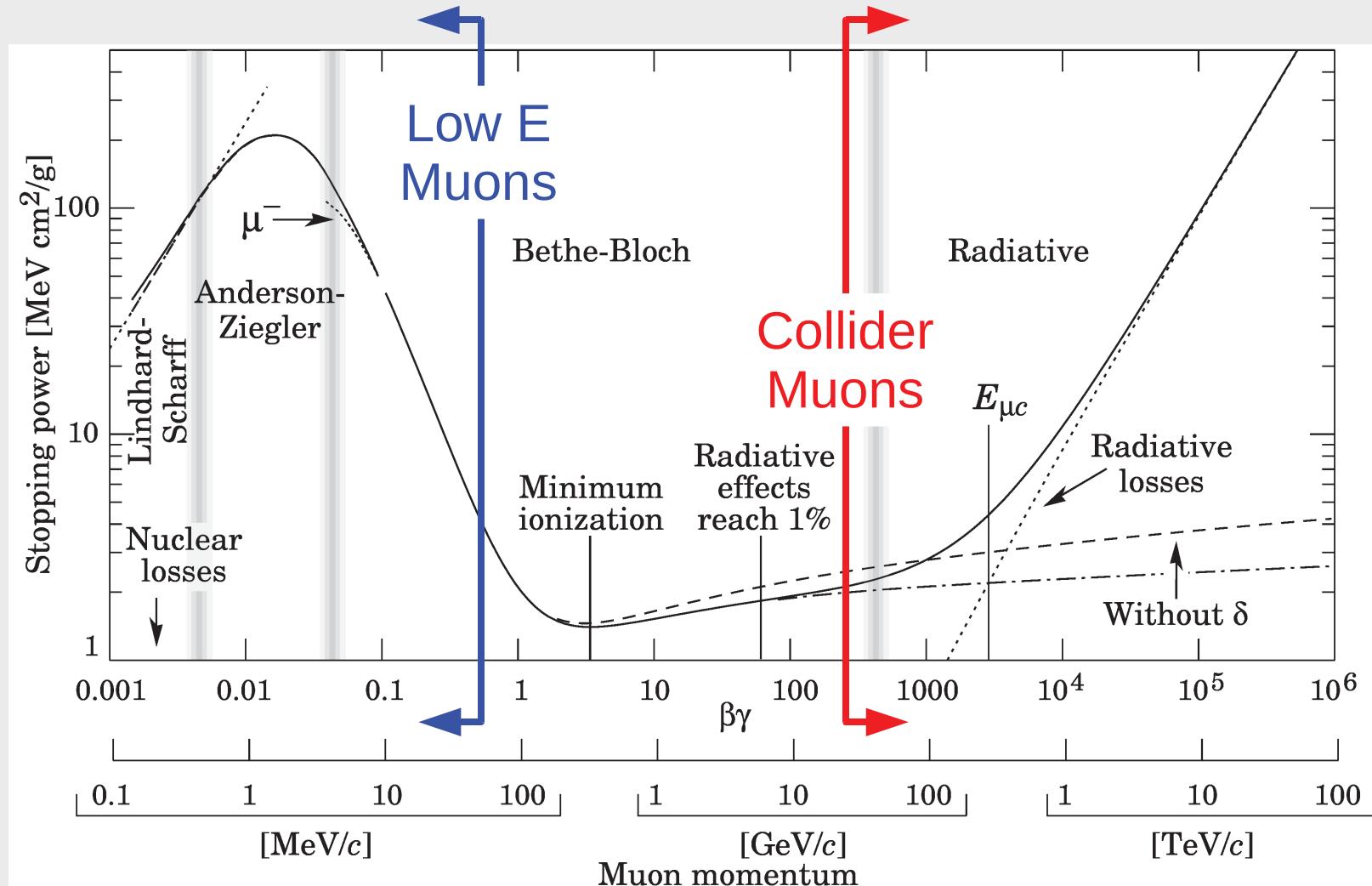
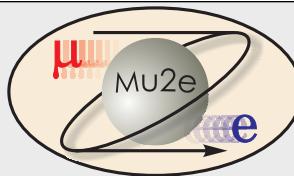
$$\mu^- N \rightarrow e^- N$$

- A Single Monoenergetic Electron
- If  $N = \text{Al}$ ,  $E_e = 105.$  MeV
  - electron energy depends on  $Z$
- Nucleus coherently recoils off outgoing electron, no breakup



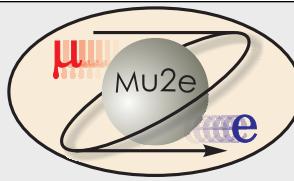


# Physics IS Different at these energies!





# “Who ordered that?”



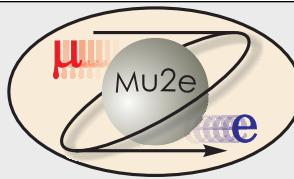
– I.I. Rabi, 1937

After the  $\mu$  was discovered, it was logical to think the  $\mu$  is just an excited electron:

- expect  $\text{BR}(\mu \rightarrow e\gamma) \approx 10^{-4}$
  - Unless another  $\nu$ , in Intermediate Vector Boson Loop, cancels (Feinberg, 1958)
- ➡ same as GIM mechanism!



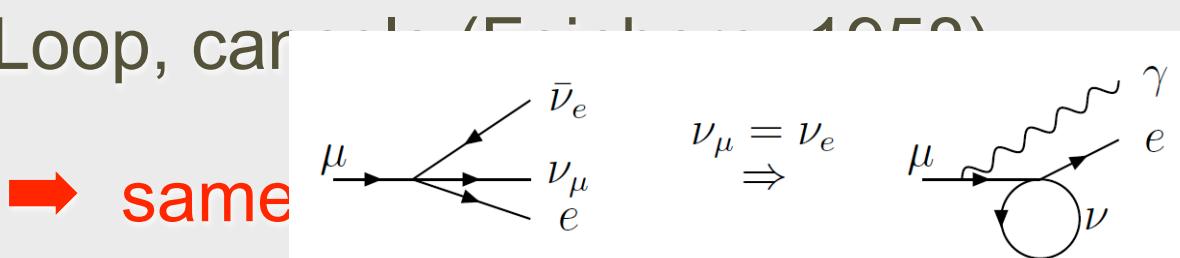
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– I.I. Rabi, 1937

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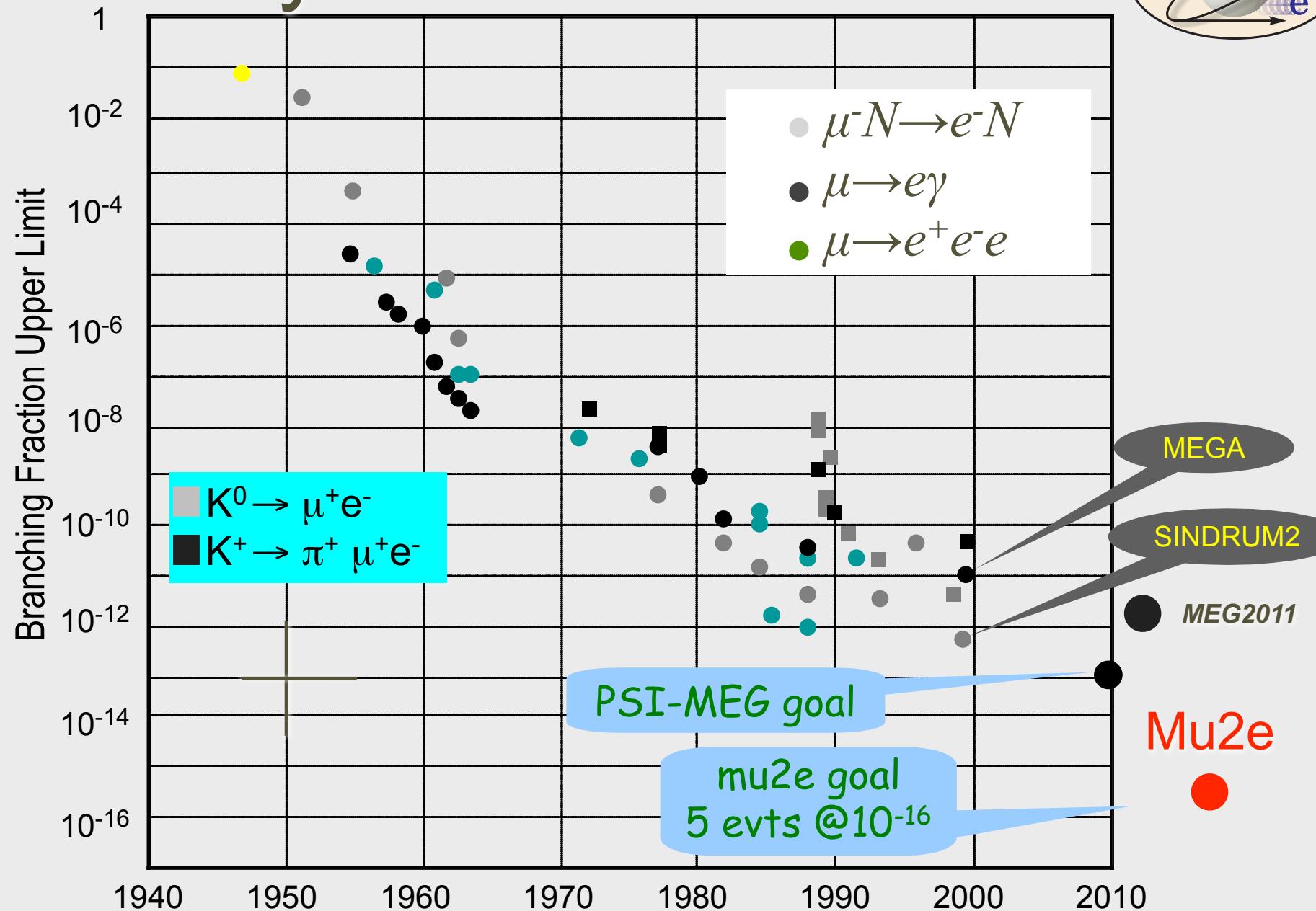
- expect  $\text{BR}(\mu \rightarrow e\gamma) \approx 10^{-4}$
- Unless another  $\nu$ , in Intermediate Vector Boson Loop, car



<sup>1</sup>Unless we are willing to give up the 2-component neutrino theory, we know that  $\mu \rightarrow e + \nu + \bar{\nu}$ .

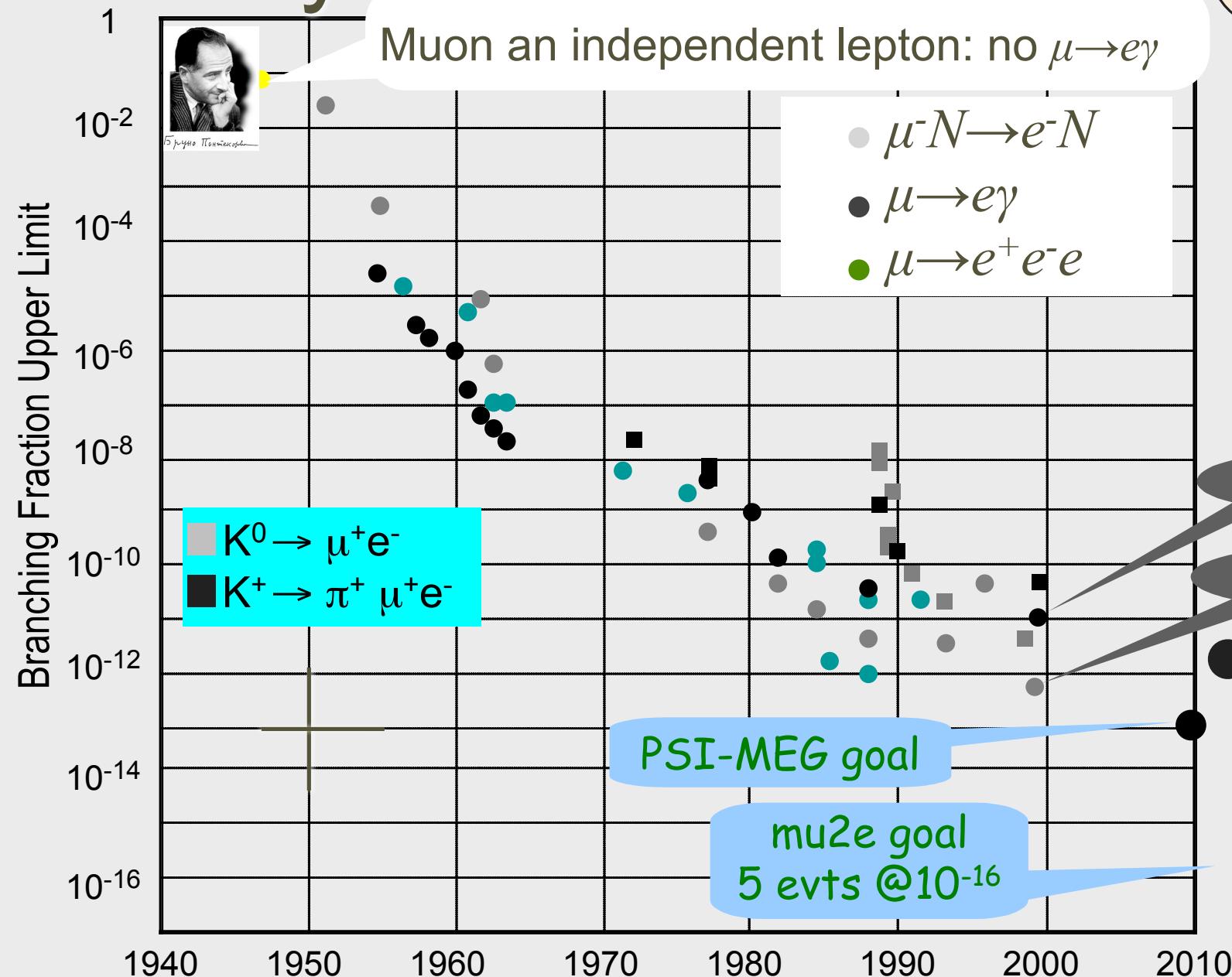


# History of CLFV Searches



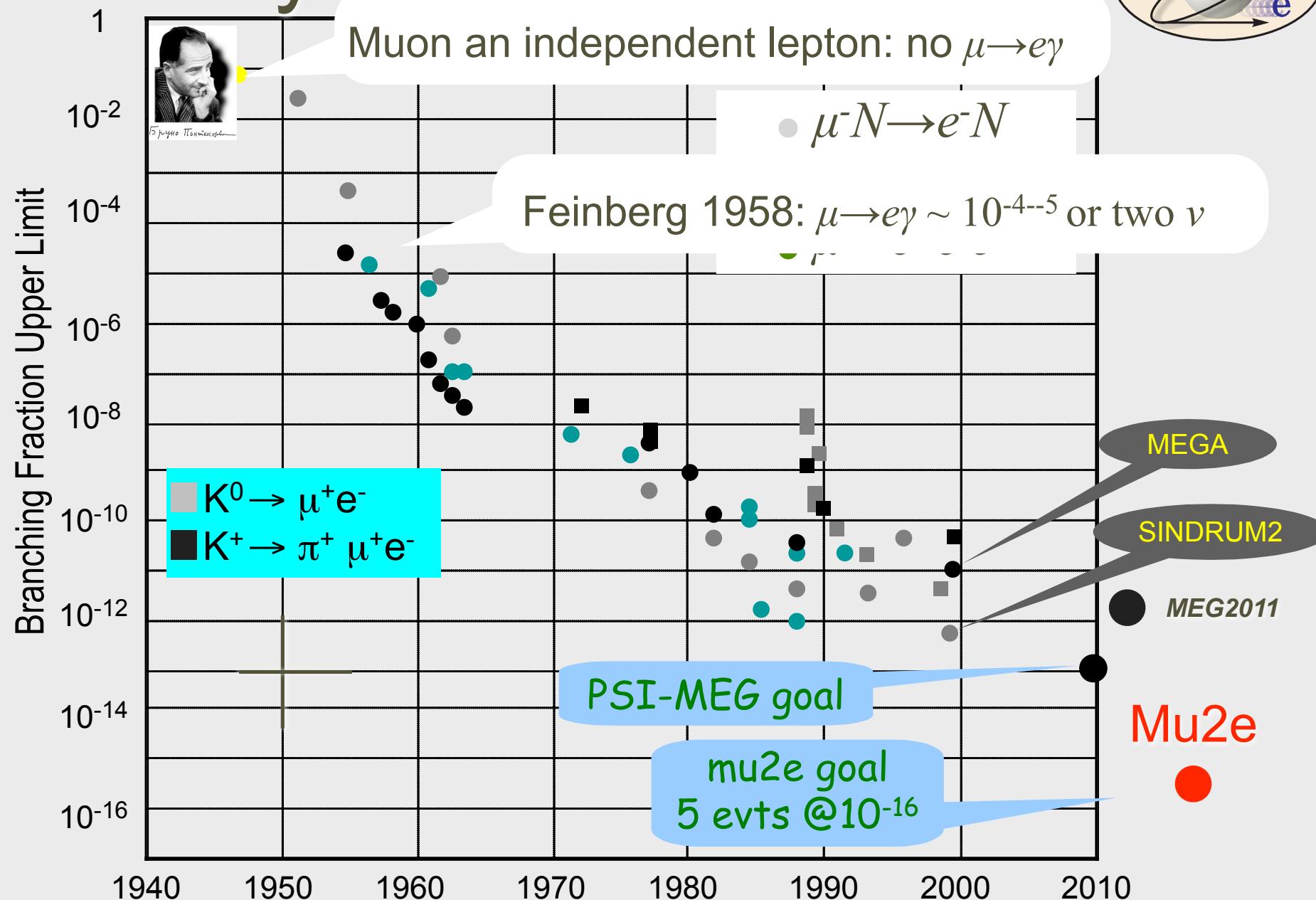


# History of CI FV Searches



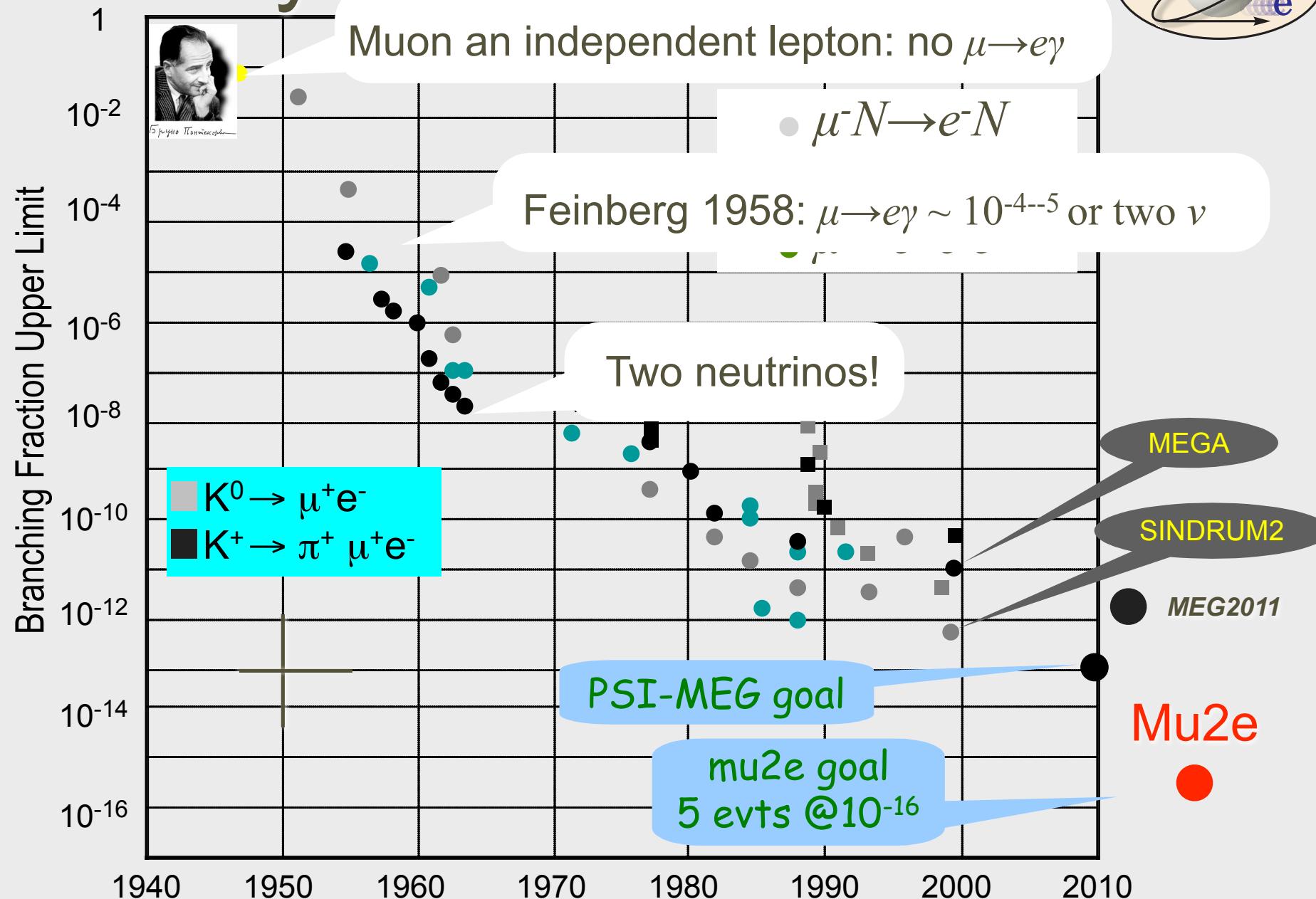
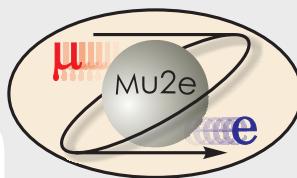


# History of CI FV Searches



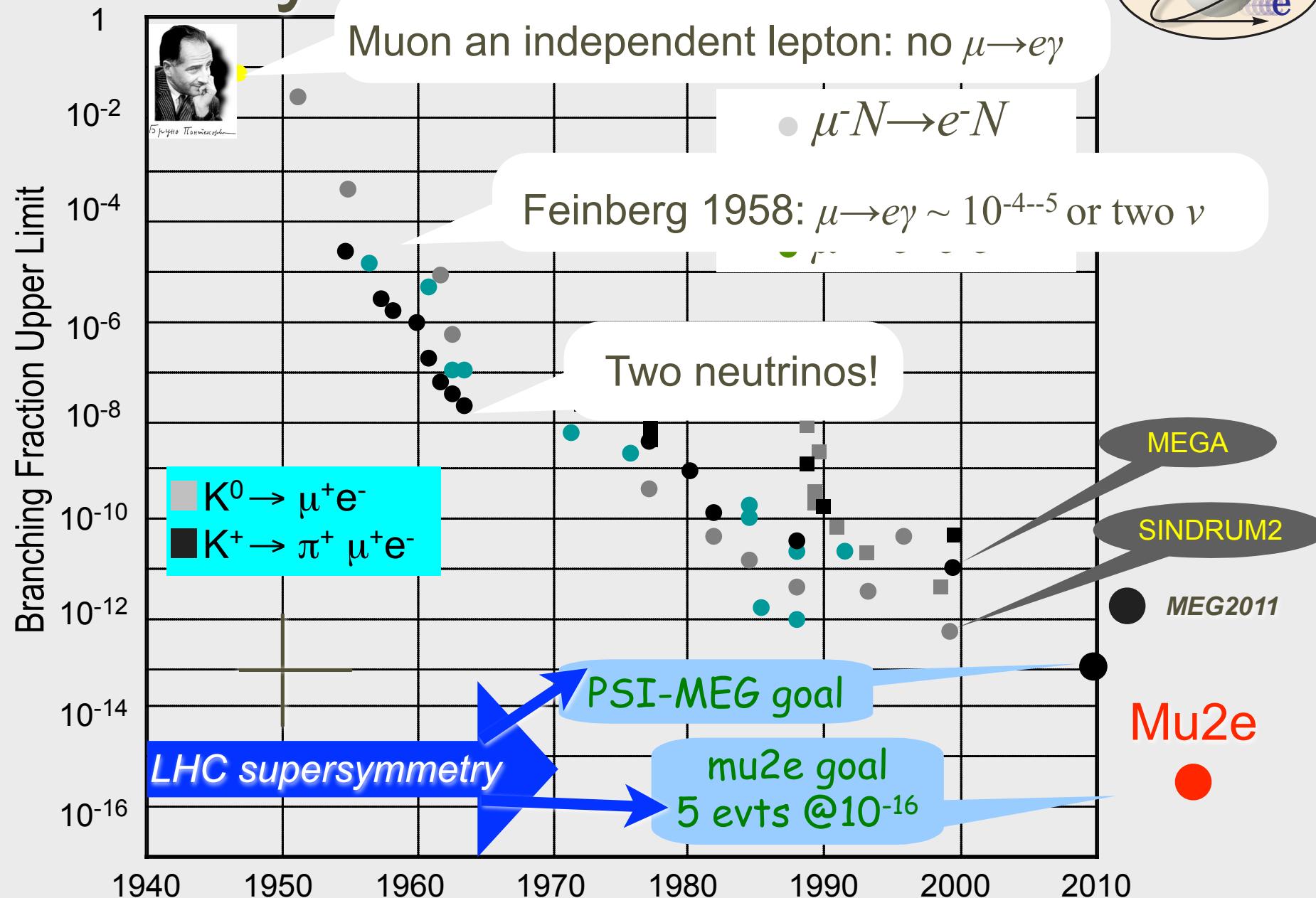


# History of CI FV Searches



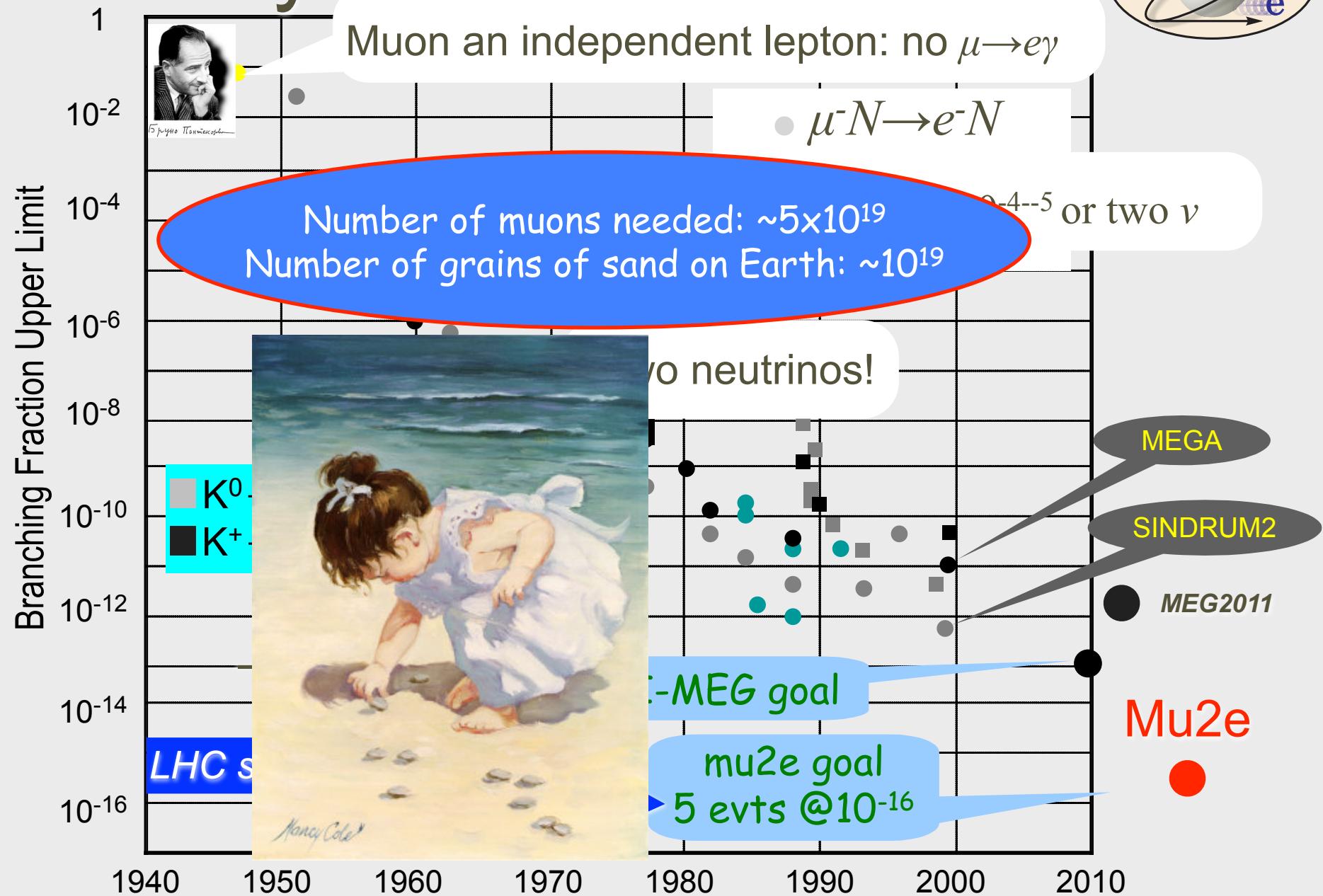


# History of CI FV Searches



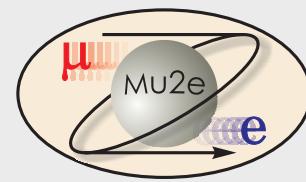


# History of CI FV Searches





# Where did that last estimate come from?



<http://www.hawaii.edu/suremath/jsand.html>

Moral: jobs at Hawaii are more fun than at Fermilab

## Estimate the number of grains of sand on all the beaches of the earth.

The number of grains of sand will be the volume of sand divided by the volume per grain.

$$N = \frac{\text{Volume\_of\_sand}}{\text{Volume\_of\_grain}}$$

The volume of sand is available from the product of the length of all beaches in the world, their average width and average depth.

$$\text{Volume\_of\_sand} = (\text{length\_of\_beaches}) (\text{width}) (\text{depth})$$

We will need to estimate the parameters on the right.

To get going we use the idea that the length of beaches is some percentage of the length of shores.

$$\text{length\_of\_beaches} = \frac{\%shores}{100} \text{length\_of\_shores}$$

The length of shores is related to the size of the earth. We choose to include the size in terms of the circumference.

$$\text{length\_of\_shores} = (\text{multiple\_of\_circumference}) (\text{earth\_circumference})$$

$$N = \frac{\left( \frac{\%shores}{100} ((\text{multiple\_of\_circumference}) \text{earth\_circumference}) \right) (\text{width}) (\text{depth})}{\text{Volume\_of\_grain}}$$

We use rough but reasonable values for the parameters.

$$\%shores = 25$$

$$\text{multiple\_of\_circumference} = 5$$

$$\text{earth\_circumference} = 40000 \text{ km} \quad \text{km} = 1000 \text{ m} \quad \text{m} = 1000 \text{ mm}$$

$$\text{width} = 30 \text{ m}$$

$$\text{depth} = 5 \text{ m}$$

$$\text{Volume\_of\_grain} = 1 \text{ mm}^3$$

$$N = \frac{\left( \frac{25}{100} ((5) (40000 (1000 (1000 \text{ mm})))) \right) (30 (1000 \text{ mm})) (5 (1000 \text{ mm}))}{1 \text{ mm}^3}$$

$$N = 7.5 \times 10^{18}$$

The numbers used can readily be varied to introduce more refined estimates of the parameters.



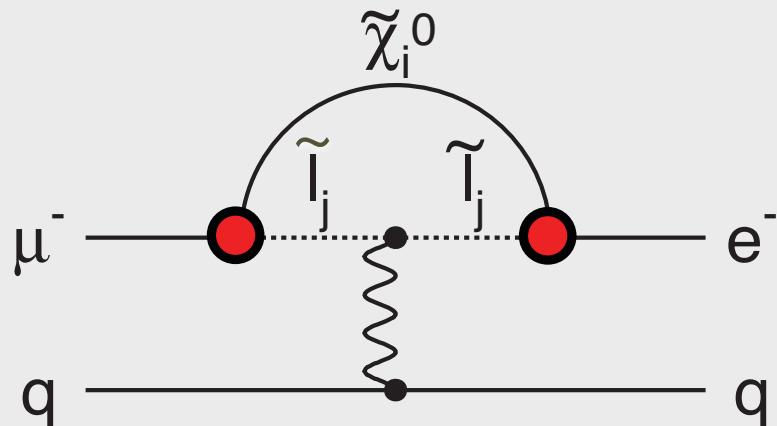


# LFV, SUSY and the LHC



## Supersymmetry

rate  $\sim 10^{-15}$

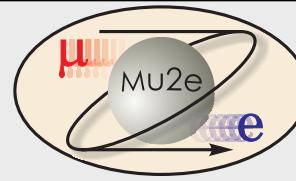


***Access SUSY  
through loops:***

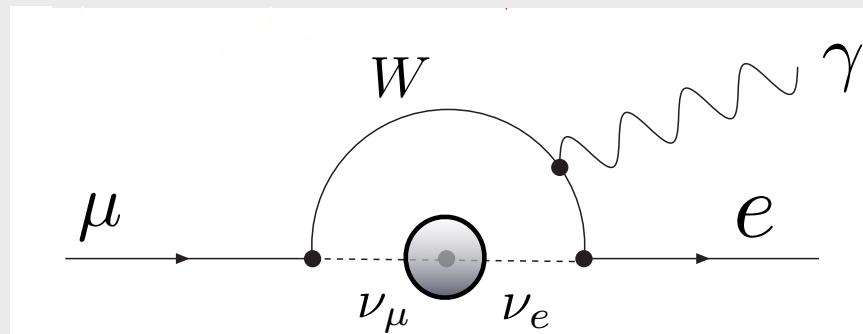
***signal of  
Terascale at LHC  
implies  
~40 event signal /  
<0.2 bkg in this  
experiment***



# Neutrino Oscillations and Muon-Electron Conversion



- ν's have mass! *individual lepton numbers are not conserved*
- Therefore Lepton Flavor Violation occurs in Charged Leptons as well



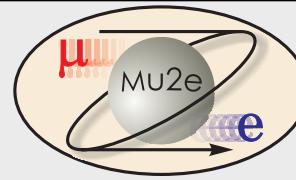
NO STANDARD  
MODEL  
BACKGROUND!

$$\text{BR}(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{e i} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 < 10^{-54}$$



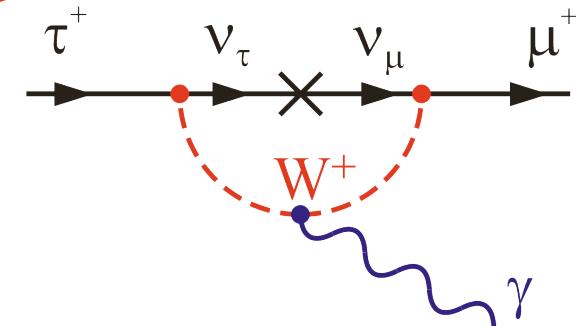


# CLFV and Tau Decays



Lee, Shrock

Phys.Rev.D16:1444,1977

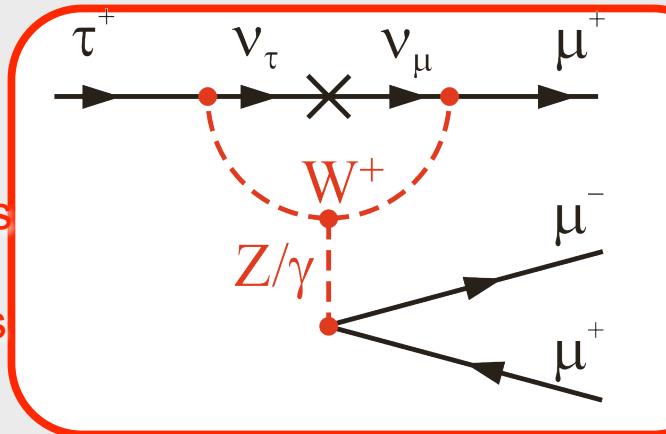


$SM \sim 10^{-40}$

Good News:  
Beyond SM rates are  
several orders of  
magnitude larger than in  
associated muon decays

but less:

Smaller  
GIM  
Cancellations  
because of  
large  $\tau$  mass



$SM \sim 10^{-14}$

Bad News:

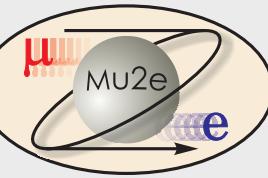
$\tau$ 's hard to produce:  
 $\sim 10^{10} \tau/\text{yr}$  vs  $\sim 10^{11} \mu/\text{sec}$  in  
fixed-target experiments  
(Mu2e/COMET)

also  $e \rightarrow \tau$  at electron-ion collider?

[M. Gonderinger](#), [M. Ramsey-Musolf](#), arXiv:1006.5063v1



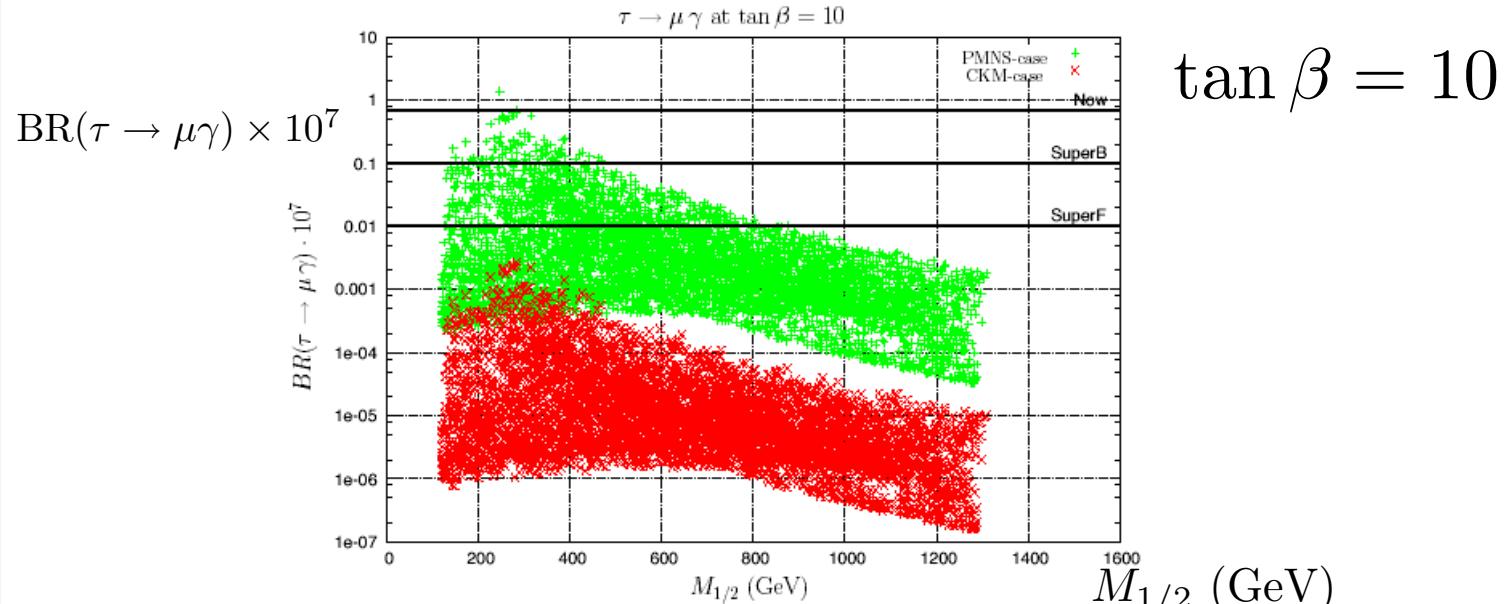
# Supersymmetry in Tau LFV



L. Calibbi, A. Faccia, A. Masiero, S. Vempati hep-ph/0605139

Neutrino-Matrix Like (PMNS)

Minimal Flavor Violation(CKM)

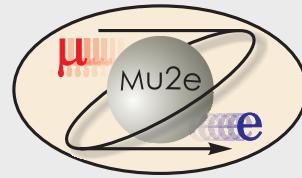


L. Calibbi, A. Faccia, A. Masiero, S. Vempati, hep-ph/0605139

neutrino mass via the see-saw mechanism, analysis is performed in an SO(10) framework



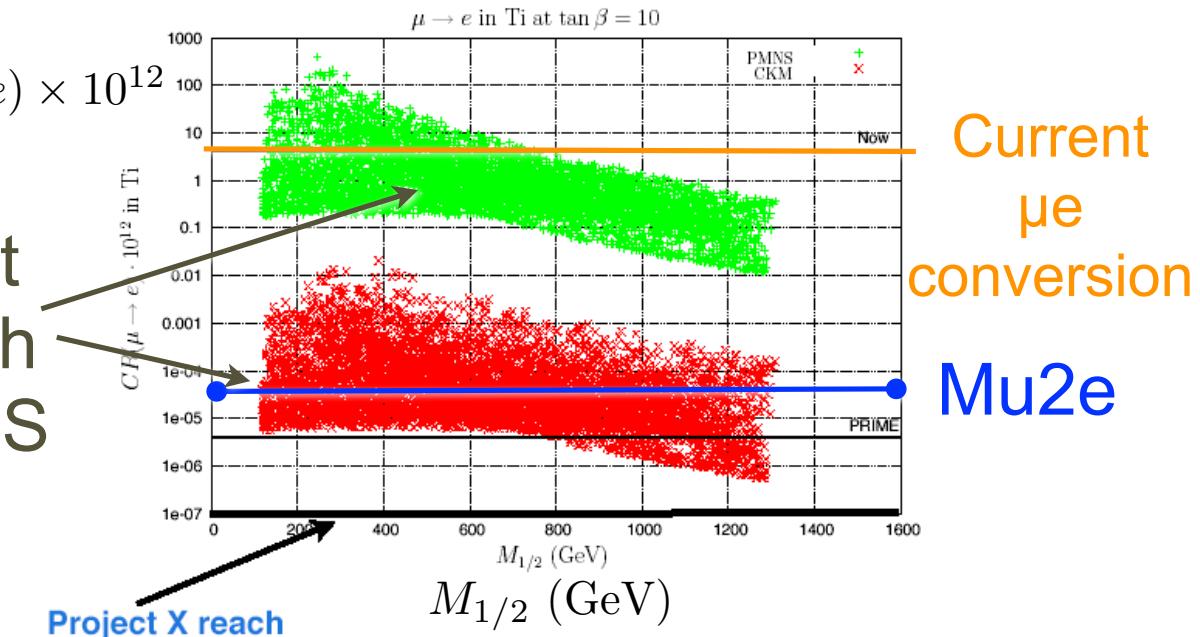
# And Muon-Electron Conversion



$$\tan \beta = 10$$

Neutrino-Matrix Like (PMNS)   Minimal Flavor Violation(CKM)

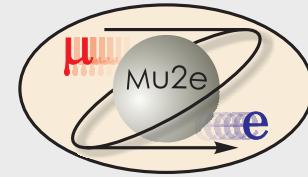
measurement  
can distinguish  
between PMNS  
and MFV



L. Calibbi, A. Faccia, A. Masiero, S. Vempati, hep-ph/0605139

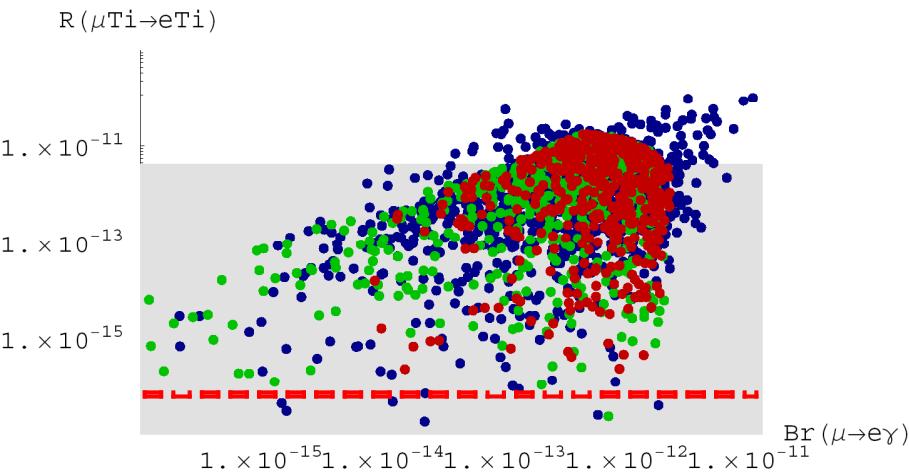
*complementarity between Lepton Flavor Violation  
(LFV) and LHC experiments*

# Combining $\mu \rightarrow e\gamma$ with $\mu \rightarrow e$ Conversion



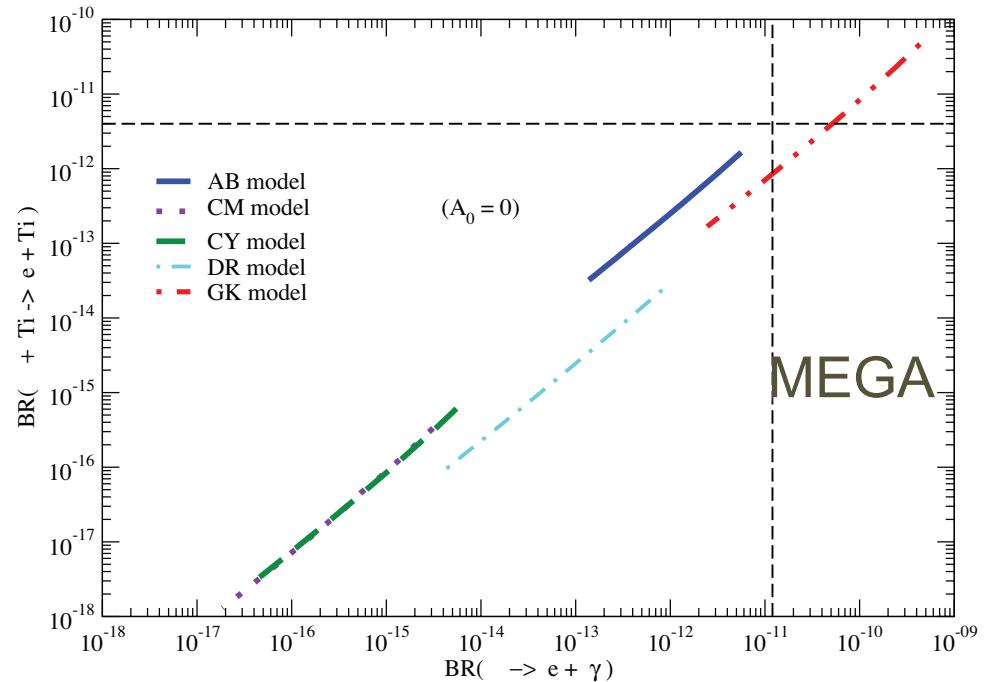
Randall-Sundrum

M. Blanke, A. J. Buras, B. Duling, A. Poschenrieder and C. Tarantino, JHEP 0705, 013 (2007).

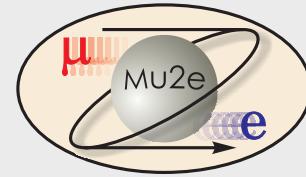


SO(10) models:

C. Albright and M. Chen, arXiv:0802.4228, PRD D77:113010, 2008.

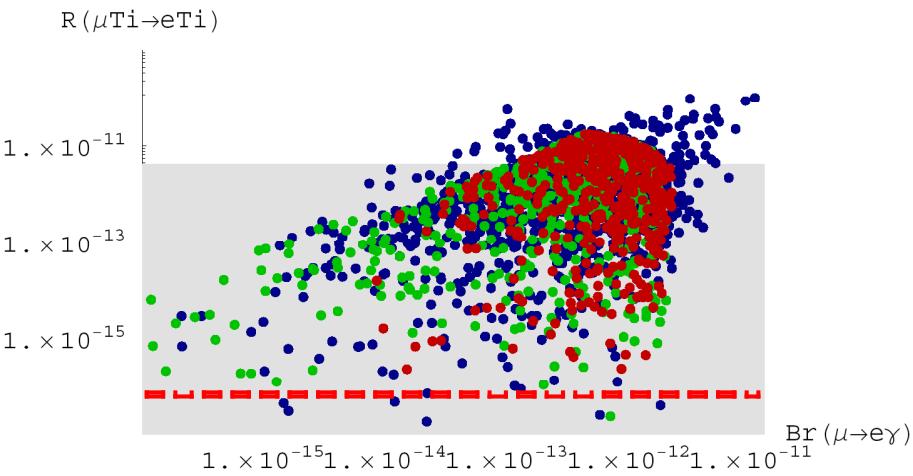


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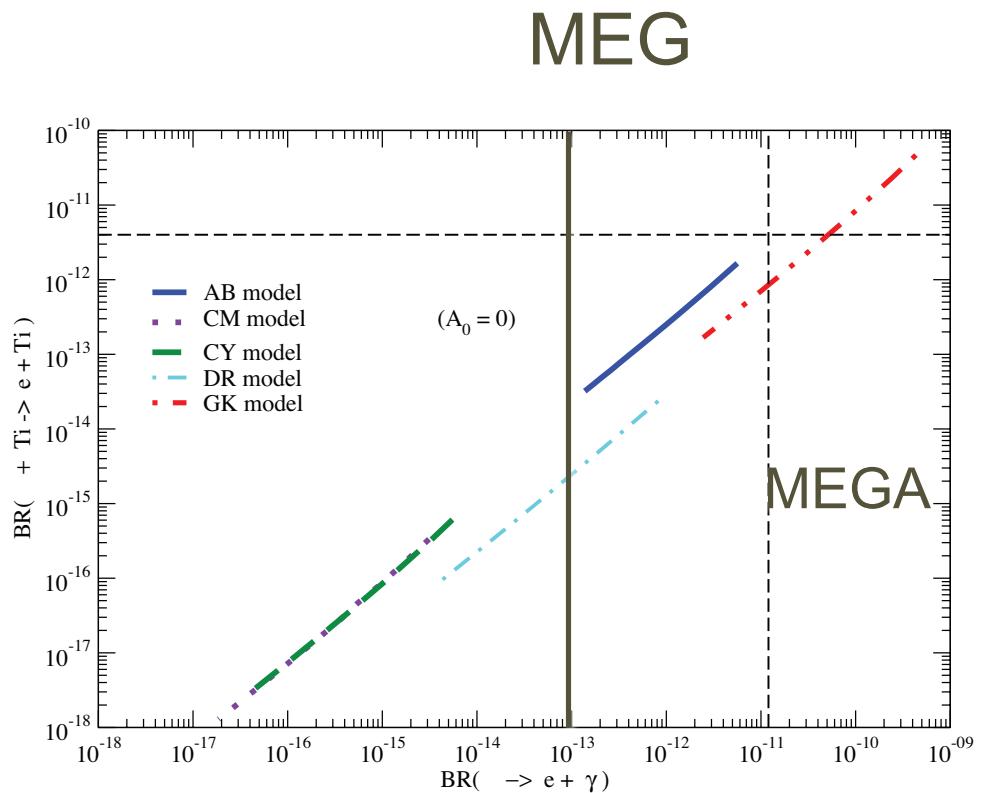
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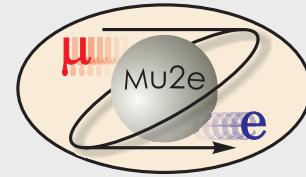


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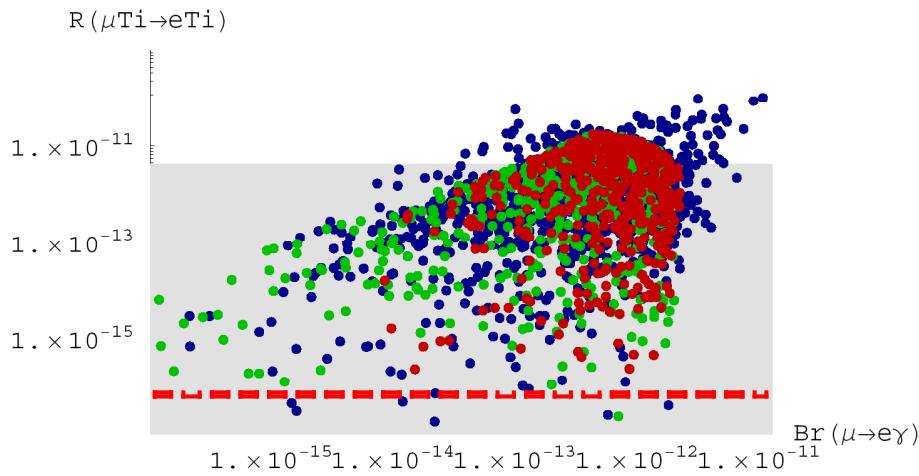


# Combining $\mu \rightarrow e\gamma$ with $\mu \rightarrow e$ Conversion



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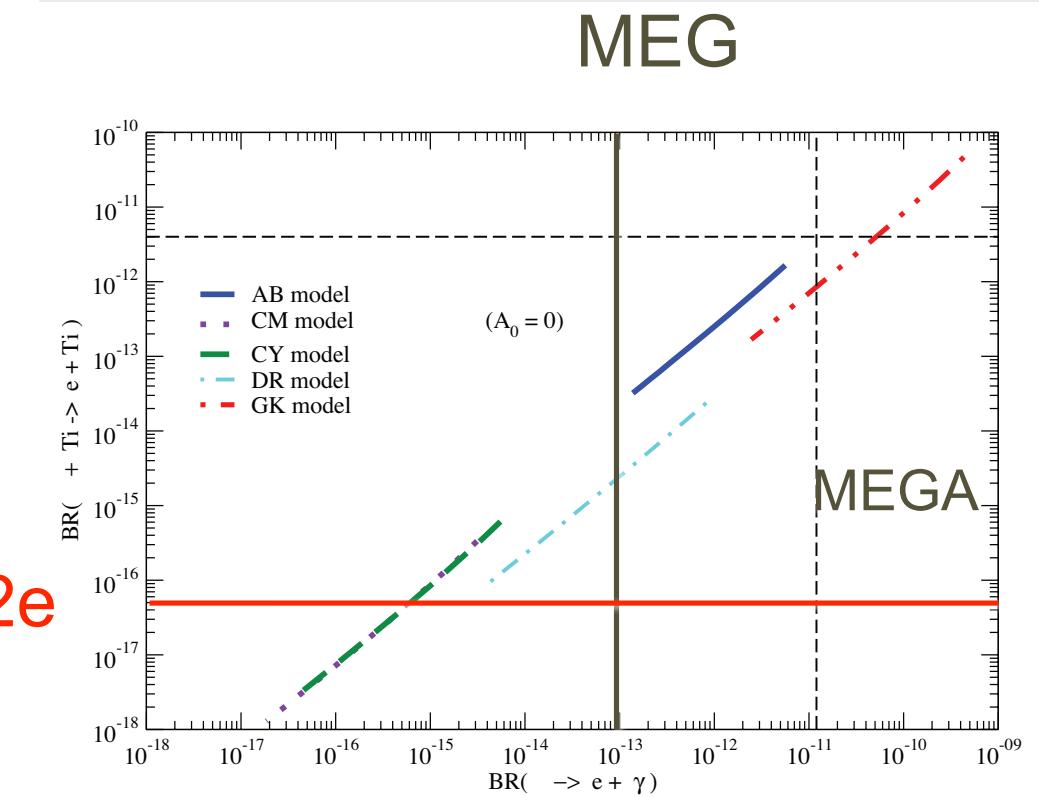
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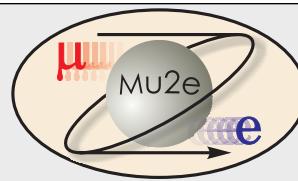


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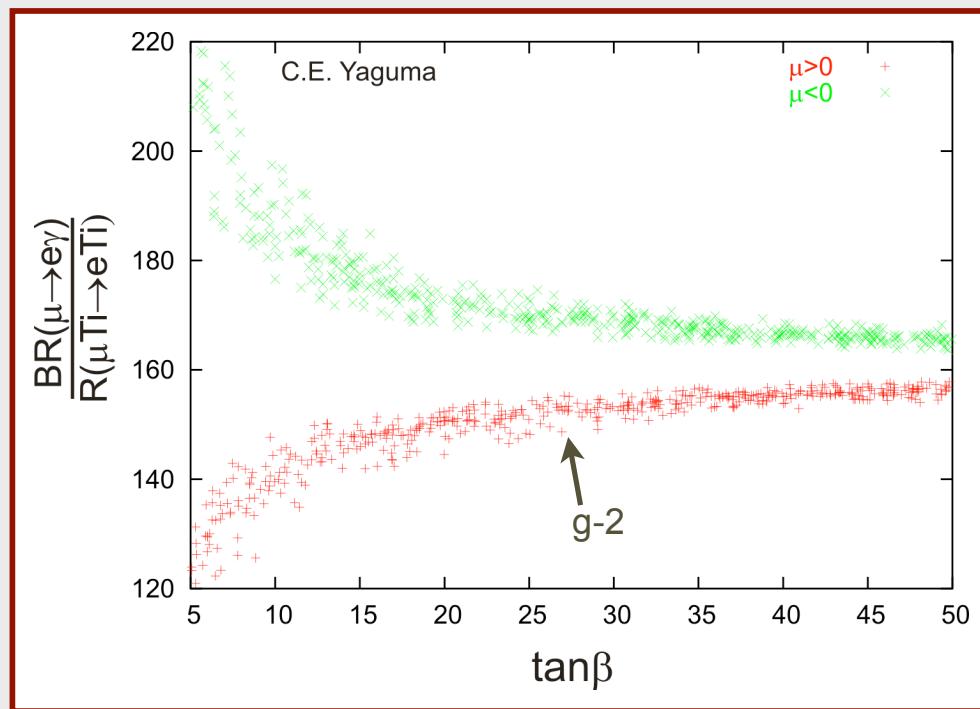
Mu2e





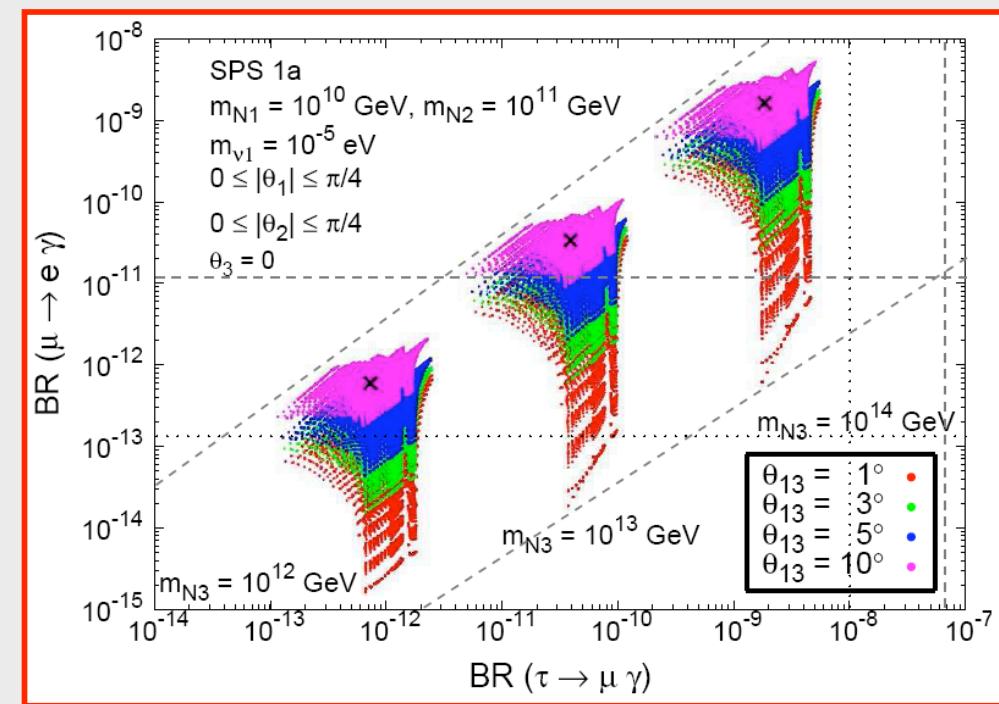
# Pinning Down SuperSymmetry

MSSM w mSUGRA

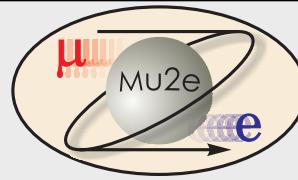


Yaguna, hep-ph/0502014v2

CMSSM - seesaw



Antusch et al., hep-ph/0610439

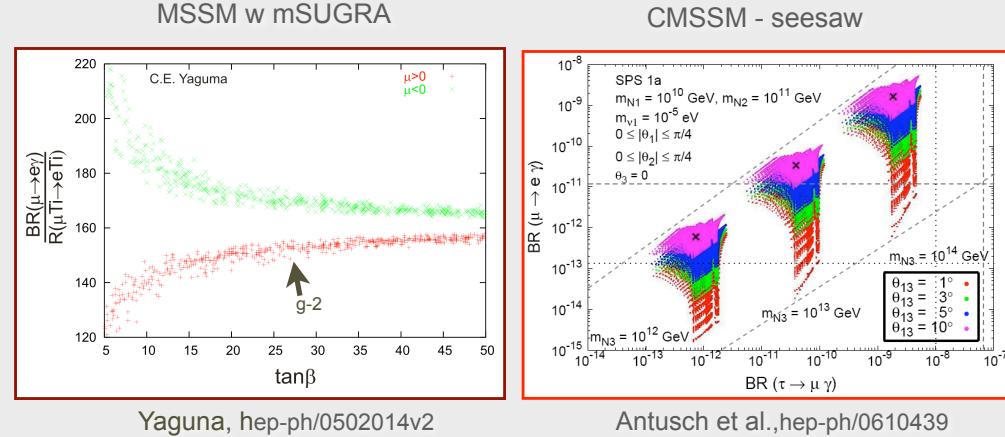


# Pinning Down SuperSymmetry

- Need:

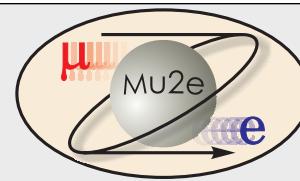
- observation of CLFV in more than one channel, and/or
- evidence from LHC, g-2, or elsewhere

to allow discrimination among different models





# In SUSY



- Large effects in many models
  - only  $\mu \rightarrow e\gamma$  and  $\mu e$  conversion

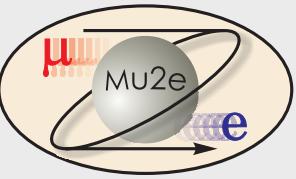


	AC	RVV2	AKM	$\delta LL$	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★	★	★	★	★	★★★	?
$\epsilon_K$	★	★★★	★★★	★	★	★★	★★★
$S_{\psi\phi}$	★★★	★★★	★★★	★	★	★★★	★★★
$S_{\phi K_S}$	★★★	★★	★	★★★	★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★	★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★	★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★	★★★	★★★	★★★	★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$\mu \rightarrow e \gamma$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$\tau \rightarrow \mu \gamma$	★★★	★★★	★	★★★	★★★	★★★	★★★
$\mu + N \rightarrow e + N$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$d_n$	★★★	★★★	★★★	★★	★★★	★	★★★
$d_e$	★★★	★★★	★★	★	★★★	★	★★★
$(g-2)_\mu$	★★★	★★★	★★	★★★	★★★	★	?

Table 8: “DNA” of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models. ★★★ signals large effects, ★★ visible but small effects and ★ implies that the given model does not predict sizable effects in that observable.

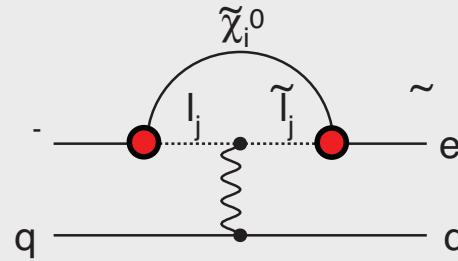


# Contributions to $\mu e$ Conversion



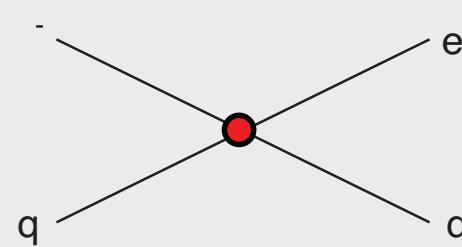
## Supersymmetry

rate  $\sim 10^{-15}$



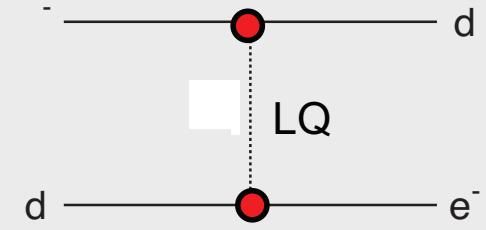
## Compositeness

$\Lambda_c \sim 3000$  TeV



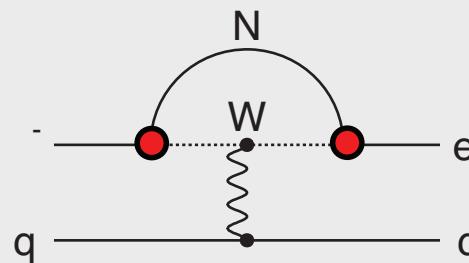
## Leptoquark

$$M_{LQ} = 3000 (\lambda_{\mu d} \lambda_{ed})^{1/2} \text{ TeV}/c^2$$



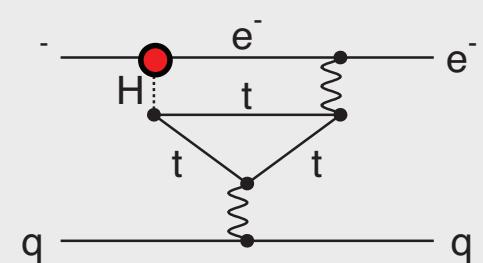
## Heavy Neutrinos

$$|U_{\mu N} U_{e N}|^2 \sim 8 \times 10^{-13}$$



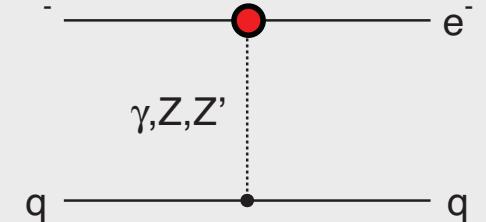
## Second Higgs Doublet

$$g(H_{\mu e}) \sim 10^{-4} g(H_{\mu \mu})$$



## Heavy Z' Anomalous Z Coupling

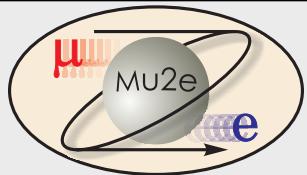
$$M_{Z'} = 3000 \text{ TeV}/c^2$$



also see Flavour physics of leptons and dipole moments, [arXiv:0801.1826](https://arxiv.org/abs/0801.1826) ;  
 Marciano, Mori, and Roney, Ann. Rev. Nucl. Sci. 58, doi:10.1146/annurev.nucl.58.110707.171126 ;



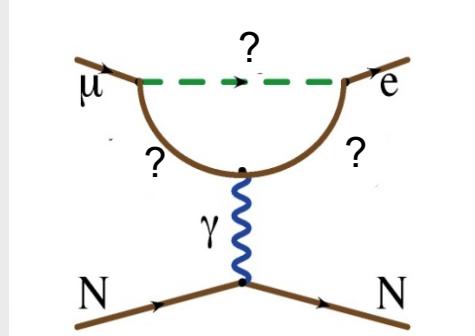
# “Model-Independent” Form



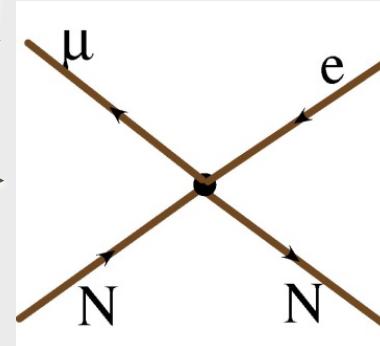
$$\mathcal{L}_{\text{CLFV}} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{u}_L \gamma_\mu u_L + \bar{d}_L \gamma_\mu d_L)$$

“Loops”

$\kappa=0$



mass scale  $\Lambda$   
 $\kappa$



$\kappa=1$

Supersymmetry and Heavy  
Neutrinos

Contributes to  $\mu \rightarrow e\gamma$

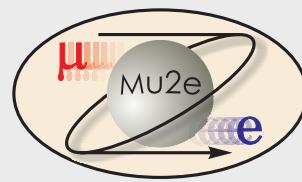
“Contact Terms”

Does not produce  $\mu \rightarrow e\gamma$

Quantitative Comparison?



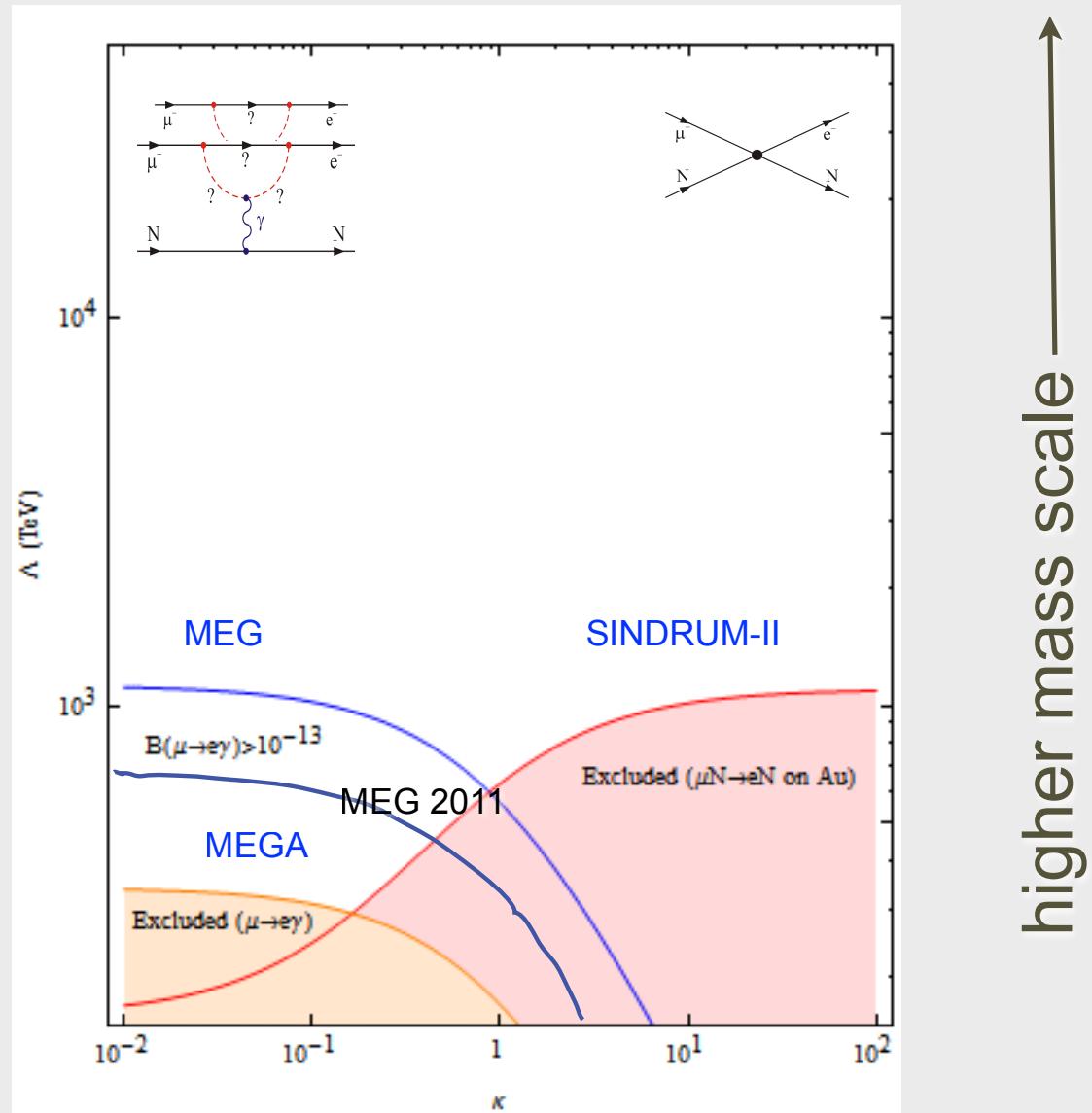
# $\mu e$ Conversion and $\mu \rightarrow e\gamma$



$\Lambda$  (TeV)

1) Mass Reach  
to  $\sim 10^4$  TeV

2) about x2  
beyond MEG in  
loop-dominated  
physics

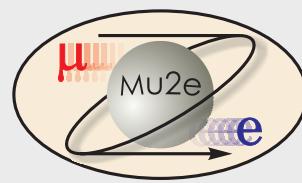


$\kappa$

André de Gouvêa, Project  
X Workshop Golden Book



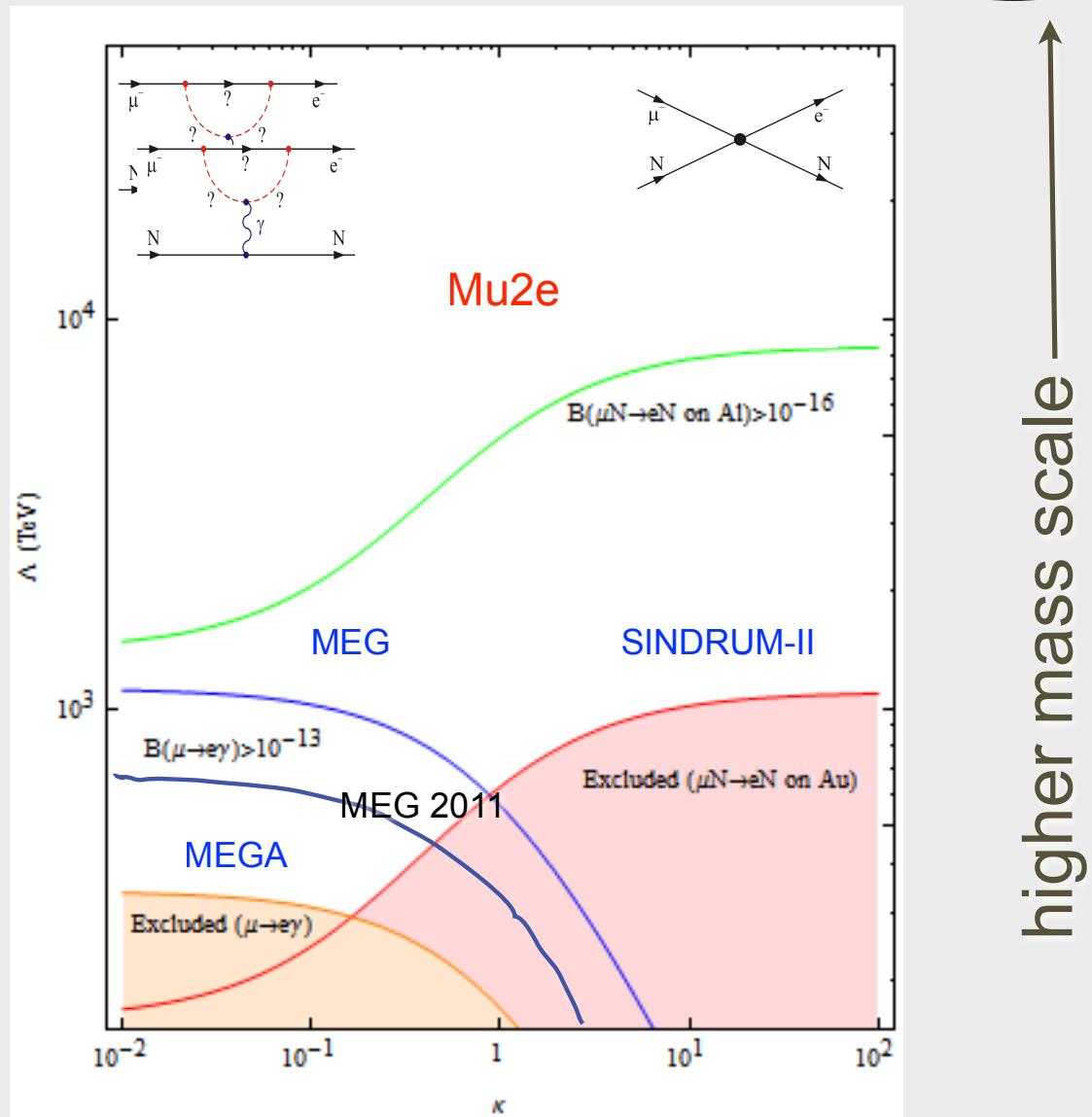
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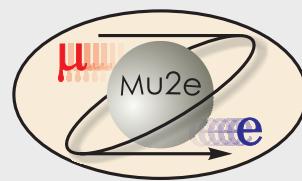


$\kappa$

André de Gouvêa, Project X Workshop Golden Book



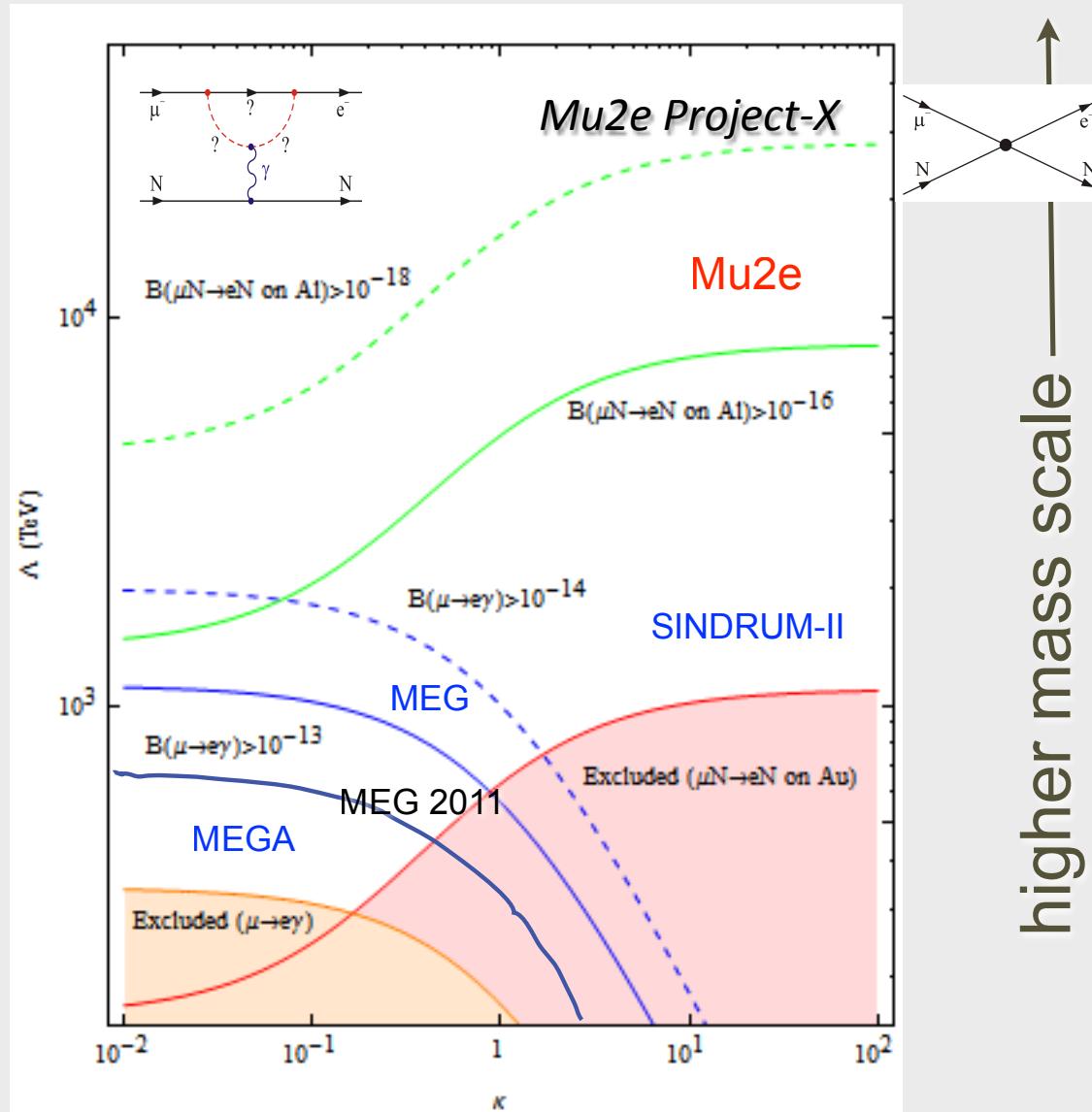
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physics

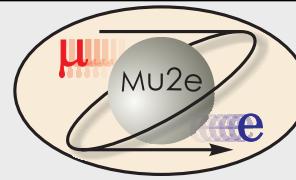


$\kappa$

André de Gouvêa, Project X Workshop Golden Book



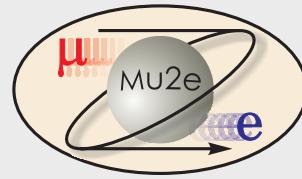
# Outline



- The search for muon-electron conversion
- *Experimental Technique*
- Fermilab Accelerator
- Project X Upgrades and Mu2e
- Cost and Schedule
- Conclusions



# Overview Of Processes

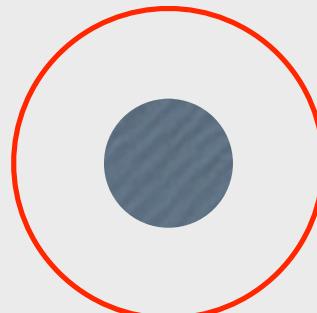


$\mu^-$  stops in thin Al foil



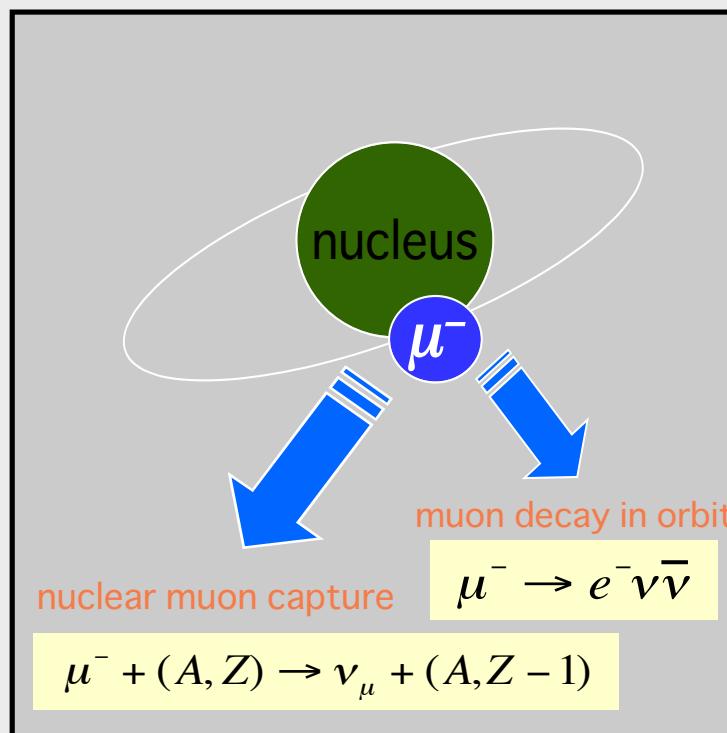
*the Bohr radius is  $\sim 20 \text{ fm}$ ,  
so the  $\mu^-$  sees the nucleus*

$\mu^-$  in 1s state



Al Nucleus  
 $\sim 4 \text{ fm}$

muon capture,  
muon “falls into”  
nucleus:  
***normalization***

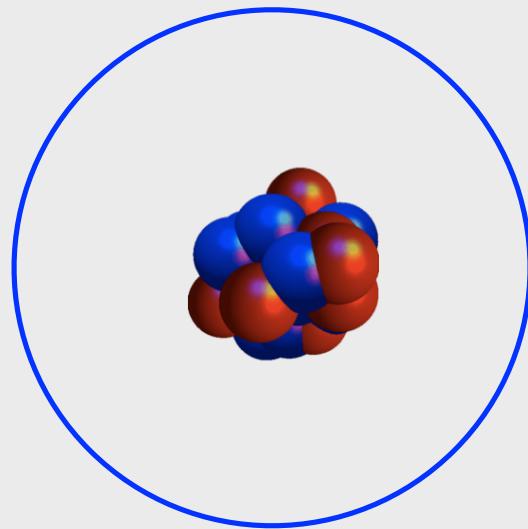
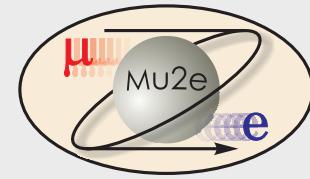


60% capture  
40% decay

Decay in Orbit:  
***background***

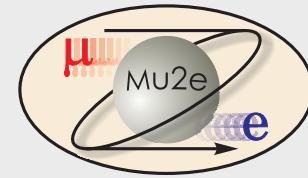


# Three Possibilities: Normalization

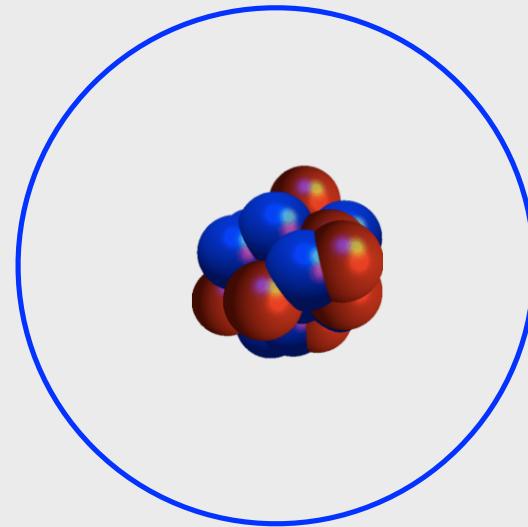




# Three Possibilities: Normalization

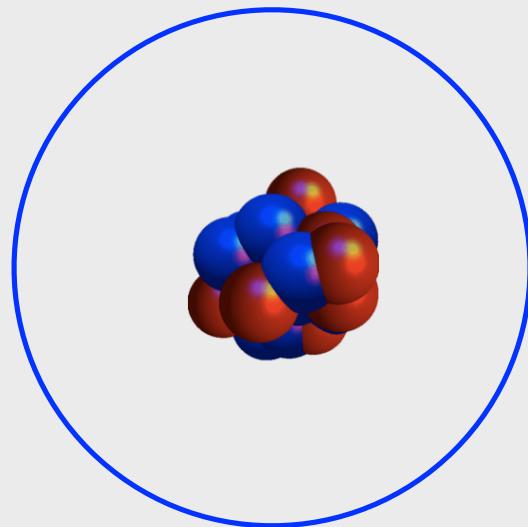
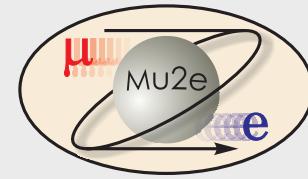


muon stops



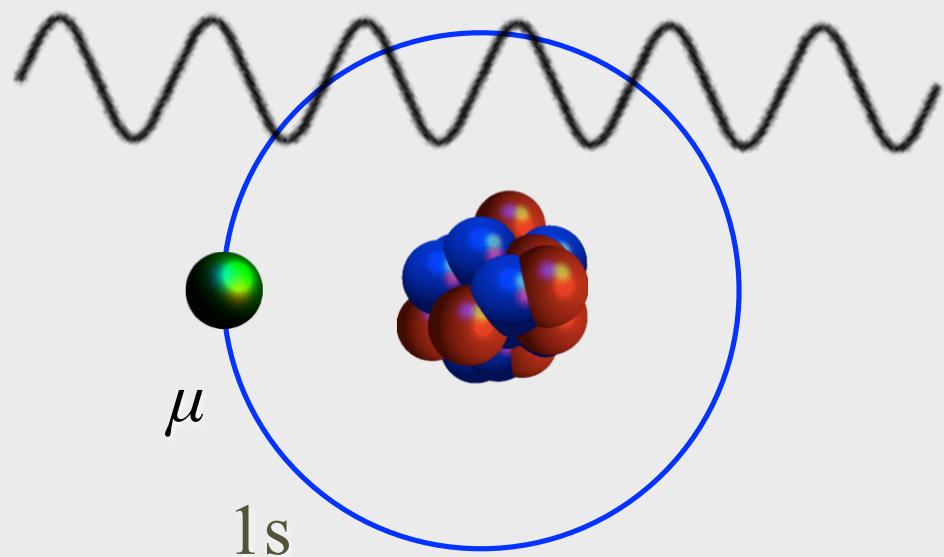
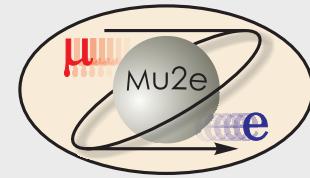


# Three Possibilities: Normalization



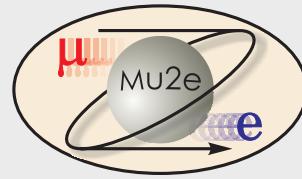


# Three Possibilities: Normalization





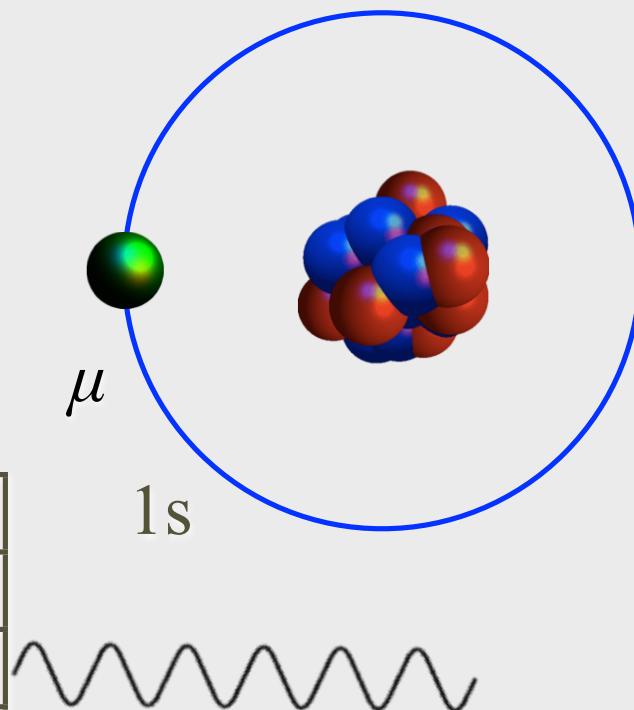
# Three Possibilities: Normalization



X-Rays from  
cascade  
(occurs in  
<psec)

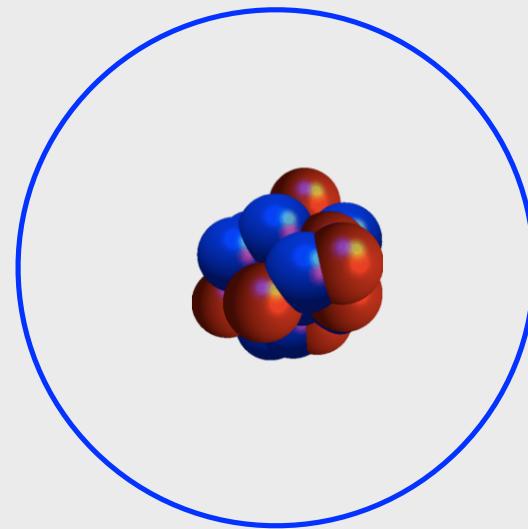
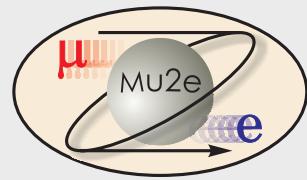
detect these  
for  
normalization

Transition	Energy
$3d \rightarrow 2p$	66 keV
$2p \rightarrow 1s$	356 keV
$3d \rightarrow 1s$	423 keV
$4p \rightarrow 1s$	446 keV



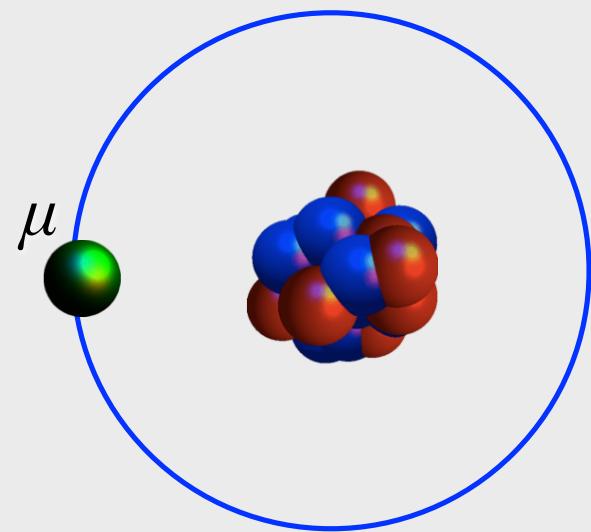
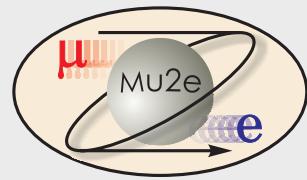


# Normalization to Nuclear Capture



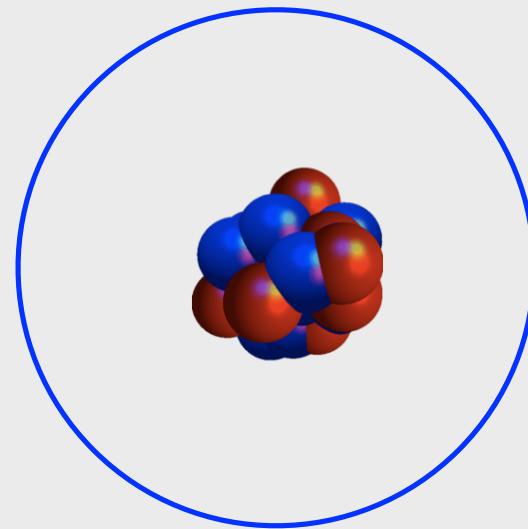
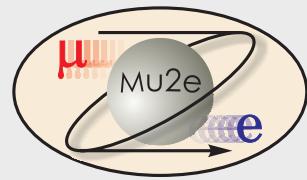


# Normalization to Nuclear Capture



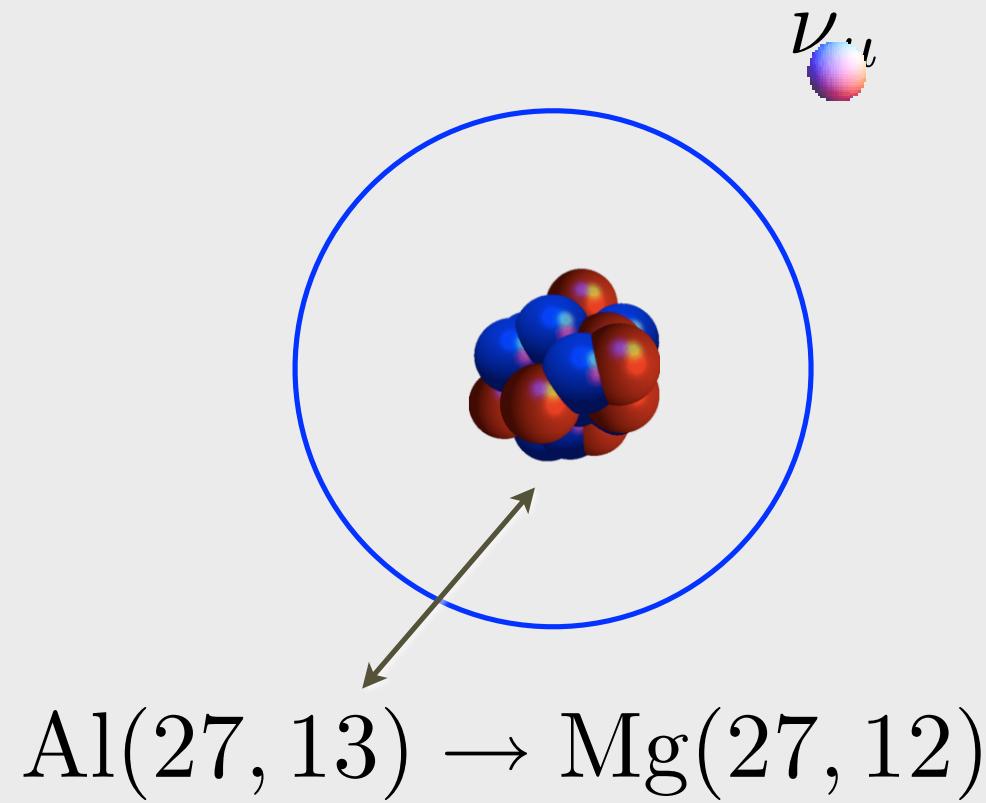
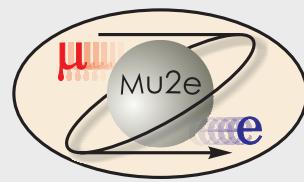


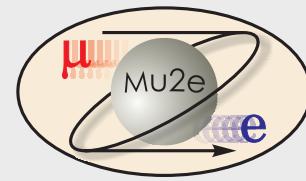
# Normalization to Nuclear Capture





# Normalization to Nuclear Capture

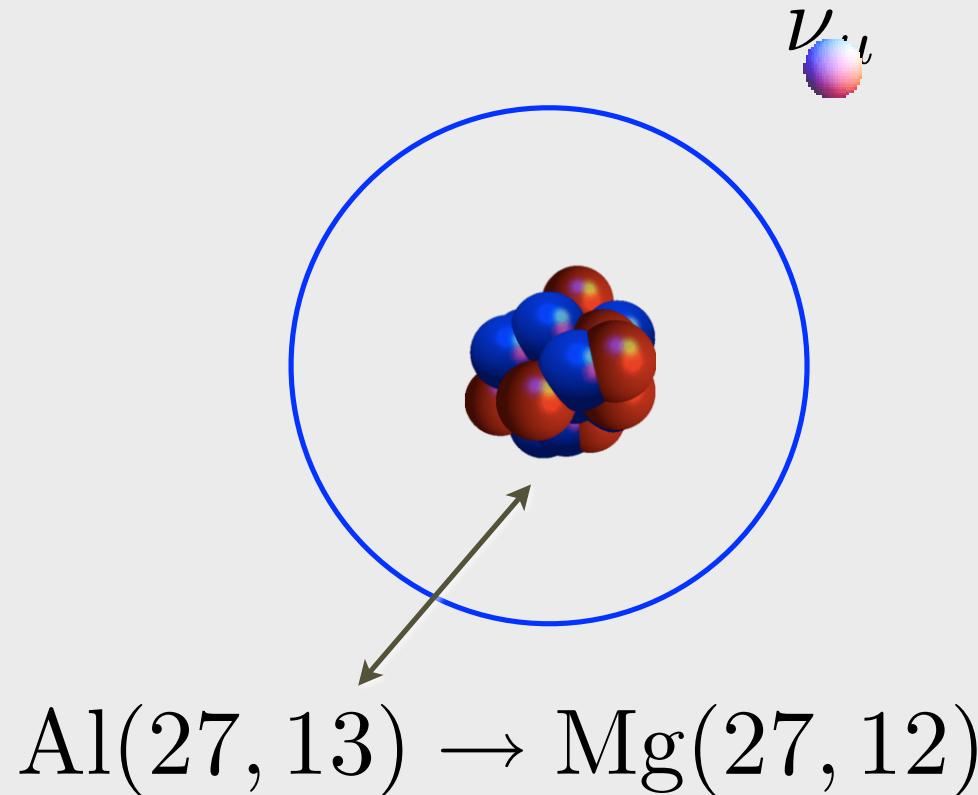




# Normalization to Nuclear Capture

1) measure stop rate 2) calculate capture rate/stop

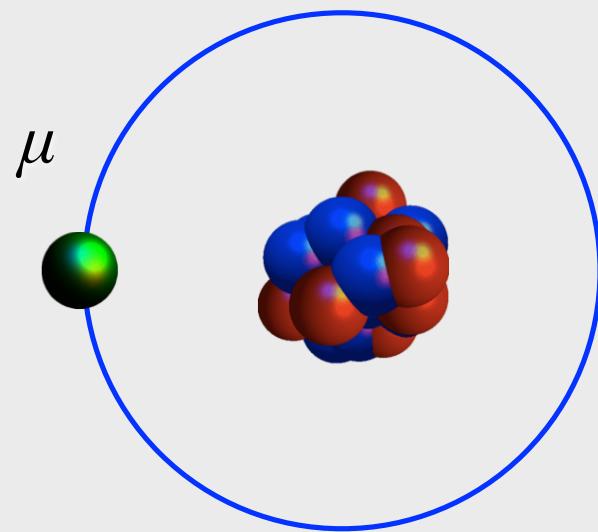
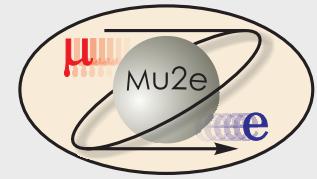
Kitano et al. ,Phys.Rev.D66:096002,2002, Erratum-ibid.D76:059902,2007. e-Print: hep-ph/0203110



$$\text{then compute } R_{\mu e} = \frac{\mu N \rightarrow eN}{\mu \text{ Al}(27, 13) \rightarrow \nu_\mu \text{ Mg}(27, 12)}$$

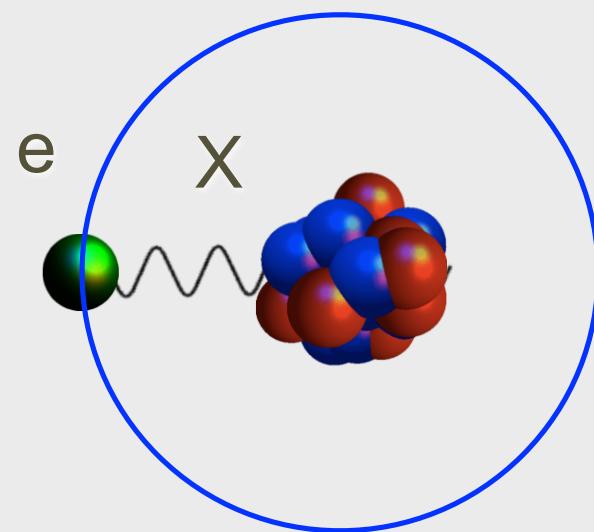
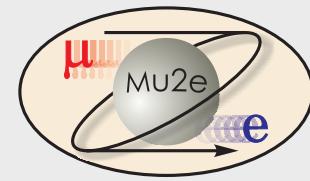


# Three Possibilities: Signal



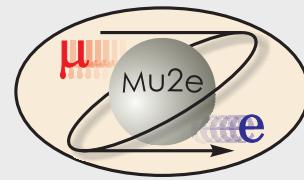


# Three Possibilities: Signal

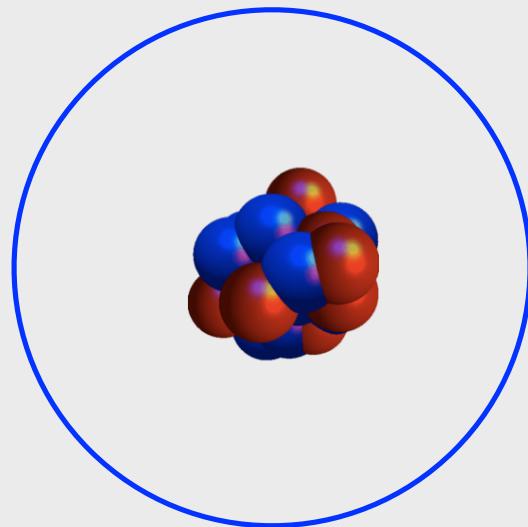




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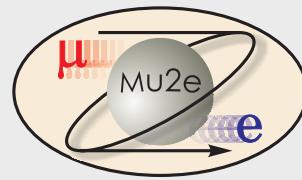


e



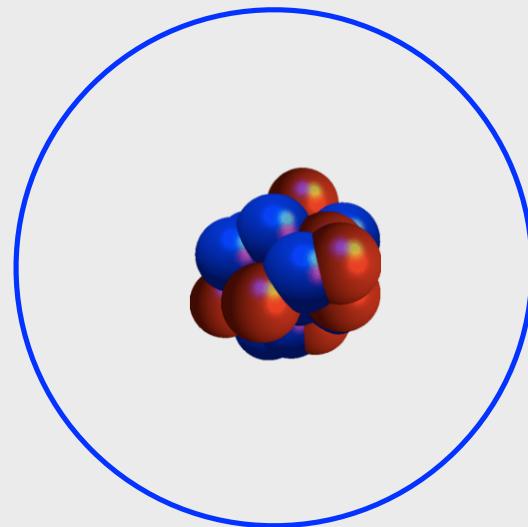


# Three Possibilities: Signal



e

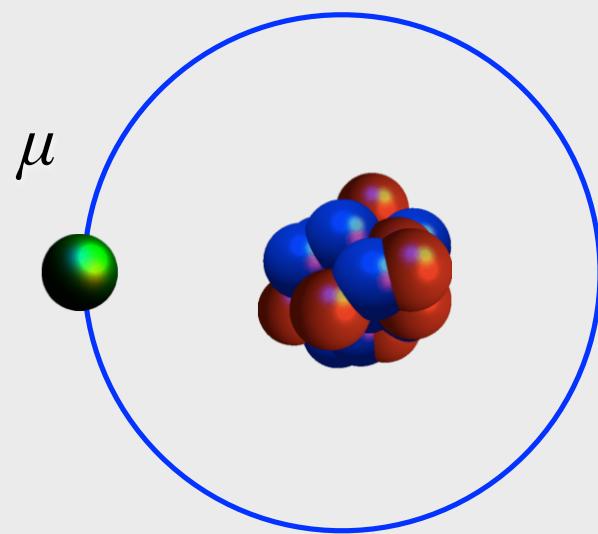
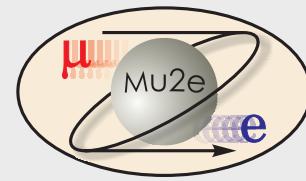
*off to detector!*



coherent recoil of nucleus

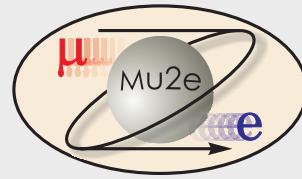


# Three Possibilities: Background

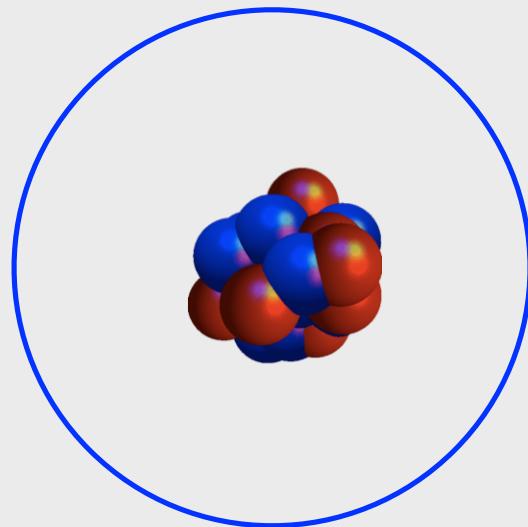




# Three Possibilities: Background

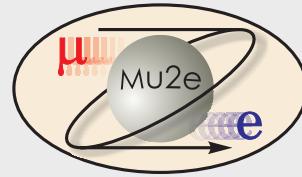


$e$   
 $\nu_\mu$   
 $\bar{\nu}_e$





# Three Possibilities: Background



this electron can be background;  
let's see how



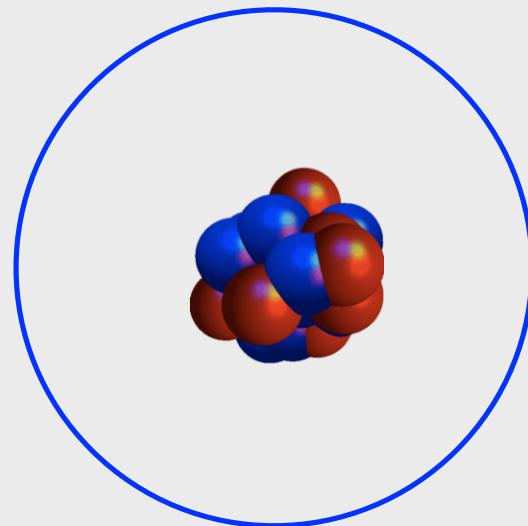
$e$

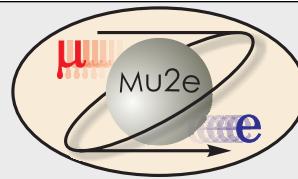


$\nu_\mu$



$\bar{\nu}_e$





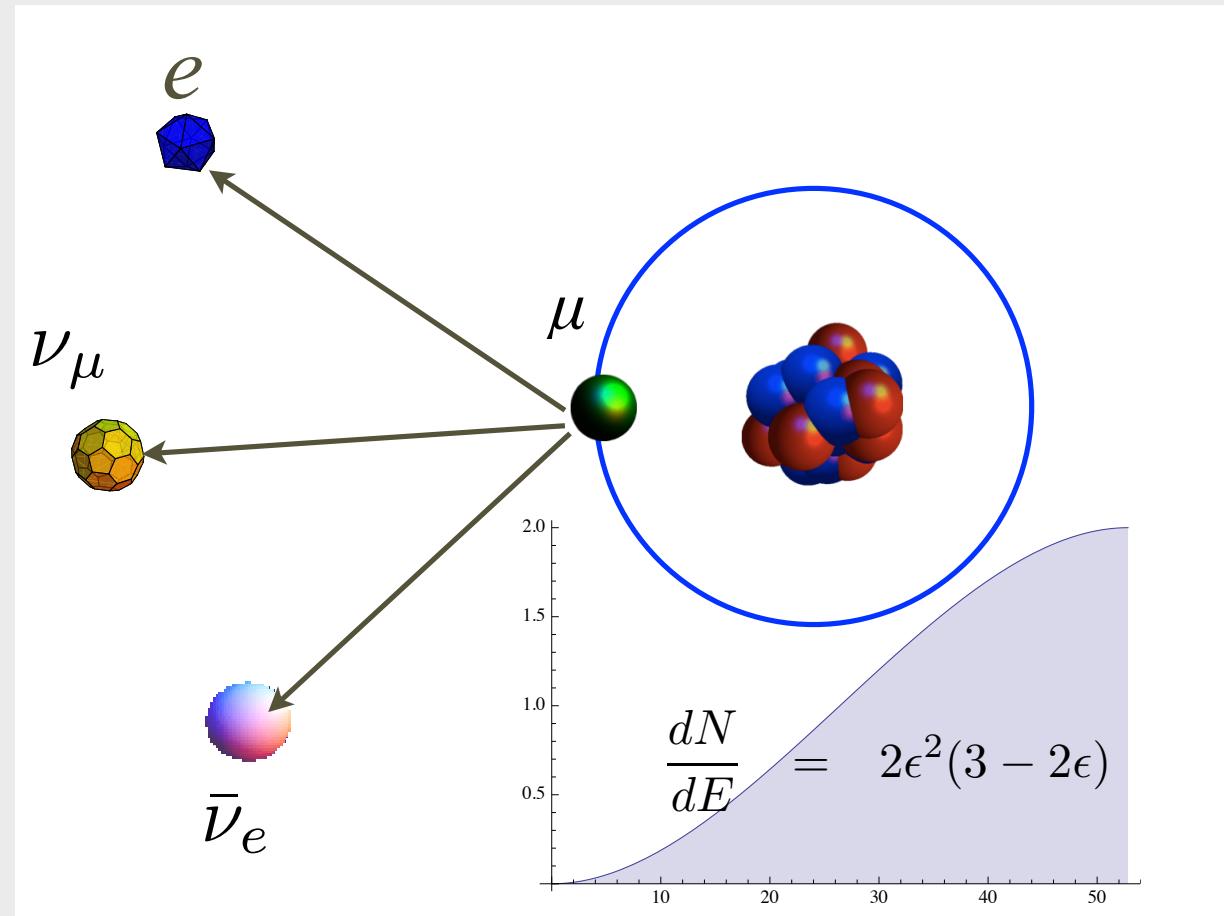
# Decay-In-Orbit: Not always Background

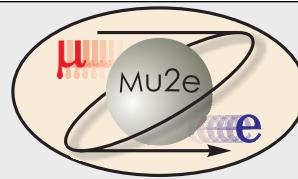
- Peak and Endpoint of Michel Spectrum is at

$$E_{\text{max}} = \frac{m_\mu^2 + m_e^2}{2m_\mu} \approx 52.8 \text{ MeV}$$

- Detector will be insensitive to electrons at this energy

- Recall *signal* at  $105 \text{ MeV} >> 52.8 \text{ MeV}$





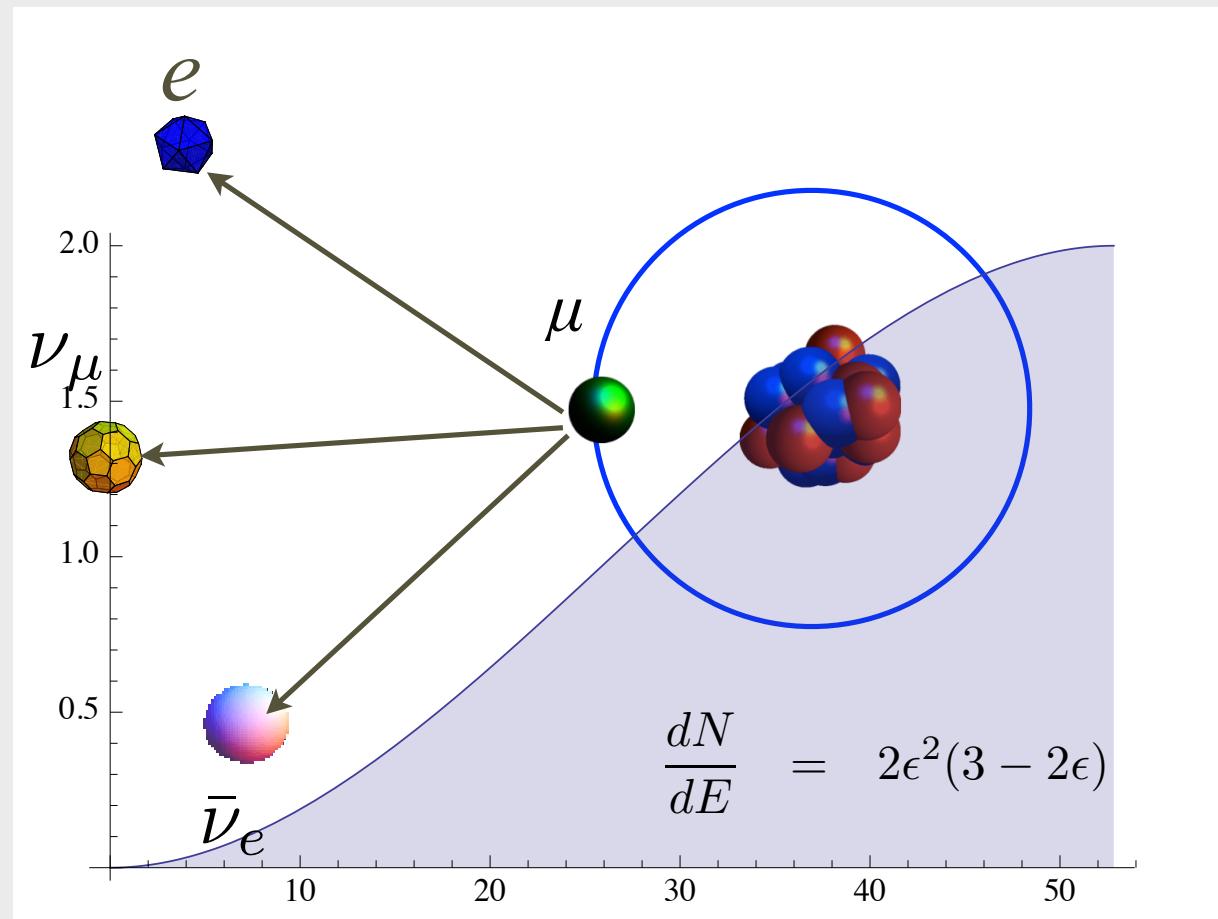
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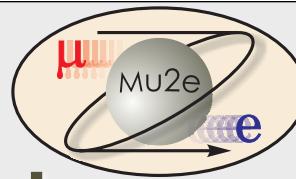
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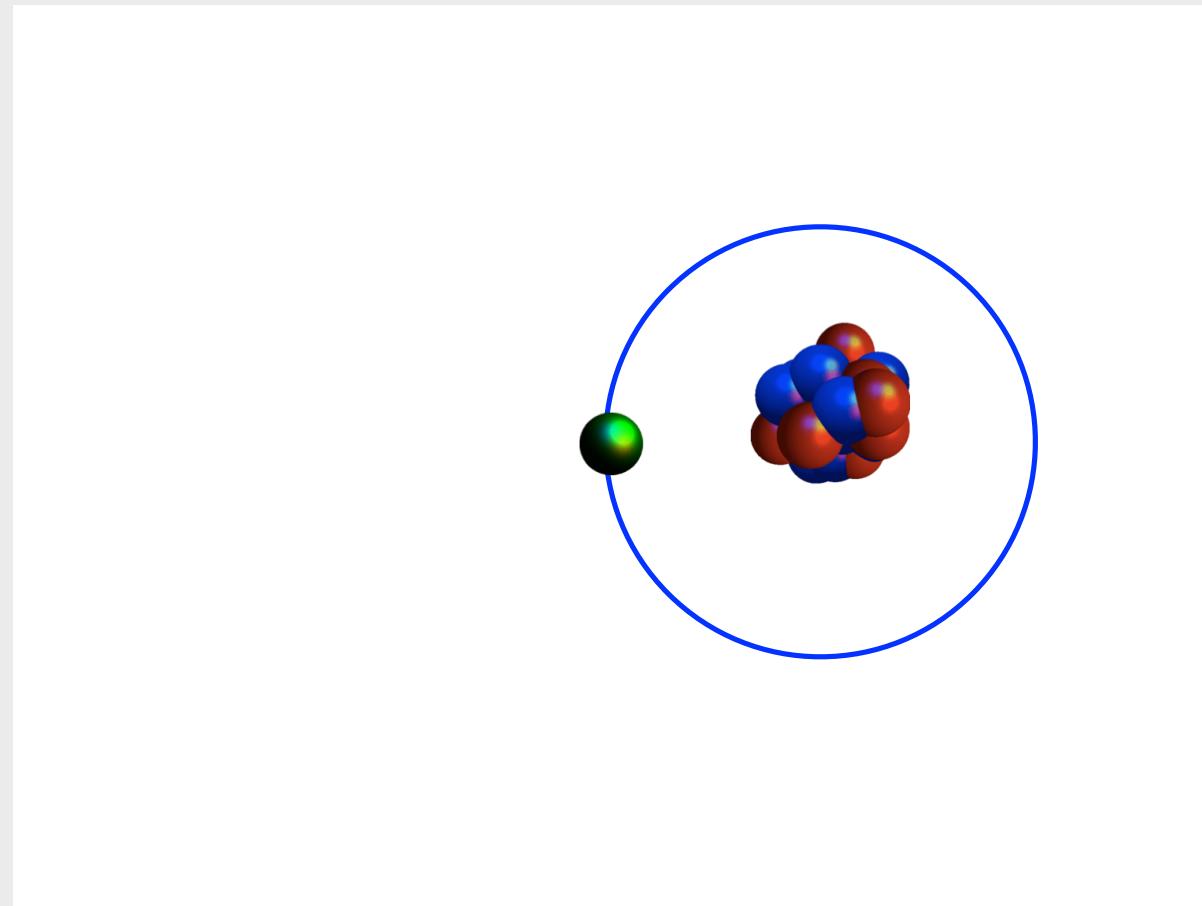
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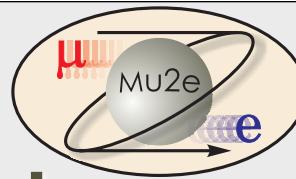




# Decay-In-Orbit Background

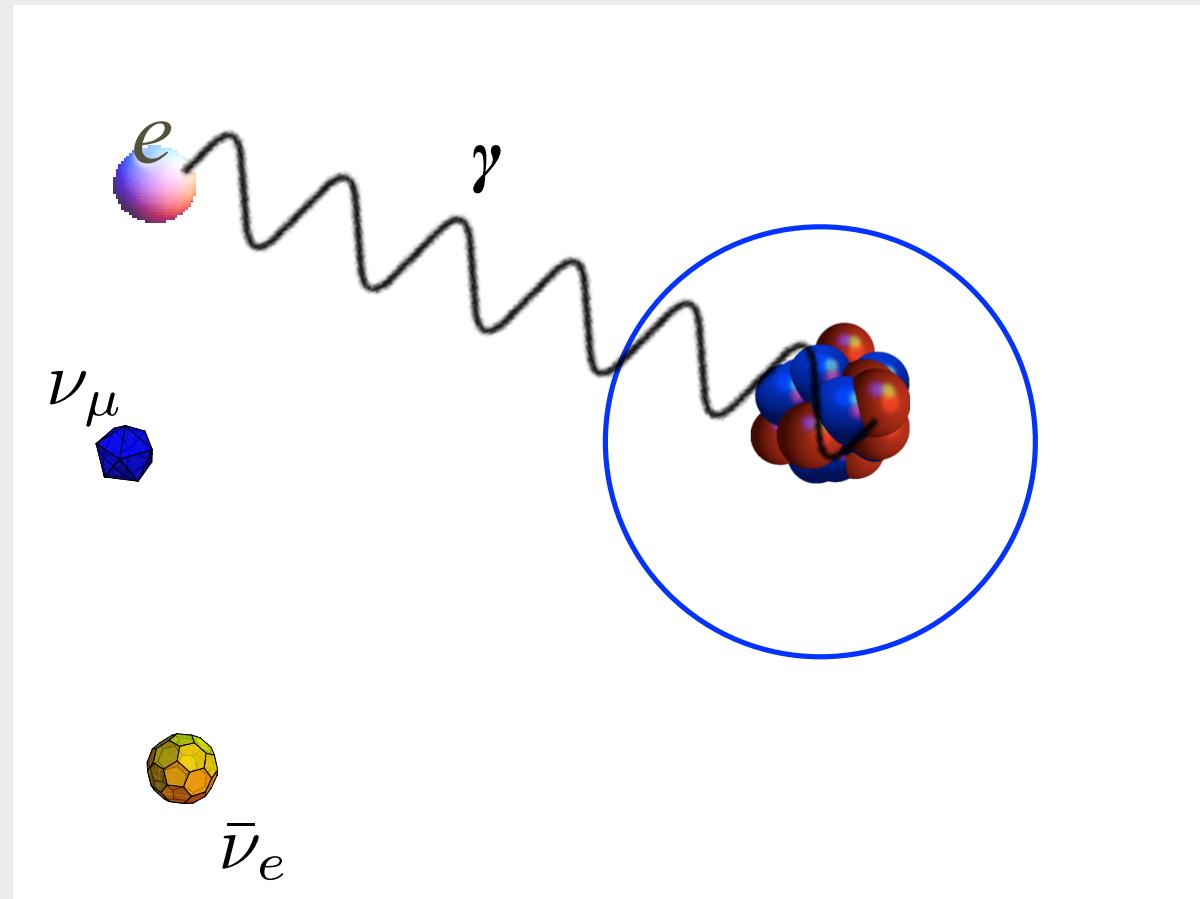
- Same process as before
- But this time, include electron recoil off nucleus
- If neutrinos are at rest, **the DIO electron can be exactly at conversion energy** (up to neutrino mass)

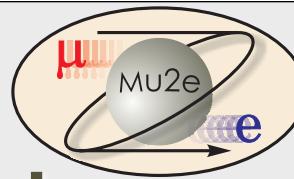




# Decay-In-Orbit Background

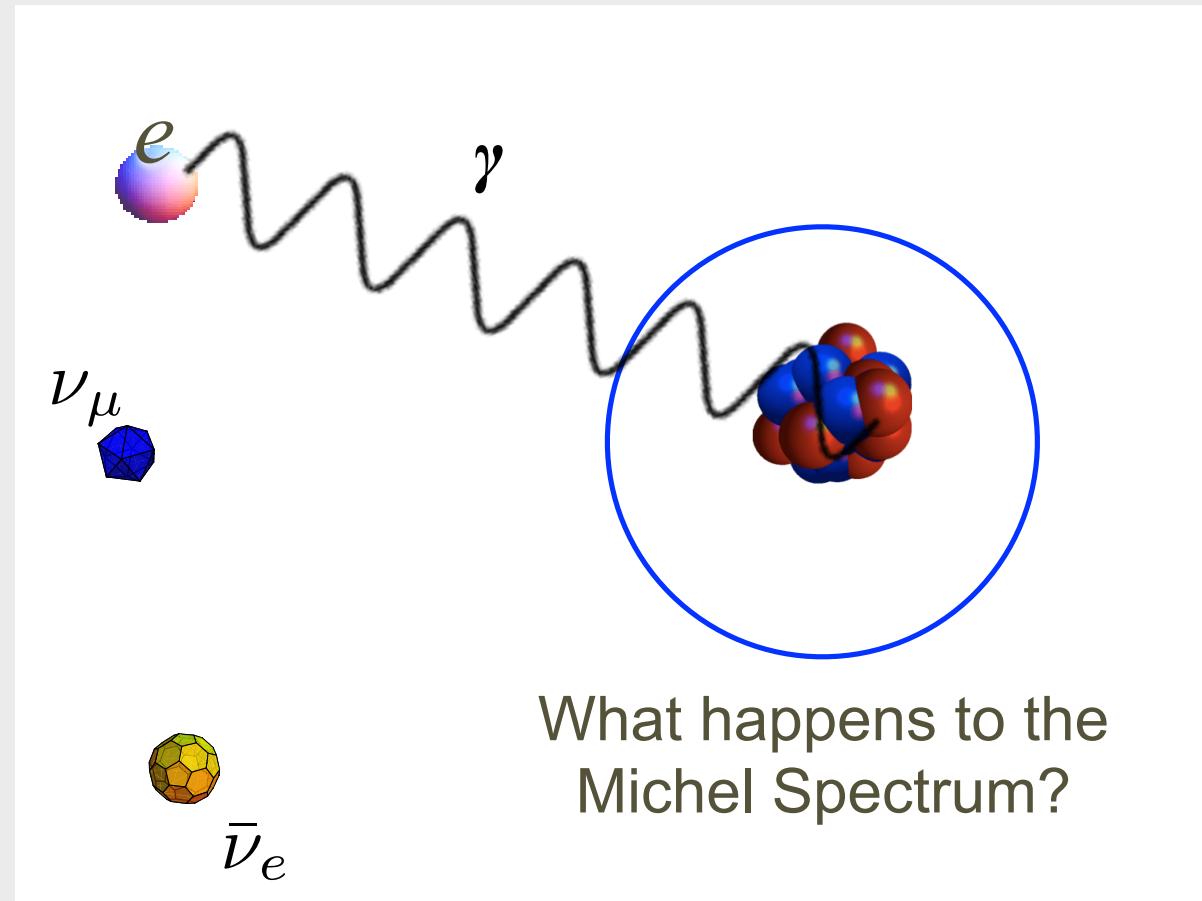
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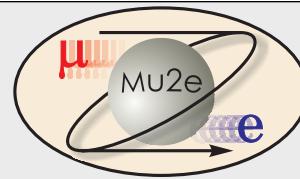
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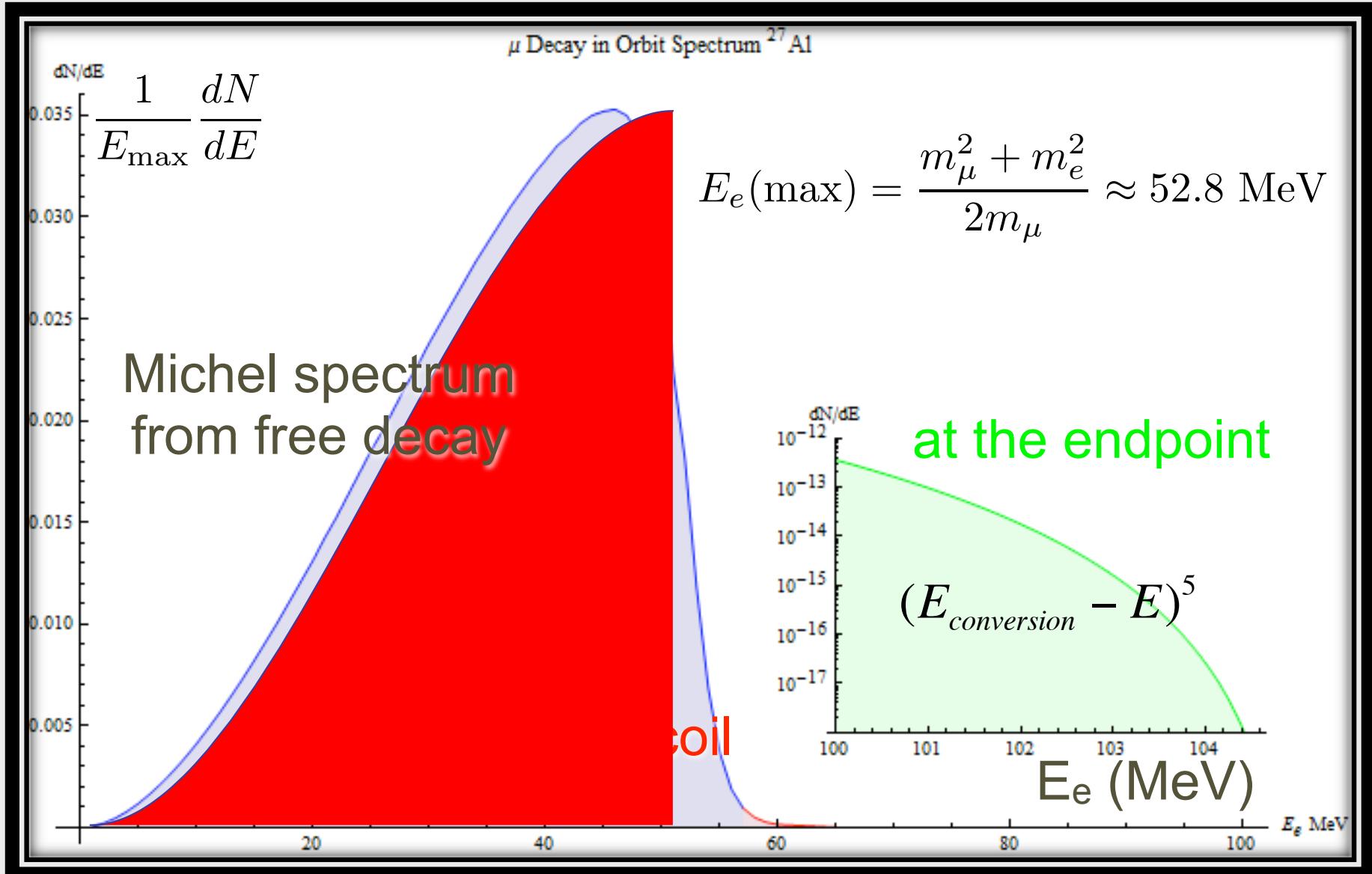




# Decay-in-Orbit Shape

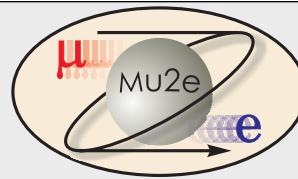


Czarnecki et al., arXiv:1106.4756v2 [hep-ph] Phys. Rev. D 84, 013006 (2011)

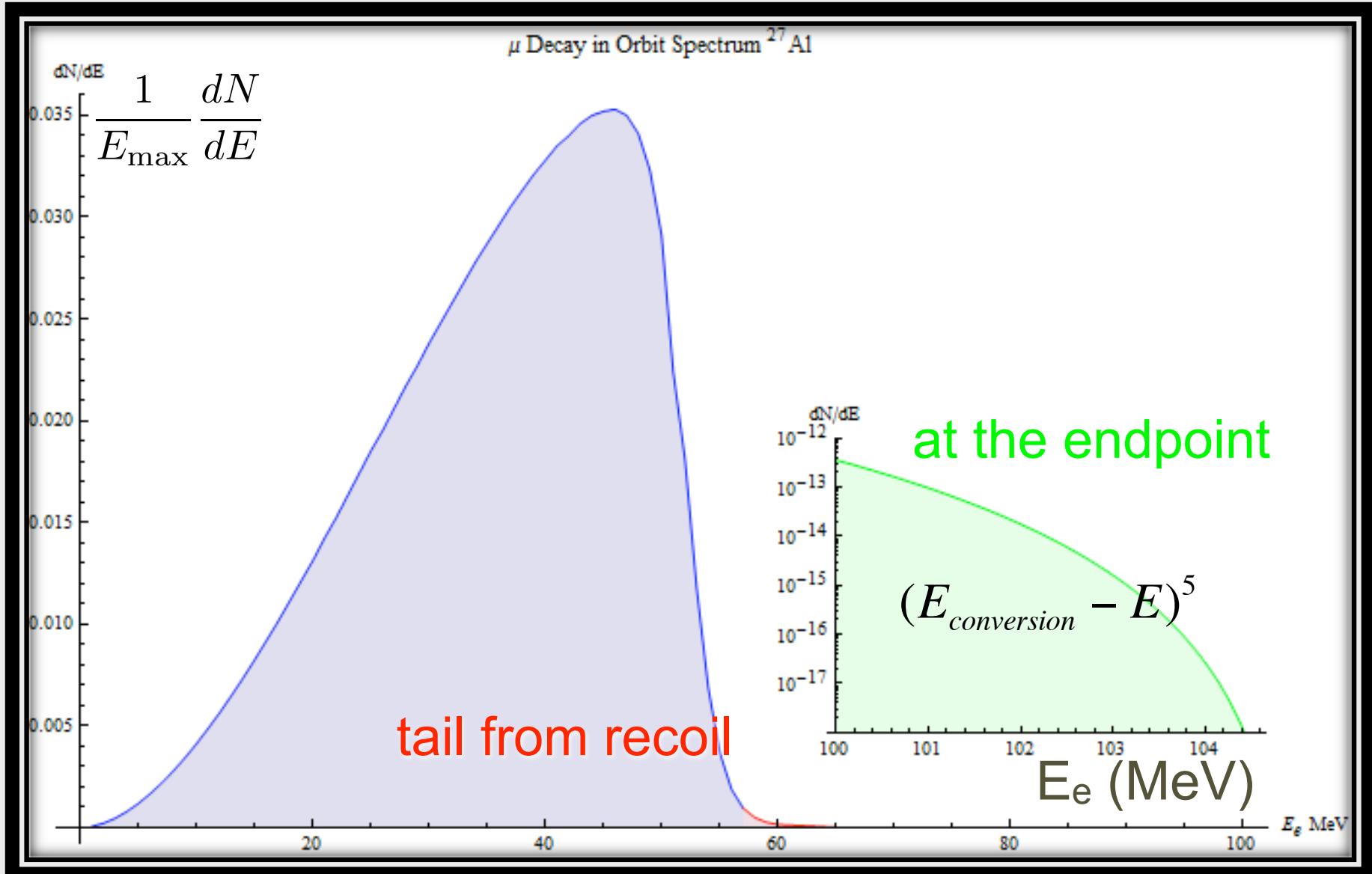




# Decay-in-Orbit Shape

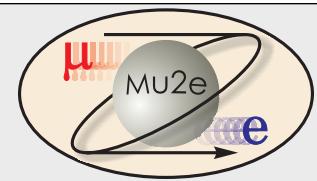


Czarnecki et al., arXiv:1106.4756v2 [hep-ph] Phys. Rev. D 84, 013006 (2011)





# Prompt Backgrounds

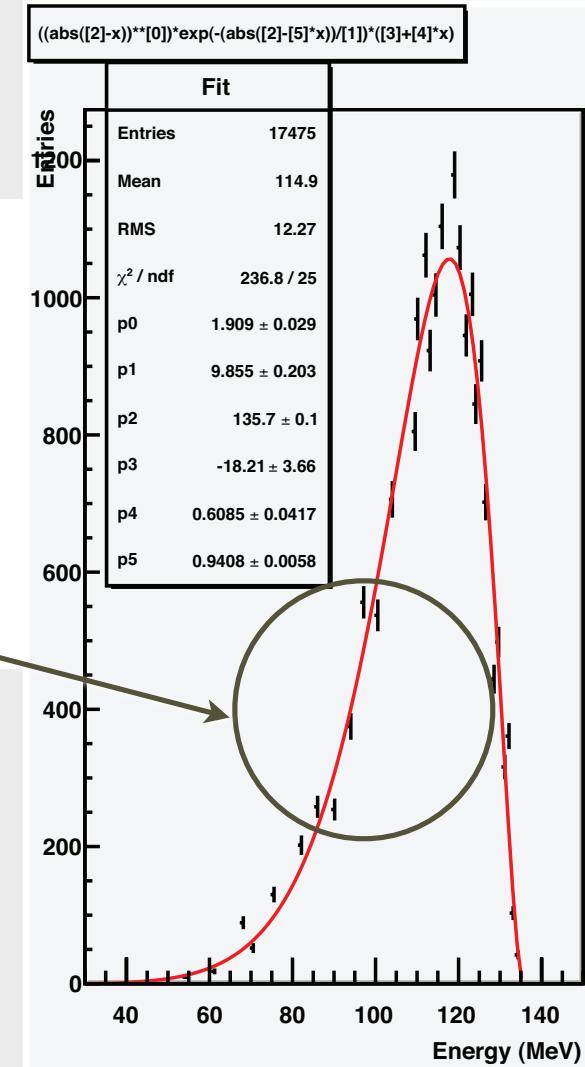


*Particles produced by proton pulse which interact almost immediately when they enter the detector:  $\pi$ , neutrons, pbars*

- **Radiative pion capture**,  $\pi^- + A(N, Z) \rightarrow \gamma + X$ .
  - $\gamma$  up to  $m_\pi$ , peak at 110 MeV;  $\gamma \rightarrow e^+e^-$ ; if one electron  $\sim 100$  MeV in the target, looks like signal: *limitation in best existing experiment, SINDRUM II?*

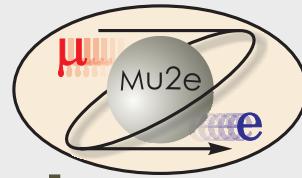
energy spectrum of  $\gamma$  measured on Mg  
J.A. Bistirlich, K.M. Crowe et al., Phys Rev C5, 1867 (1972)

also included internal conversion,  $\pi^- N \rightarrow e^+ e^- X$





# Review:

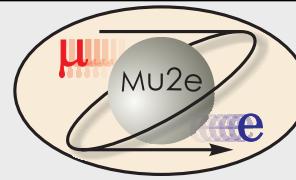


## Two Classes of Backgrounds

	Decay-In-Orbit	Prompt
Source	Intrinsic Physics Background	Radiative $\pi$ Capture: Mostly $\pi$ 's produced in production target
Solution	Spectrometer Design: resolution and pattern recognition	Design of Muon Beam, formation, transport, and time structure

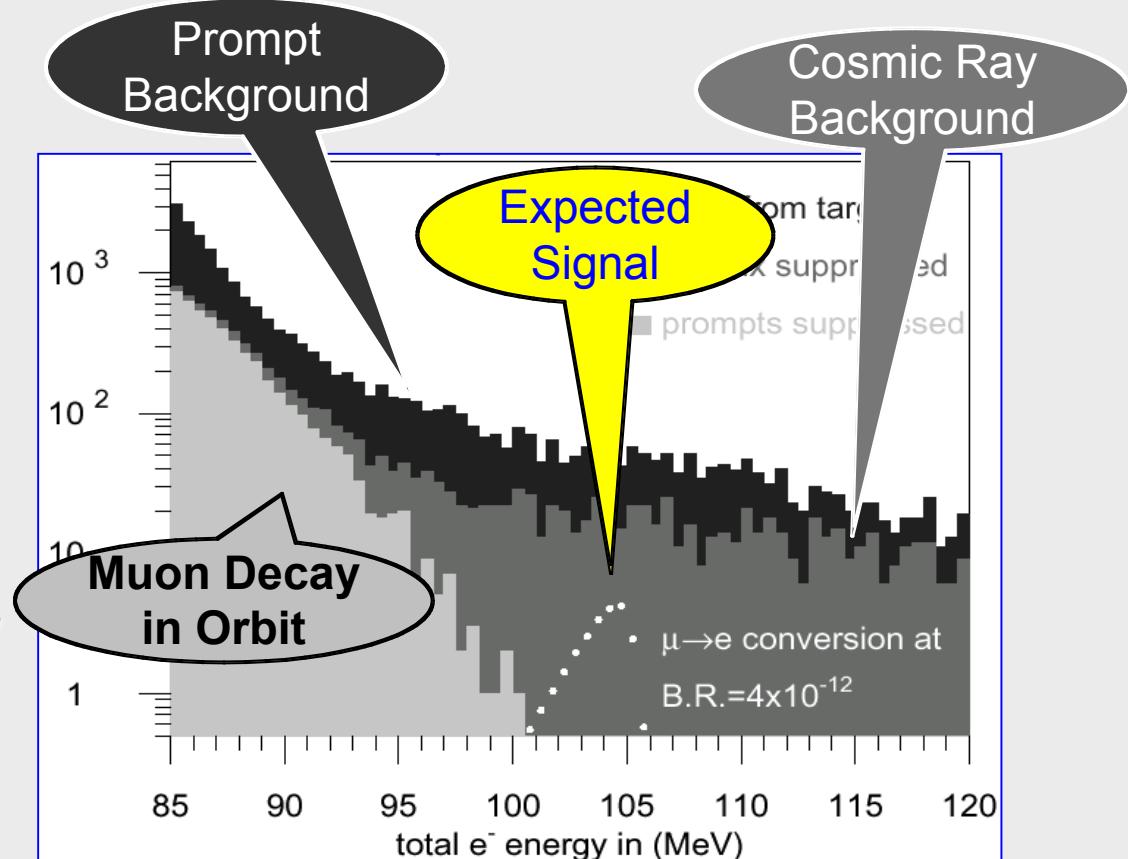


# Previous Best Experiment



## SINDRUM-II

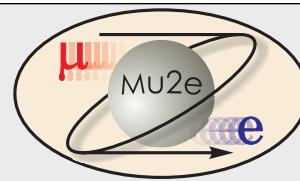
- $R_{\mu e} < 6.1 \times 10^{-13}$  in Au
- Want to probe to  $6 \times 10^{-17}$
- $\approx 10^4$  improvement



**Experimental signature is 105 MeV  $e^-$  originating in a thin Ti stopping target**



# SINDRUM-II Results



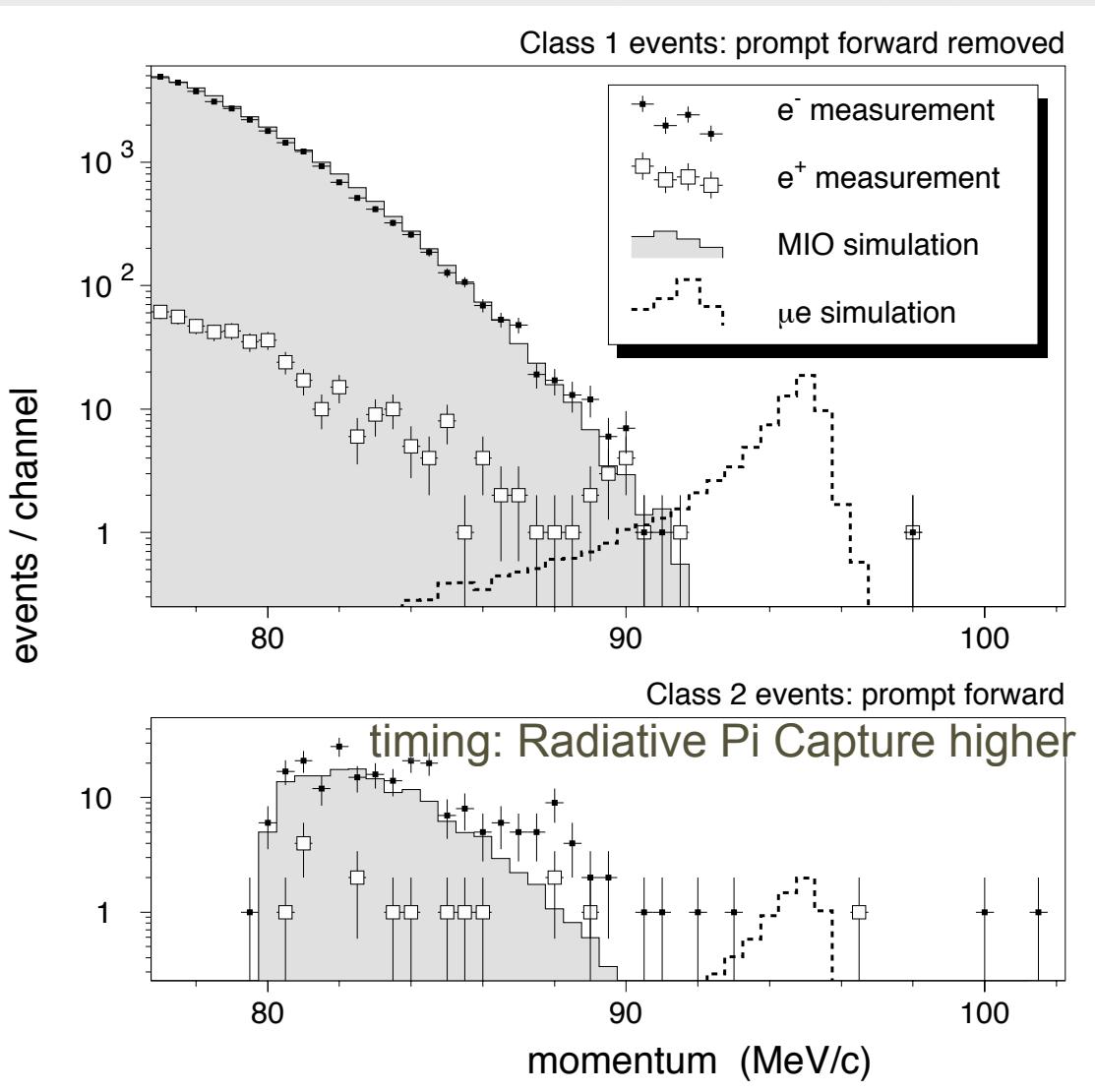
W. Bertl et al., Eur. Phys. J. C 47, 337–346 (2006)

- Final Results on Au:

$$B_{\mu e}^{\text{Au}} < 7 \times 10^{-13} \text{ @ 90% CL}$$

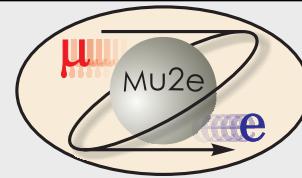
**51 MHz (20 nsec)  
repetition rate,  
width of pulse  
~0.3 nsec**

little time separation  
between  
signal and prompt  
background





# How Can We Do Better?



>10<sup>3</sup> increase in muon intensity from SINDRUM

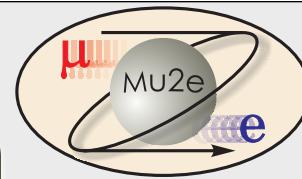
## *Requiring*

Pulsed Beam to Eliminate prompt backgrounds like  
radiative  $\pi$  capture and CR

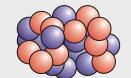
protons out of beam pulse/ protons in beam-pulse < 10<sup>-10</sup>  
*and we must measure it*



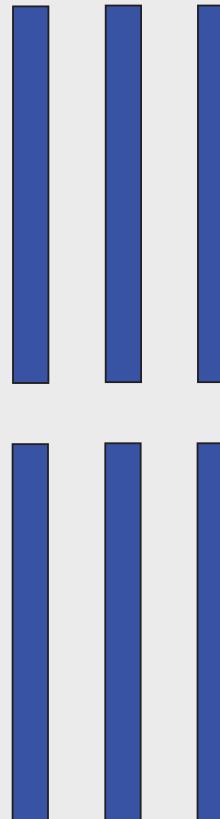
# Advantage of Pulsed Beam



target foils: muon converts here

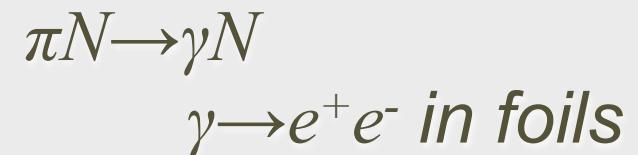


= muons, electrons, pions



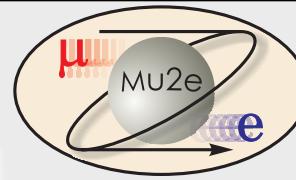
pulsed beam lets us  
wait until after prompt  
backgrounds  
disappear and rate  
lowered

RPC:

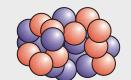




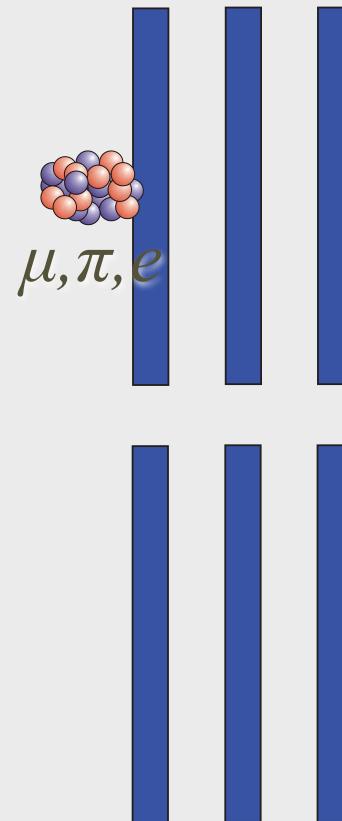
# Advantage of Pulsed Beam



target foils: muon converts here

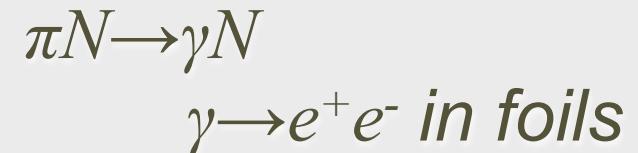


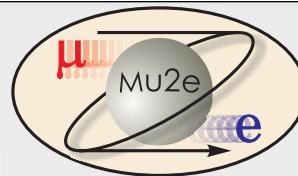
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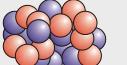
RPC:

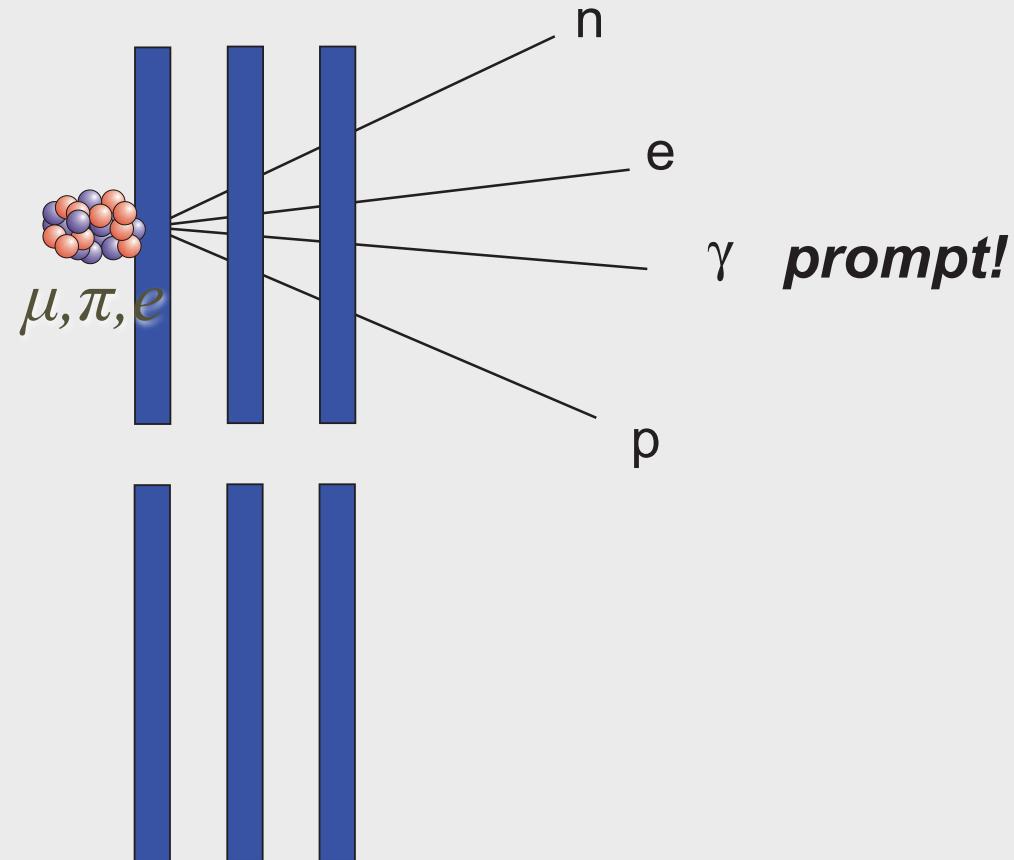




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target foils: muon converts here

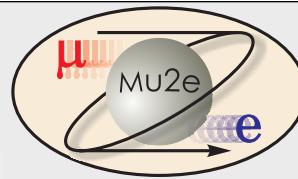
 = muons, electrons, pions



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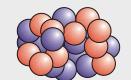
RPC:

$$\pi N \rightarrow \gamma N$$
$$\gamma \rightarrow e^+ e^- \text{ in foils}$$

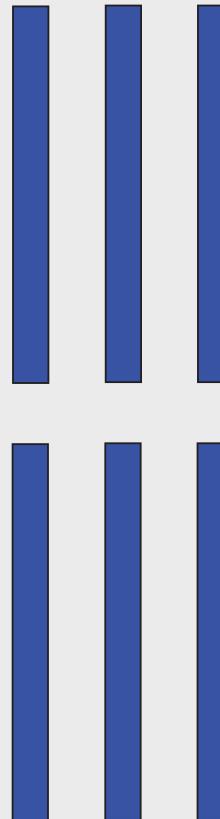


# Advantage of Pulsed Beam

target foils: muon converts here

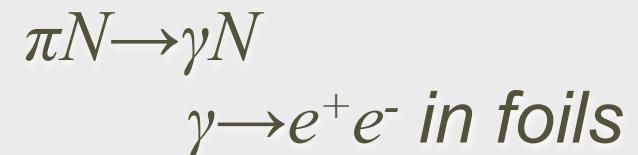


= muons, electrons, pions



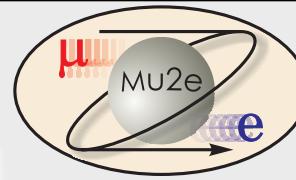
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RPC:

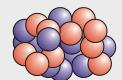




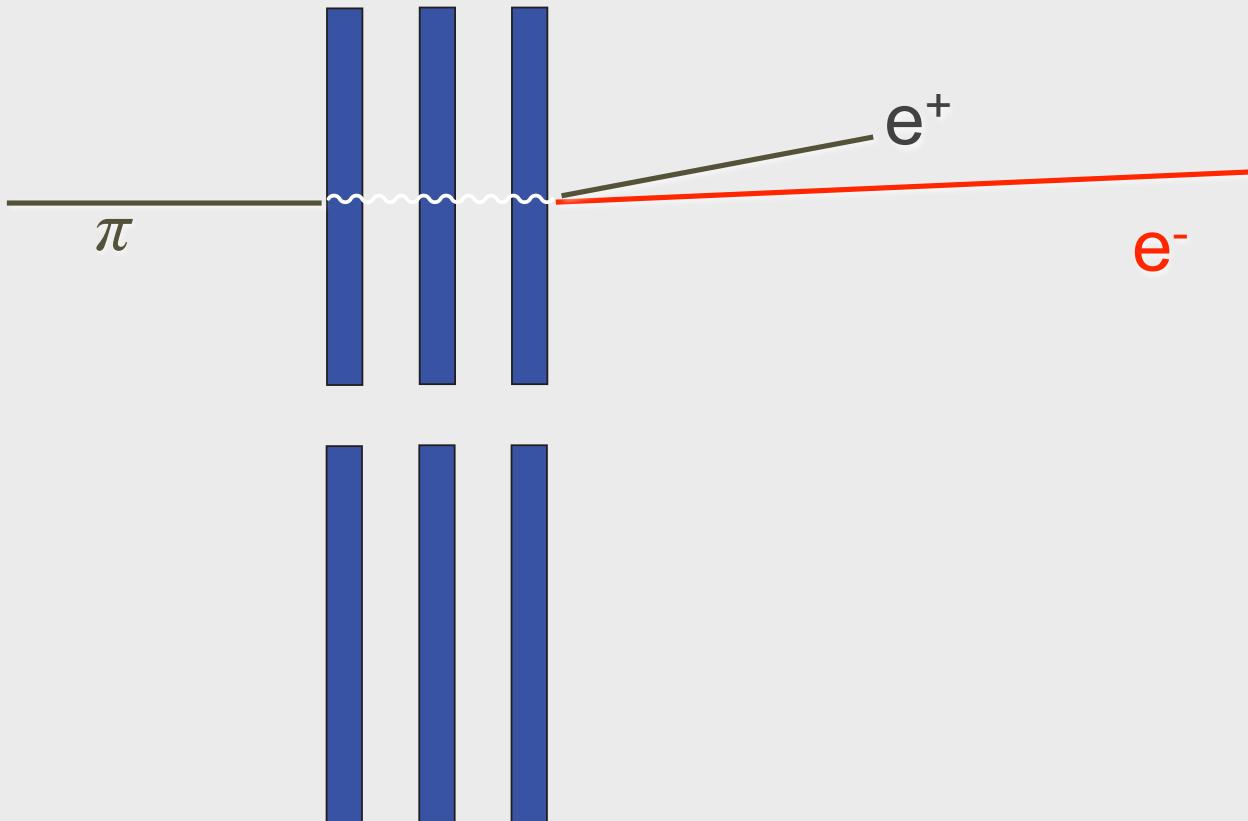
# Advantage of Pulsed Beam



target foils: muon converts here

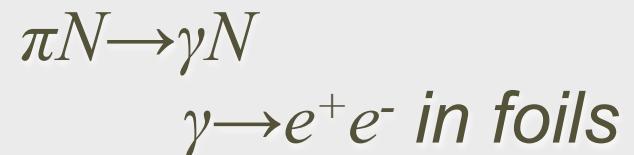


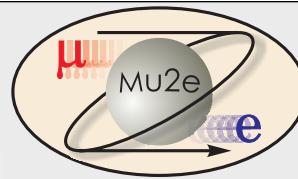
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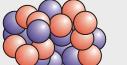
RPC:

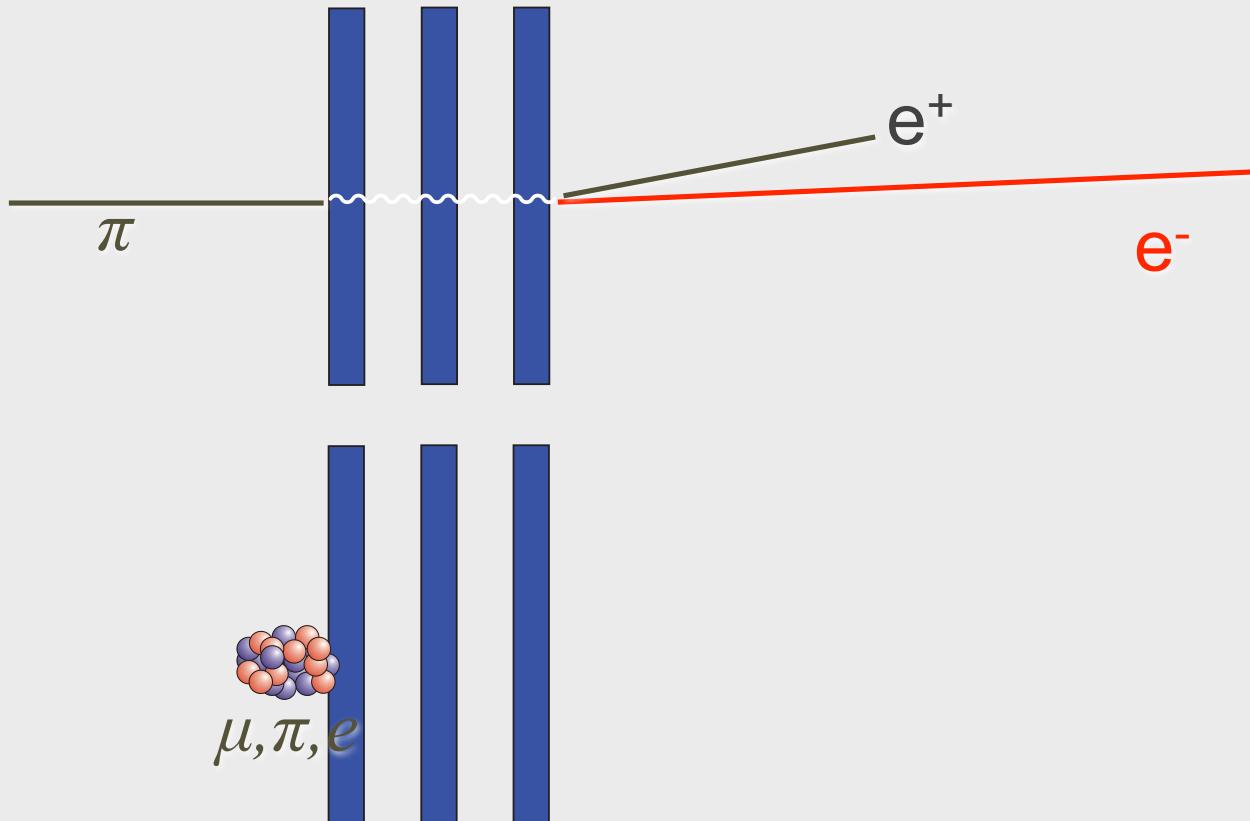




# Advantage of Pulsed Beam

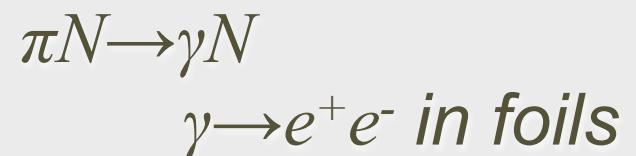
target foils: muon converts here

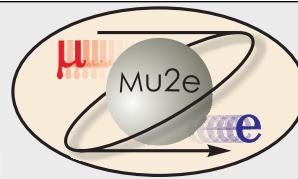
 = muons, electrons, pions



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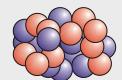
RPC:



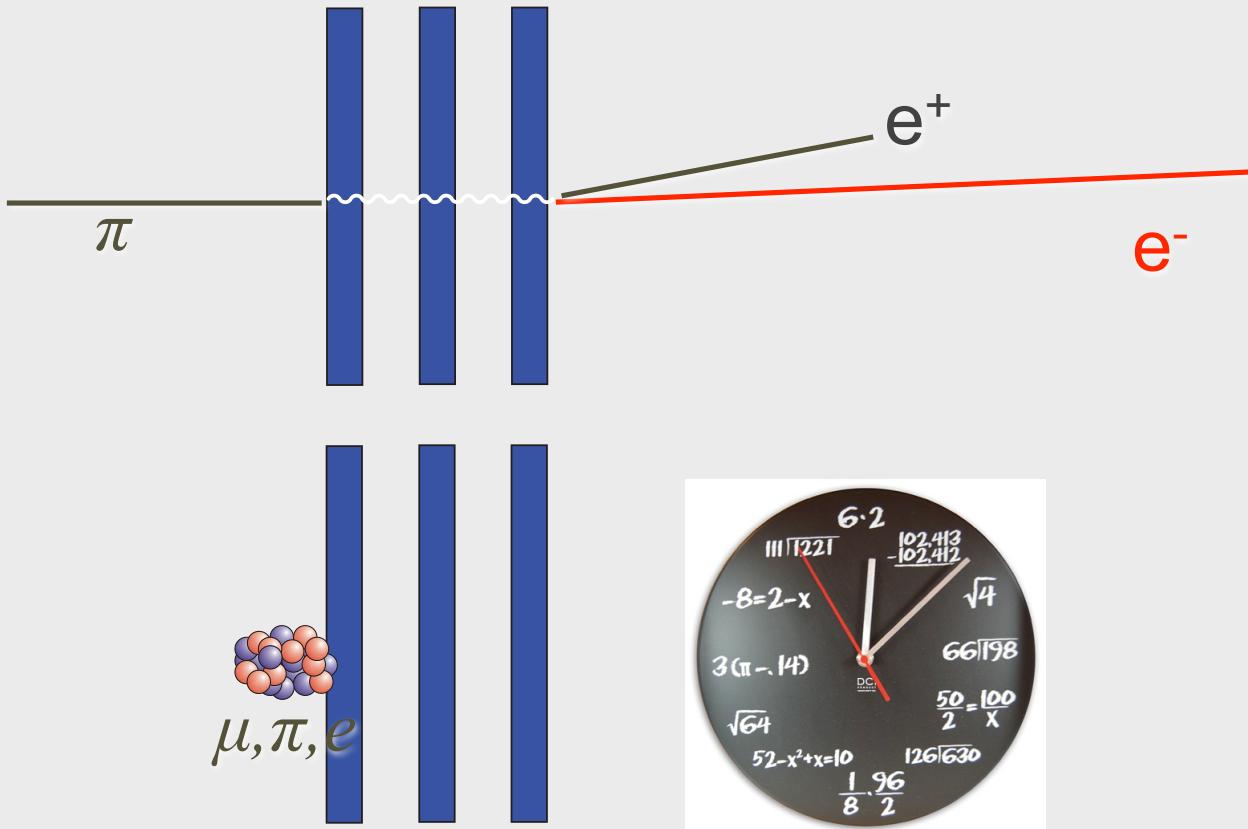


# Advantage of Pulsed Beam

target foils: muon converts here



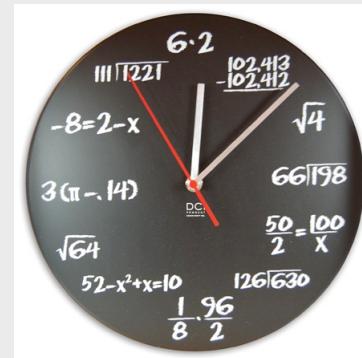
= muons, electrons, pions

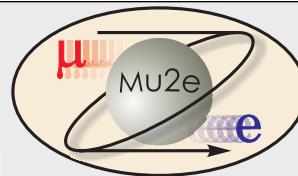


pulsed beam lets us  
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lowered

RPC:

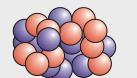
$\pi N \rightarrow \gamma N$   
 $\gamma \rightarrow e^+ e^- \text{ in foils}$

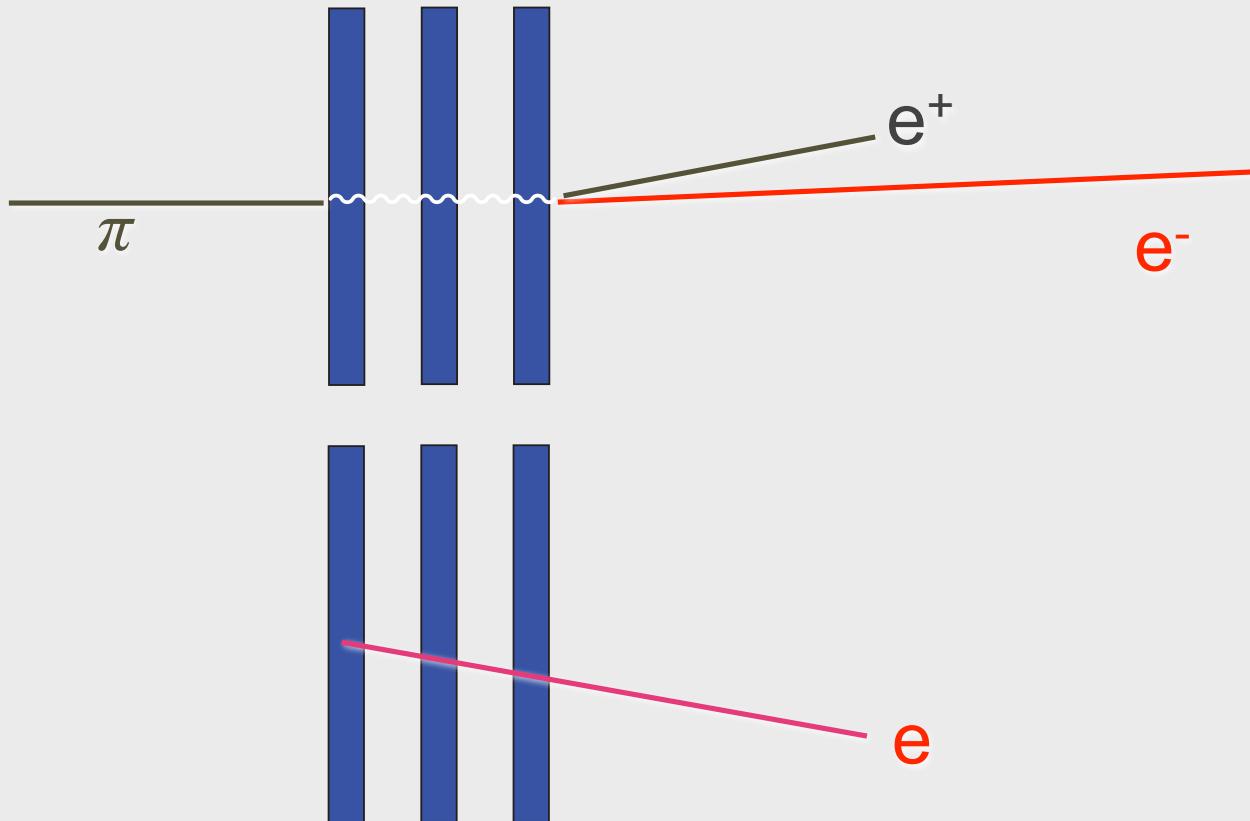




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 = muons, electrons, pions



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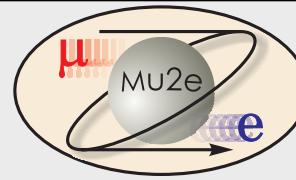
RPC:

$\pi N \rightarrow \gamma N$   
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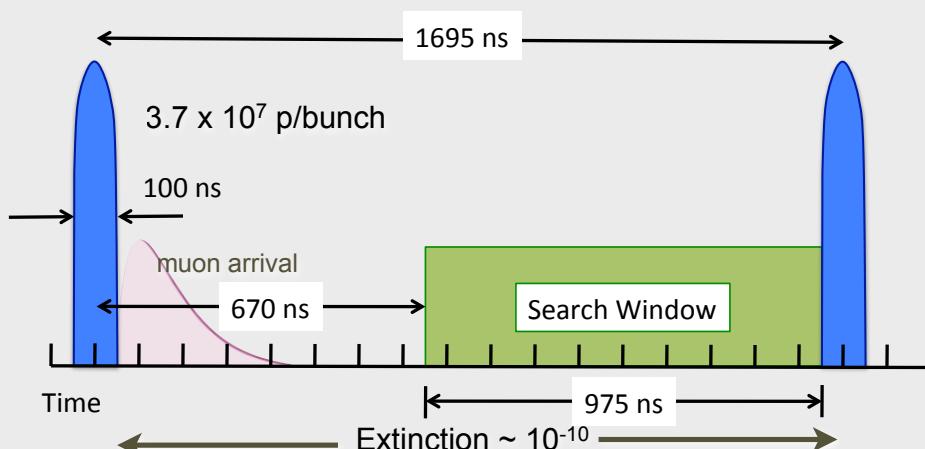
delayed 105 MeV electron



# Pulsed Beam Structure



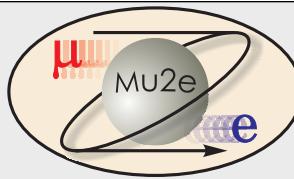
- Tied to prompt rate and machine: FNAL “perfect”
- Want **pulse duration  $\ll \tau_\mu^{\text{Al}}$** , **pulse separation  $\approx \tau_\mu^{\text{Al}}$** 
  - FNAL Debuncher has circumference  **$1.7\mu\text{sec}$** ,  $\sim \times 2 \tau_\mu^{\text{Al}}$
- Extinction between pulses  $< 10^{-10}$  needed
  - = # protons out of pulse/# protons in pulse



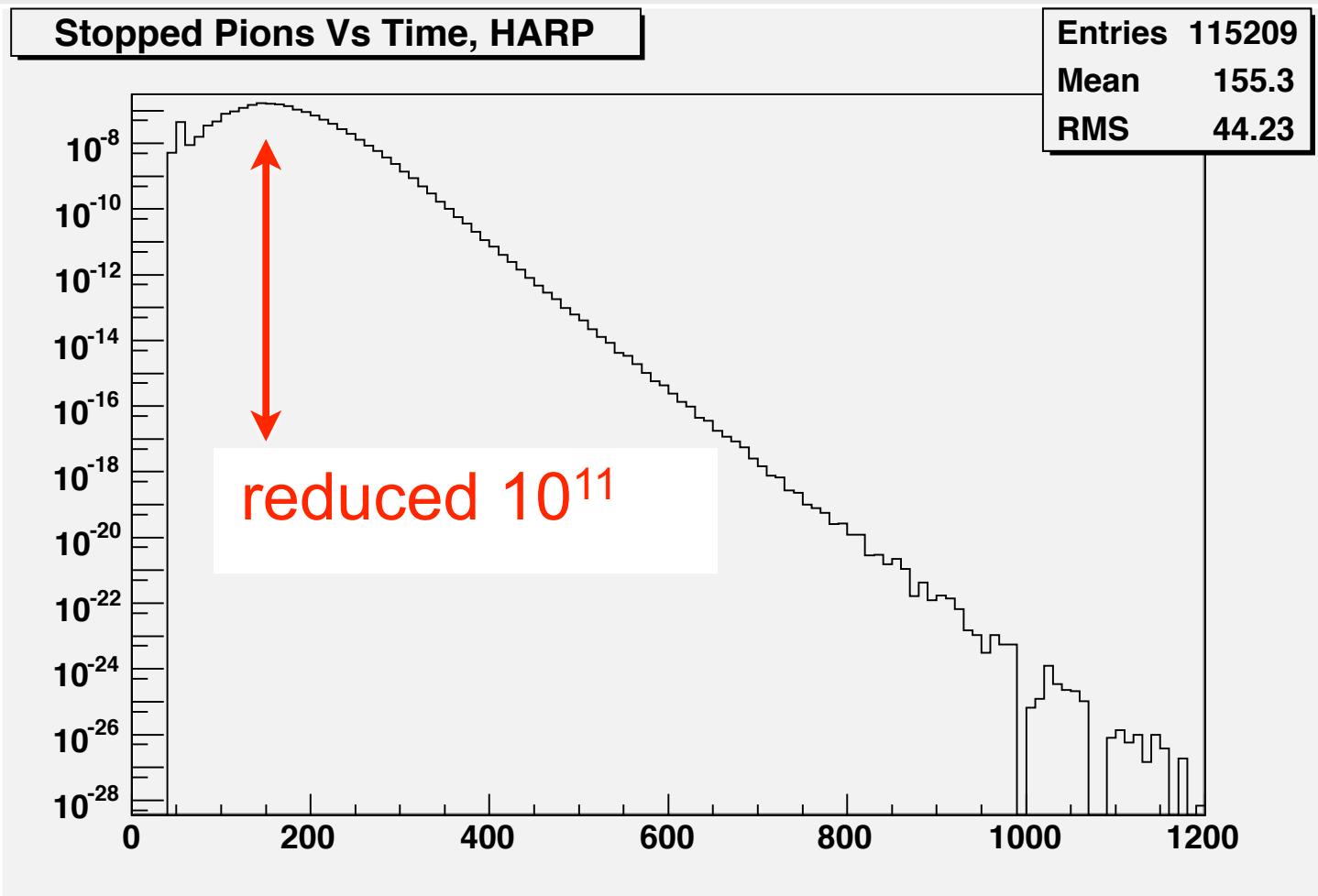
- $10^{-10}$  based on simulation of prompt backgrounds and beamline



# Pulsed Beam Structure and Radiative $\pi$ Capture



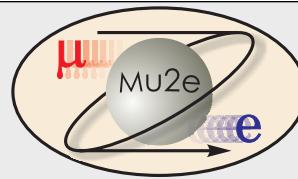
$$\pi N \rightarrow \gamma N, \gamma \rightarrow e^+ e^-$$



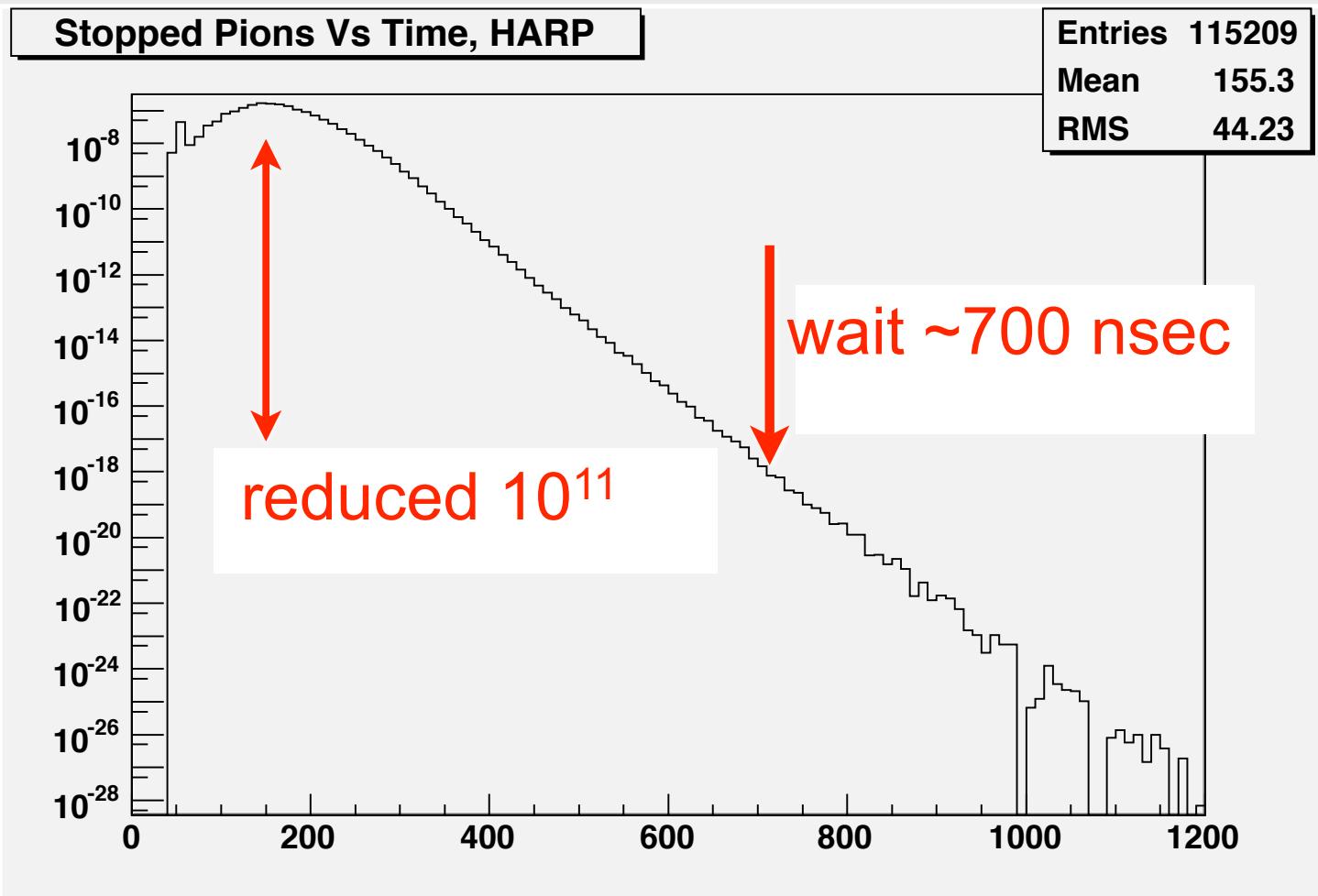
*need a beam that lets us wait this long: FNAL*



# Pulsed Beam Structure and Radiative $\pi$ Capture



$$\pi N \rightarrow \gamma N, \gamma \rightarrow e^+ e^-$$



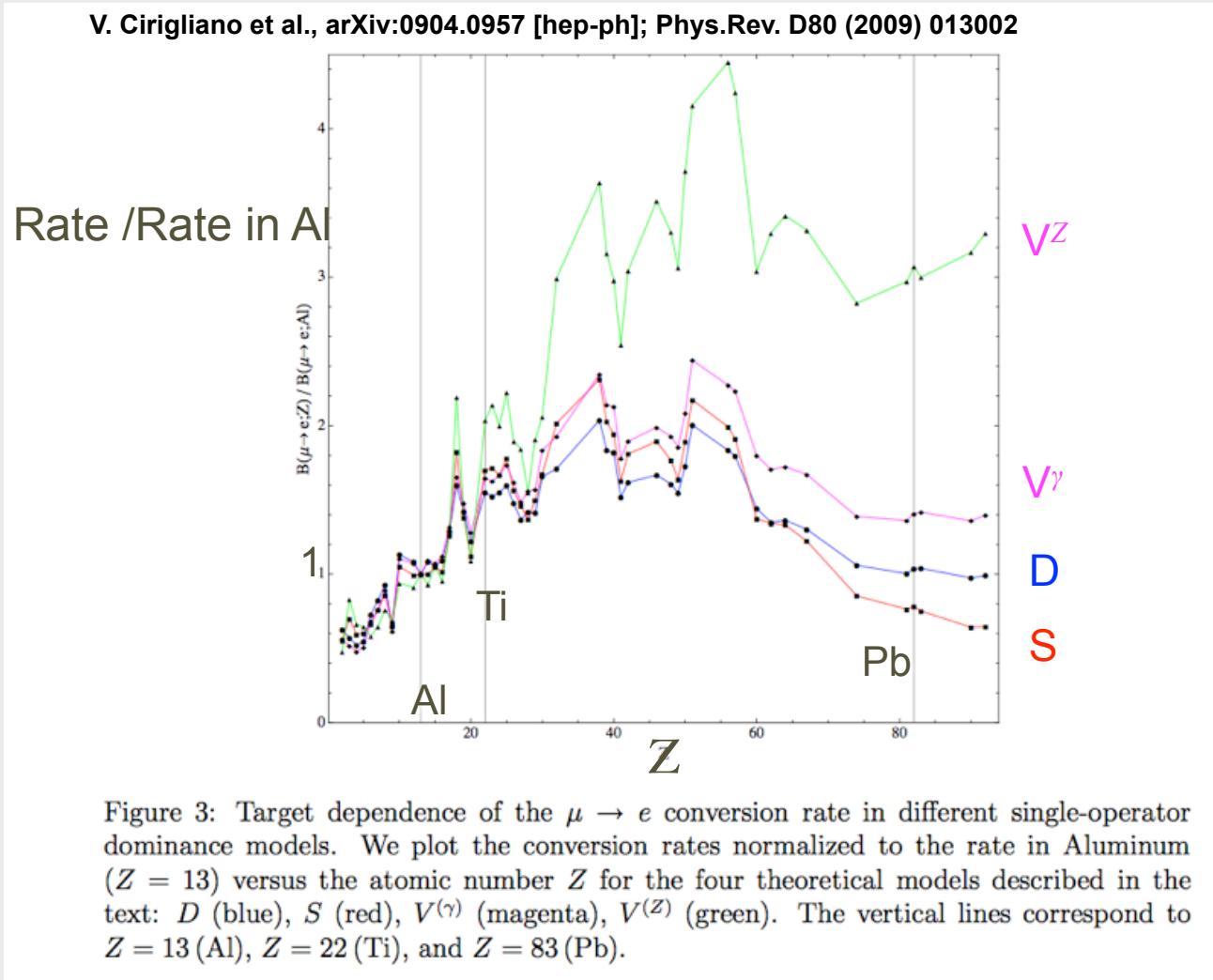
*need a beam that lets us wait this long: FNAL*



# Choice of Stopping Material: rate vs wait



- Stop muons in target ( $Z, A$ )
- Physics sensitive to  $Z$ : with signal, can switch target to probe source of new physics
- Why start with Al?



shape governed by relative conversion/capture rate, form factors, ...

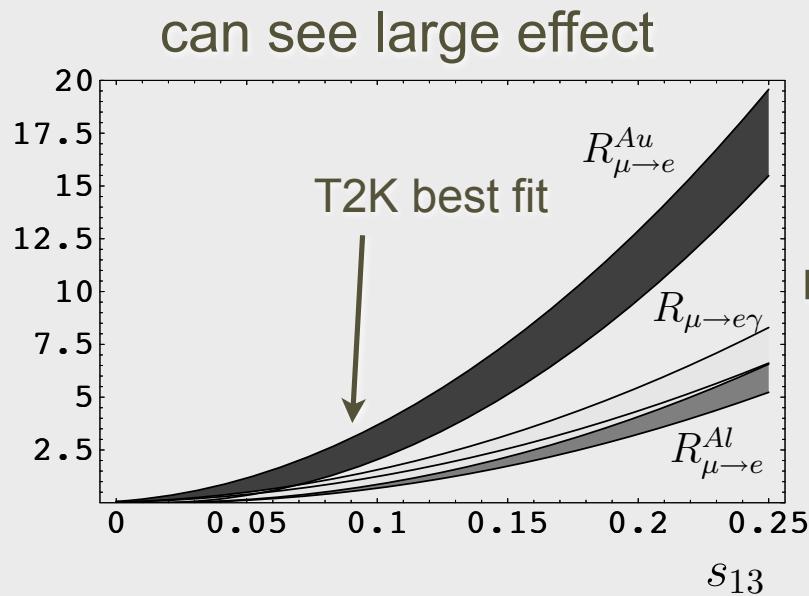


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V. Cirigliano, B. Grinstein, G. Isidori, M. Wise **Nucl.Phys.B728:121-134,2005.**  
e-Print: [hep-ph/0507001](https://arxiv.org/abs/hep-ph/0507001)

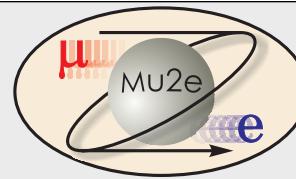


$s_{13}$  is NOvA  
mixing angle  
 $< 0.2$  or so

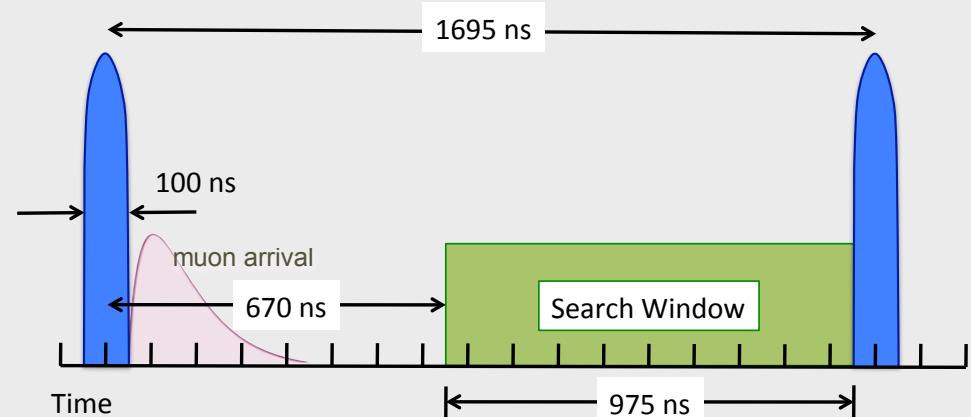
shape governed by relative conversion/capture rate, form factors, ...



# Prompt Background and Choice of Z



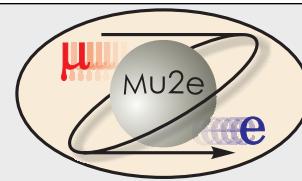
choose Z based on tradeoff  
between rate and lifetime:  
longer lived reduces prompt  
backgrounds



Nucleus	$R_{\mu e}(Z) / R_{\mu e}(\text{Al})$	Bound Lifetime	Conversion Energy
Al(13,27)	1.0	864 nsec	104.96 MeV
Ti(22,~48)	1.7	328 nsec	104.18 MeV
Au(79,~197)	~0.8-1.5	72.6 nsec	95.56 MeV

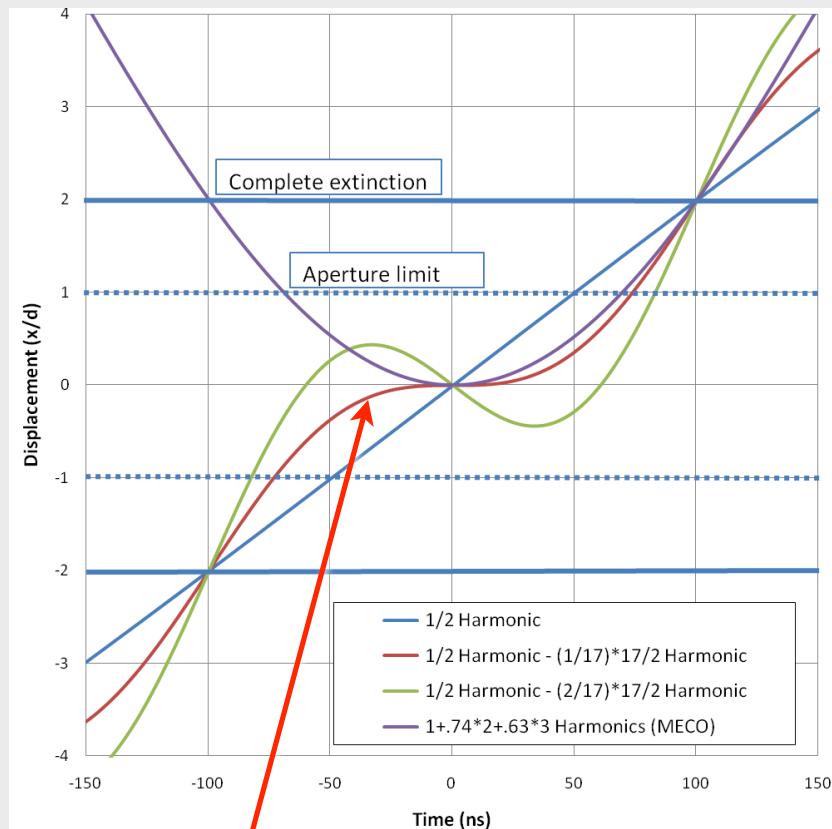


# Extinction Scheme

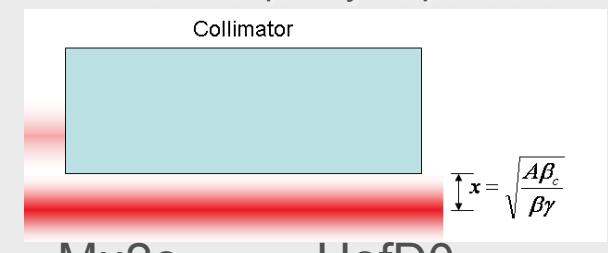


*achieving  $10^{-10}$  is hard; normally get  $10^{-2} - 10^{-3}$*

- Internal (momentum scraping) and bunch formation in Accumulator
- External: oscillating (AC) dipole
  - high frequency (300 KHz) dipole with smaller admixture of 17th harmonic (5.1 MHz)
  - Sweep Unwanted Beam into collimators
- Calculations (MARS) show this combination gets  $\sim 10^{-12}$

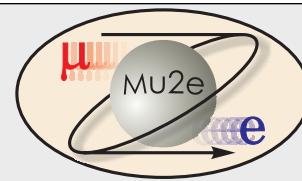


choose a field as flat as possible  
during the pulse that kicks beam  
out as quickly as possible

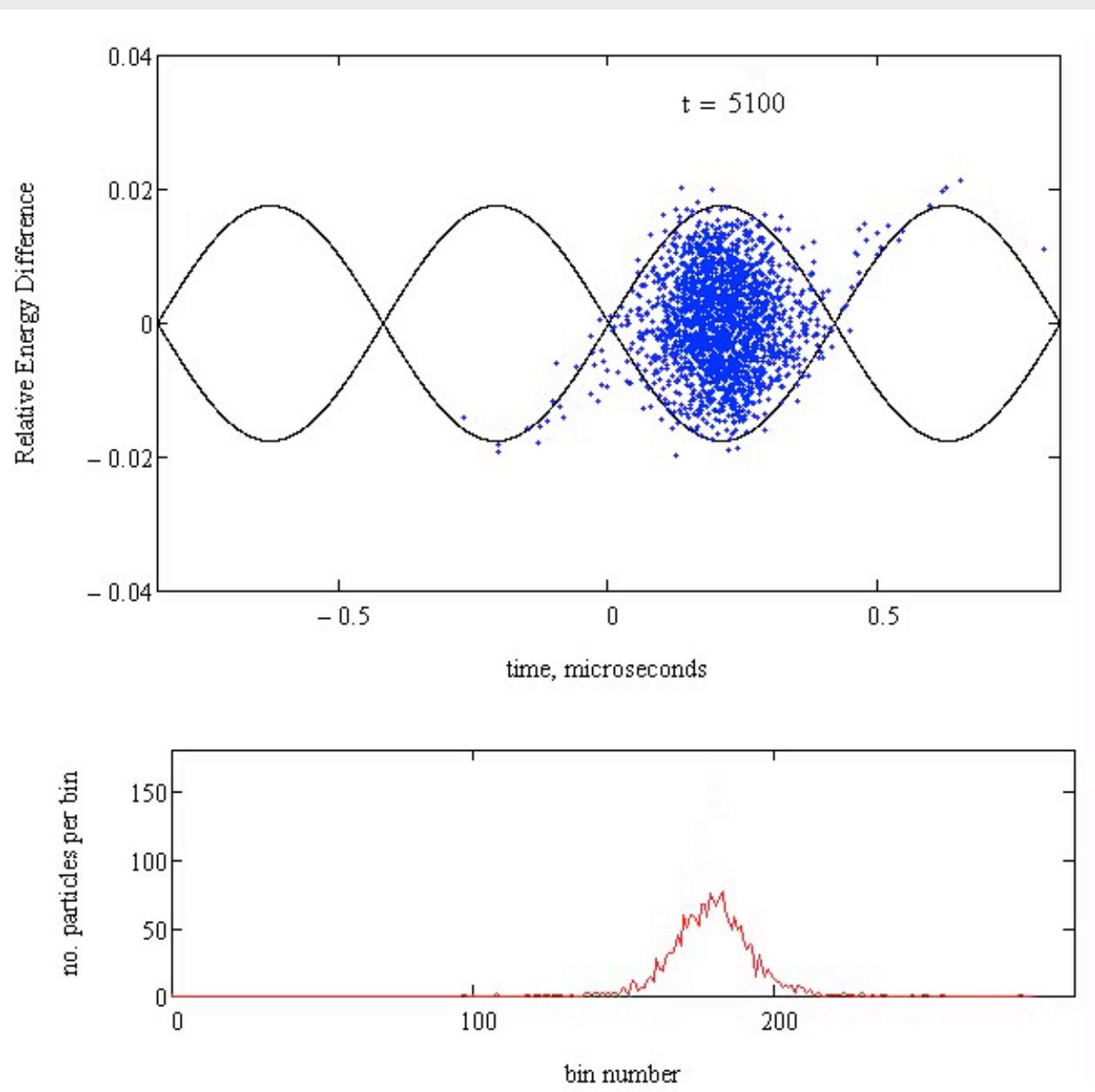




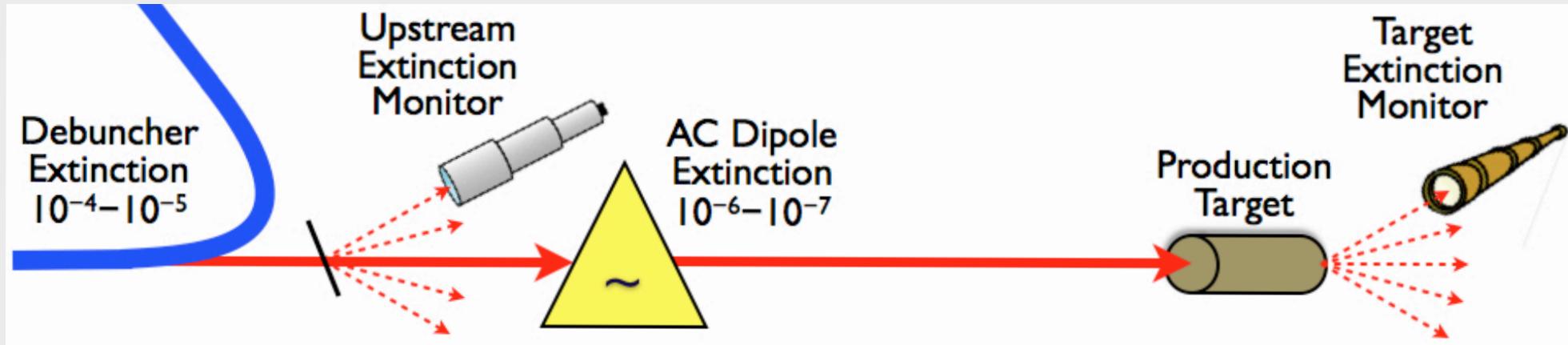
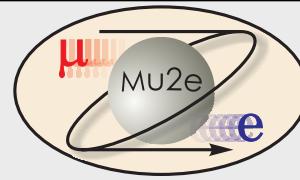
# Extinction (Internal)



- RF jitter, scattering, etc can cause protons to wander out of their bucket



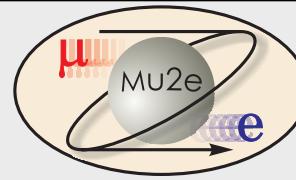
# Extinction Measurement



- Beam Formation(Debuncher, Momentum Scraping) and AC-Dipole provide extinction
- Measure with:
  - Thin foils in 8 GeV transport line (fast feedback on machine performance)
  - Off-axis telescope looking at production target (time scale of hour)
  - Also considering detecting individual out-of-time protons but very difficult



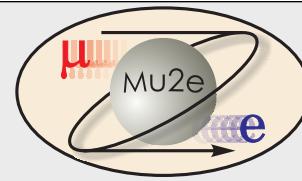
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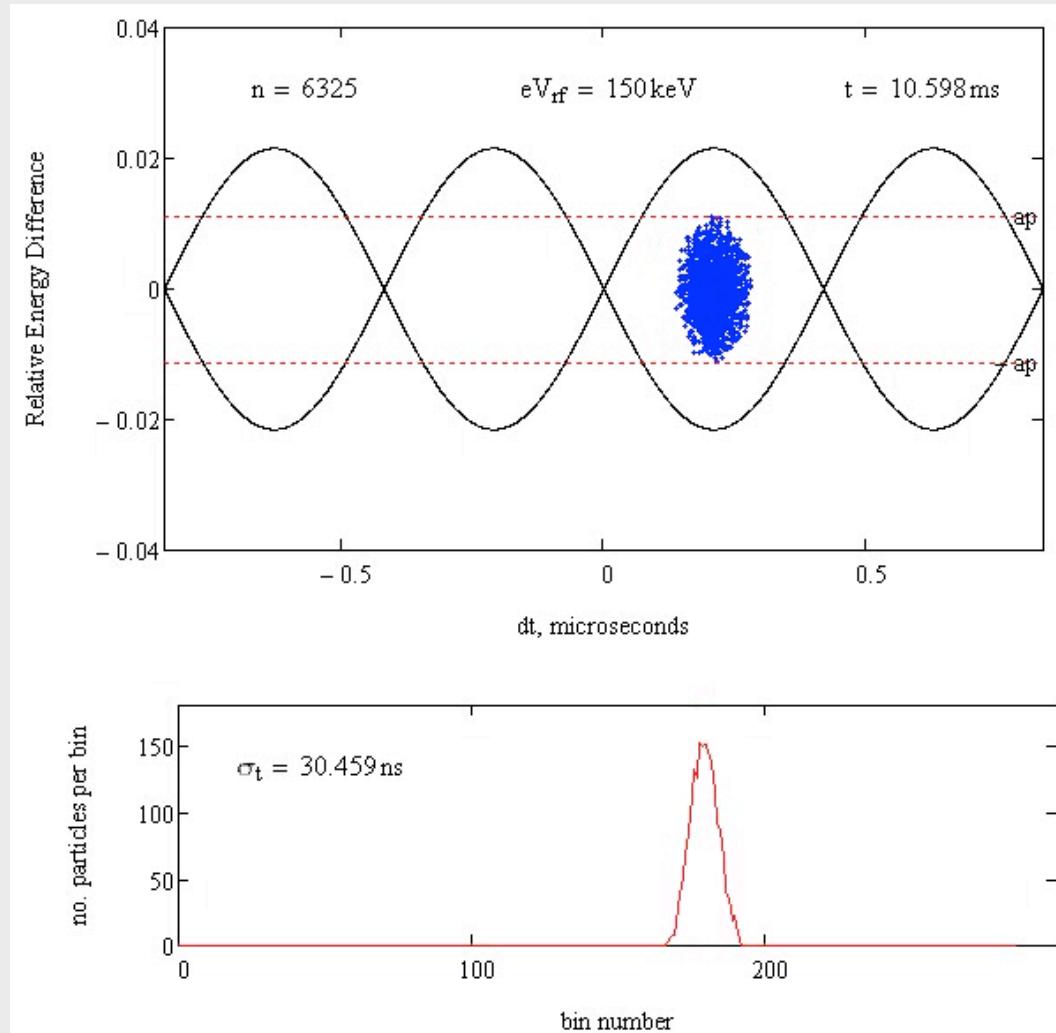
- *Internal*: “momentum scraping”: wait for beam to be wide, then use collimators
- Perform in Debuncher
- Have also modified beam transfer scheme from Accumulator to Debuncher to improve cleanliness



# Extinction (internal)

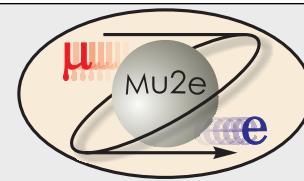


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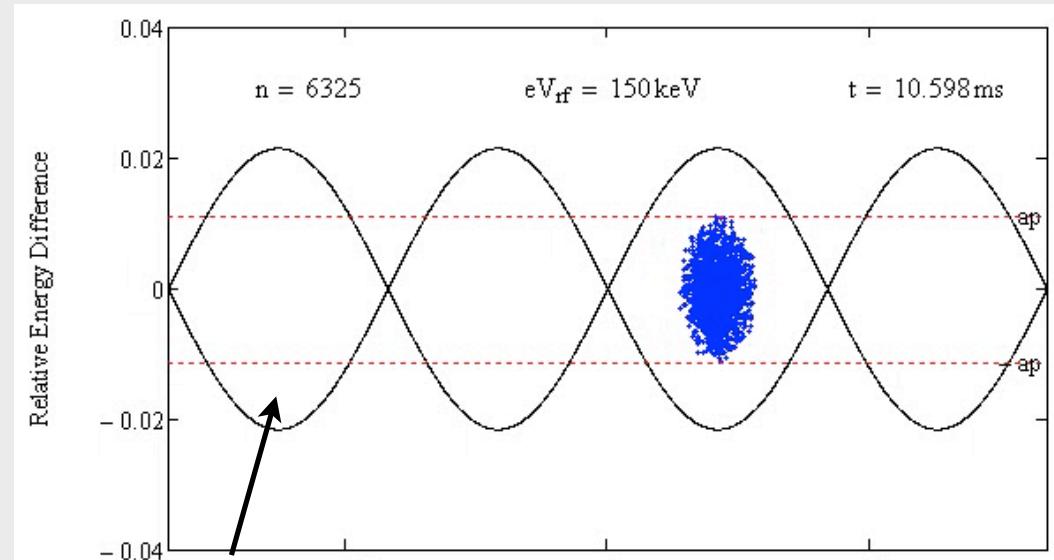




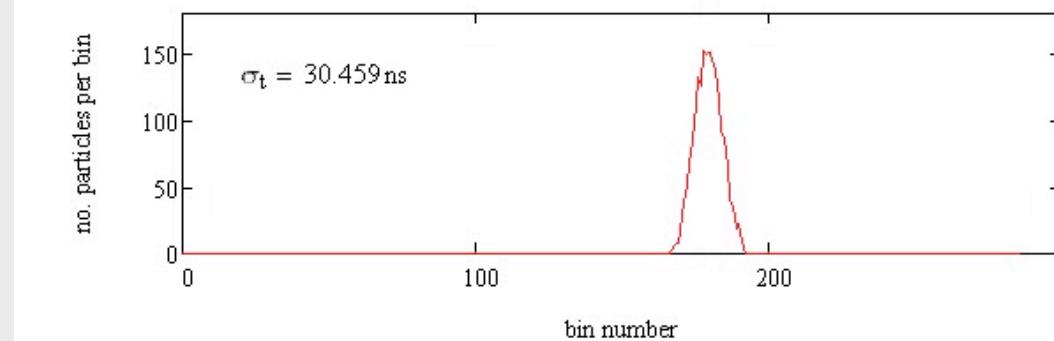
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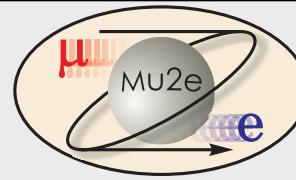


*momentum scrape:*  $|dE/E| = x_{max}/D$





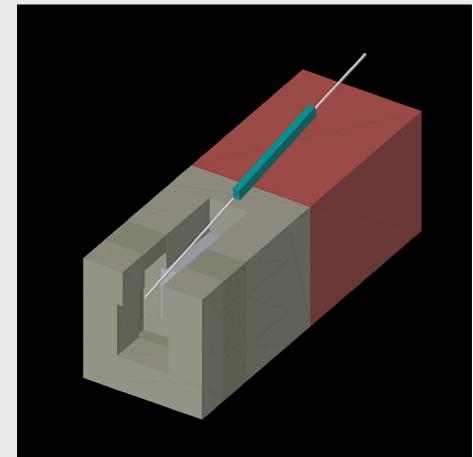
# Current Extinction Plan



- Look at secondaries from target
  - Use a magnet to pick out signal
  - Count hits in-and-out of pulse, compare
  - Current plan -- *not proton-by-proton*

Nominal Momentum (GeV/c)	Transmitted tracks / bunch		Integration Time (min)		Proton Percentage	MeV Deposited / Proton-on-Target		
			No Bkgnd.	5 / hr		Entry Channel	Magnet	Exit Channel
2.0	40.5	± 4.4	6.9	10.8	77.6	0.293	0.167	0.007
3.0	151.4	± 8.5	1.8	2.1	82.7	0.258	0.168	0.037
4.0	240.0	± 10.7	1.2	1.3	89.9	0.246	0.164	0.037
5.0	191.0	± 9.5	1.5	1.6	93.0	0.201	0.128	0.037
6.0	119.1	± 5.8	2.3	2.8	96.4	0.172	0.112	0.026
7.0	69.4	± 4.5	4.0	5.3	97.1	0.149	0.100	0.021
8.0	39.8	± 2.9	7.0	11.0	99.0	0.133	0.095	0.015
9.0	26.1	± 2.3	10.6	20.1	100.0	0.116	0.089	0.011

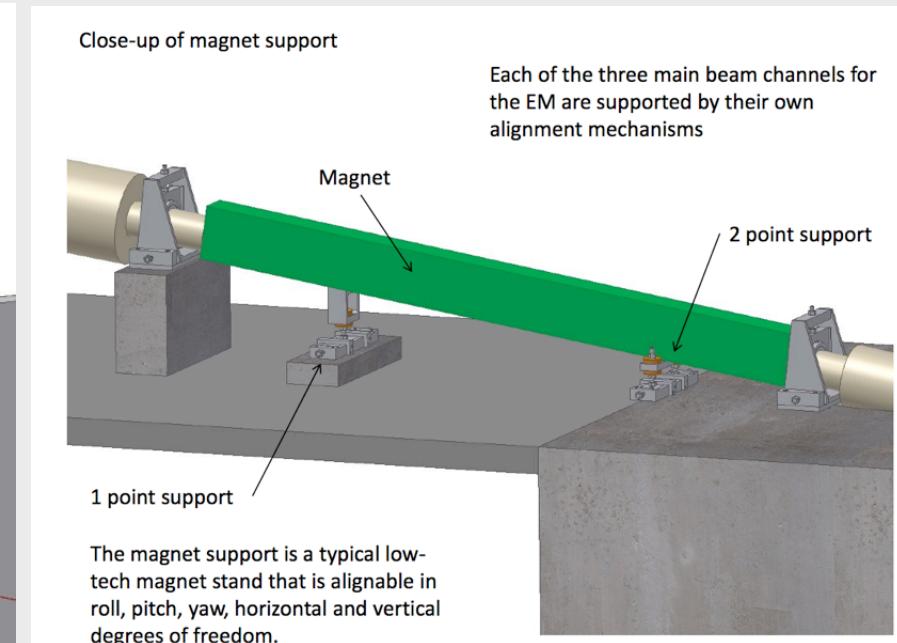
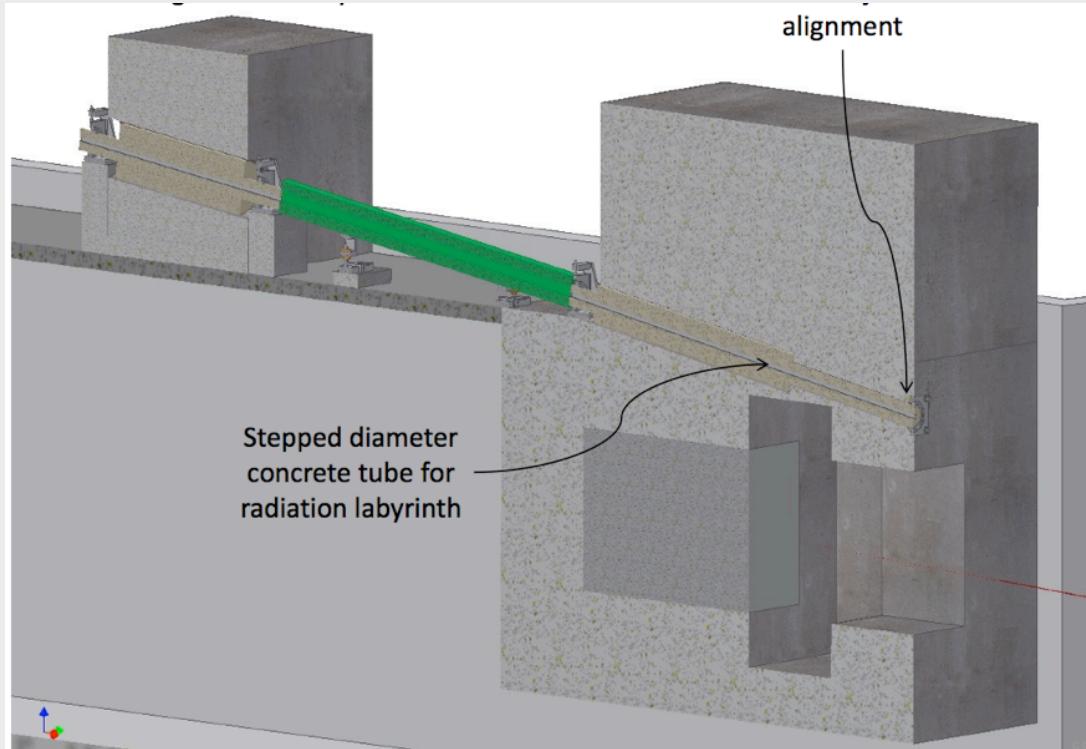
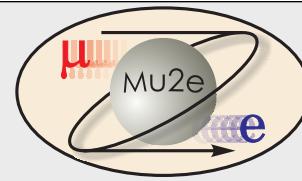
G4Beamline Model



working on using diffractive protons:  
little-to-no background, easier to model

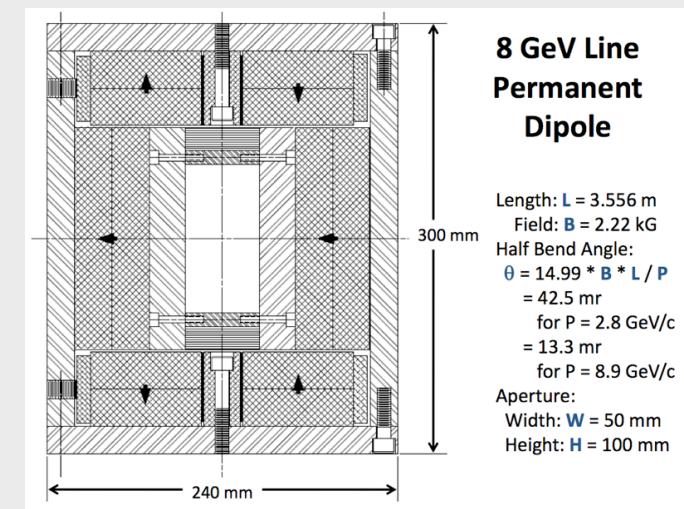


# Wide Variety of choices



- Working on detector modeling
  - ATLAS pixels?
  - TOF?
  - Calorimeter?

*interesting problem  
for post-docs!!*

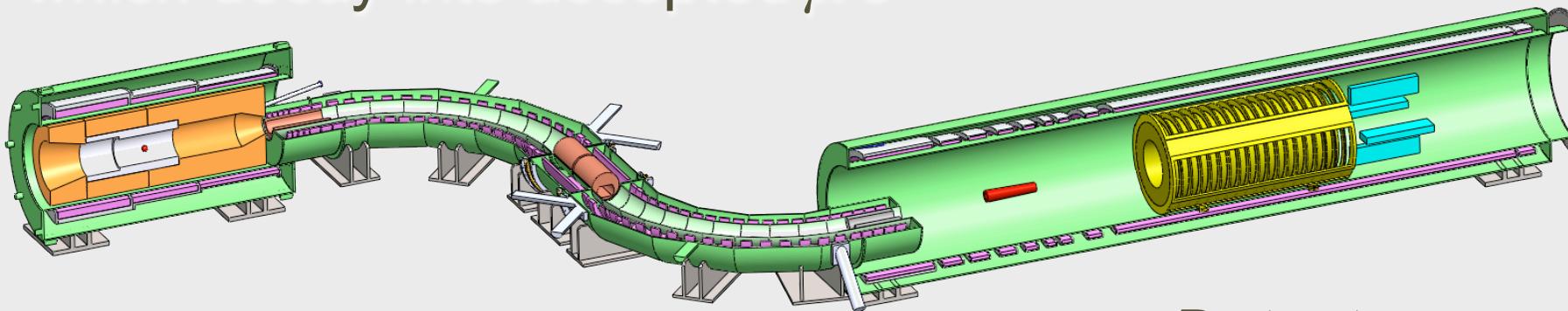
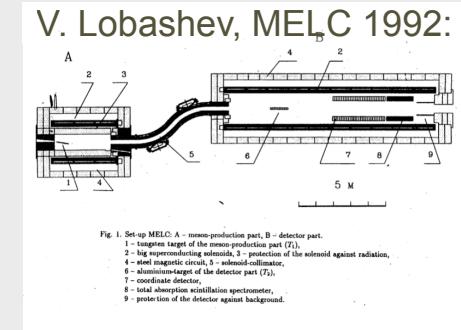




# Mu2e Overview



- *Production:* Magnetic bottle traps  $\pi$ 's, which decay into accepted  $\mu$ 's

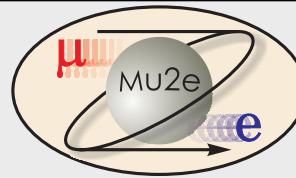


- *Transport:* S-curve eliminates backgrounds and sign-selects

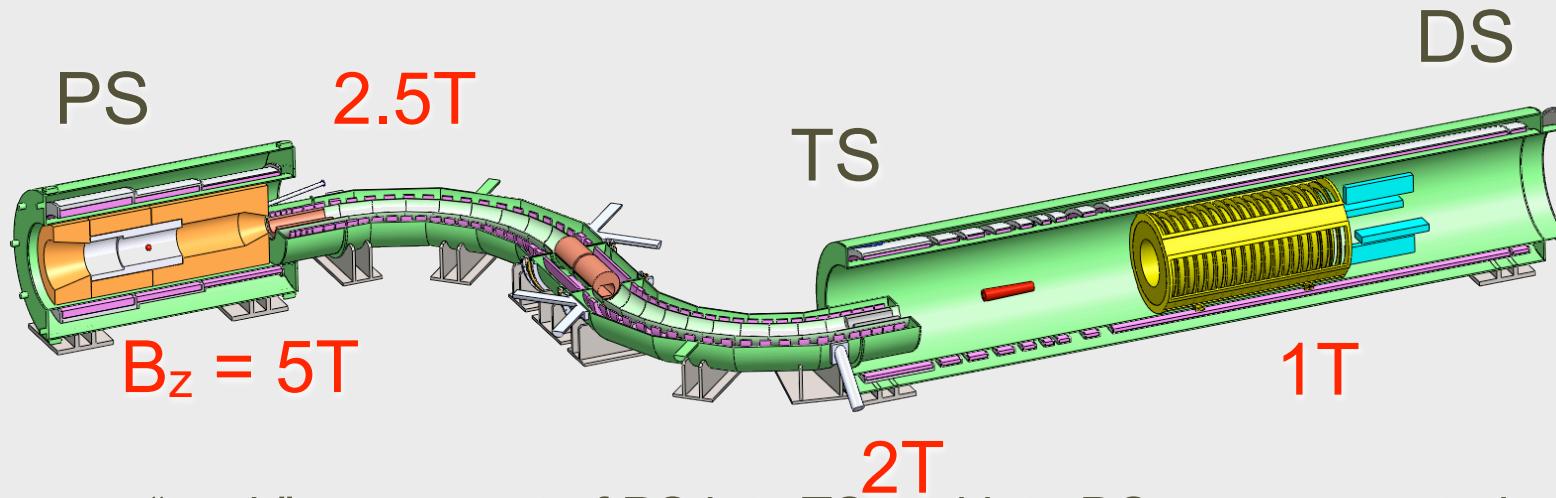
- *Detector:* Stopping Target, Tracking and Calorimeter



# Gradient Fields in Mu2e



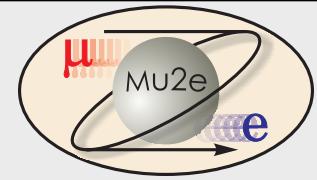
- Play a vital role throughout the design
- A lot of our intellectual effort!!



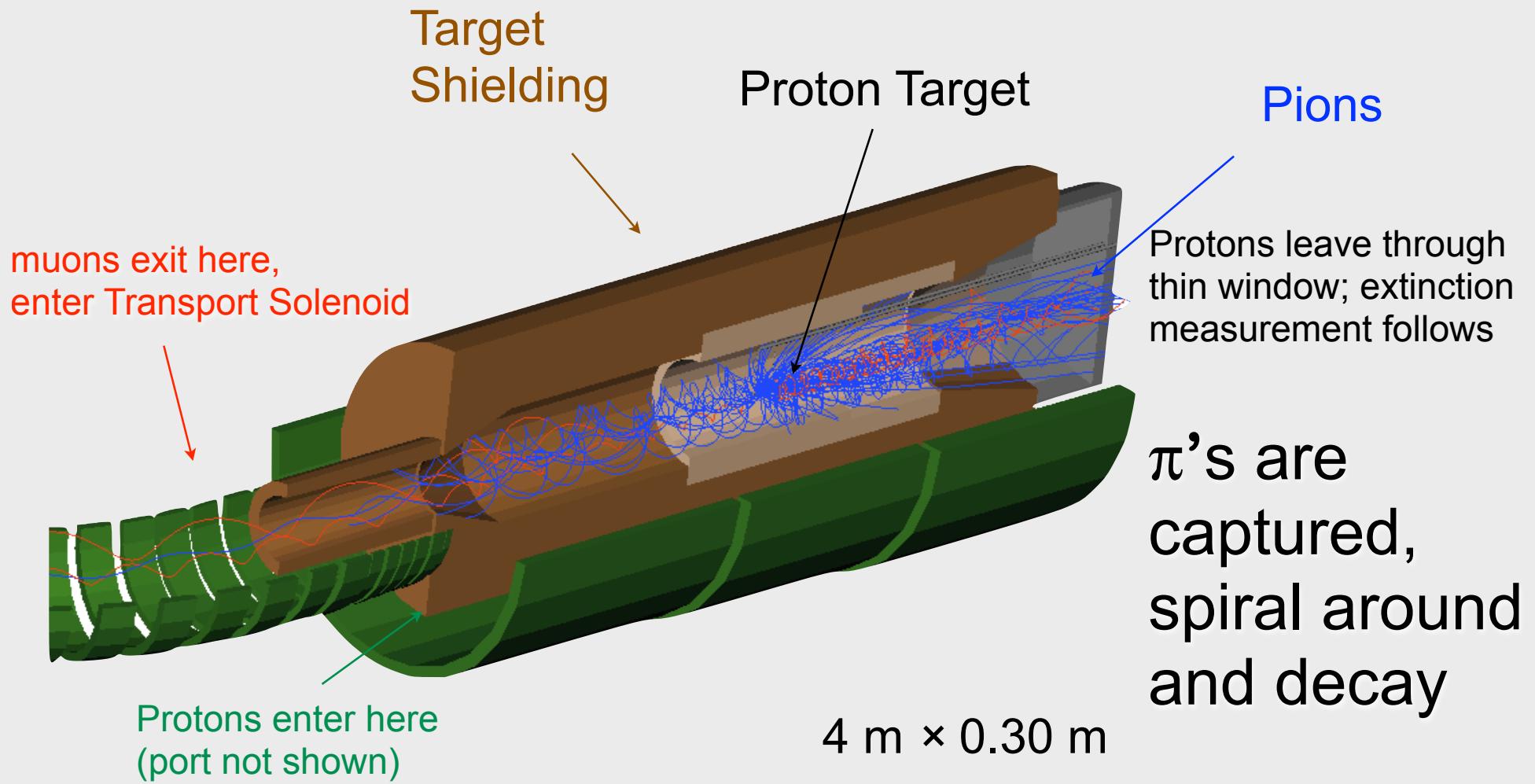
- “push” muons out of PS into TS and into DS so we can study them
- keep particles from spiraling around, arriving late
- conversions are isotropic in stopping target; the gradient over stopping target “reflects” backward going muons and nearly doubles the acceptance



# Production Solenoid:

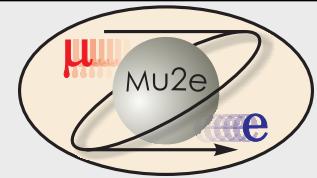


Protons enter opposite to outgoing muons

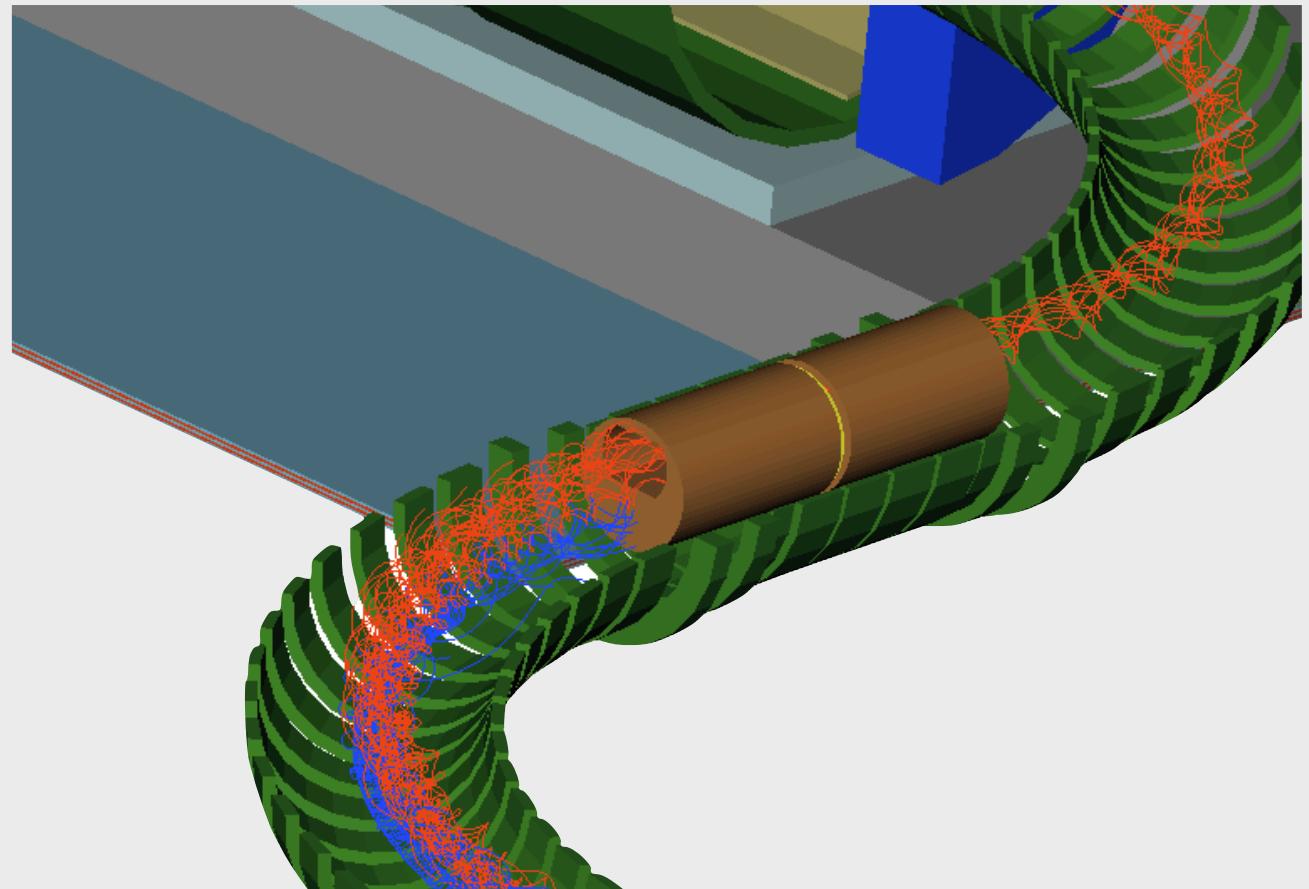




# Transport Solenoid



- Curved solenoid eliminates line-of-sight transport of photons and neutrons
- Curvature drift and collimators sign and momentum select beam



13.1 m along axis  $\times$   $\sim$ 0.25 m



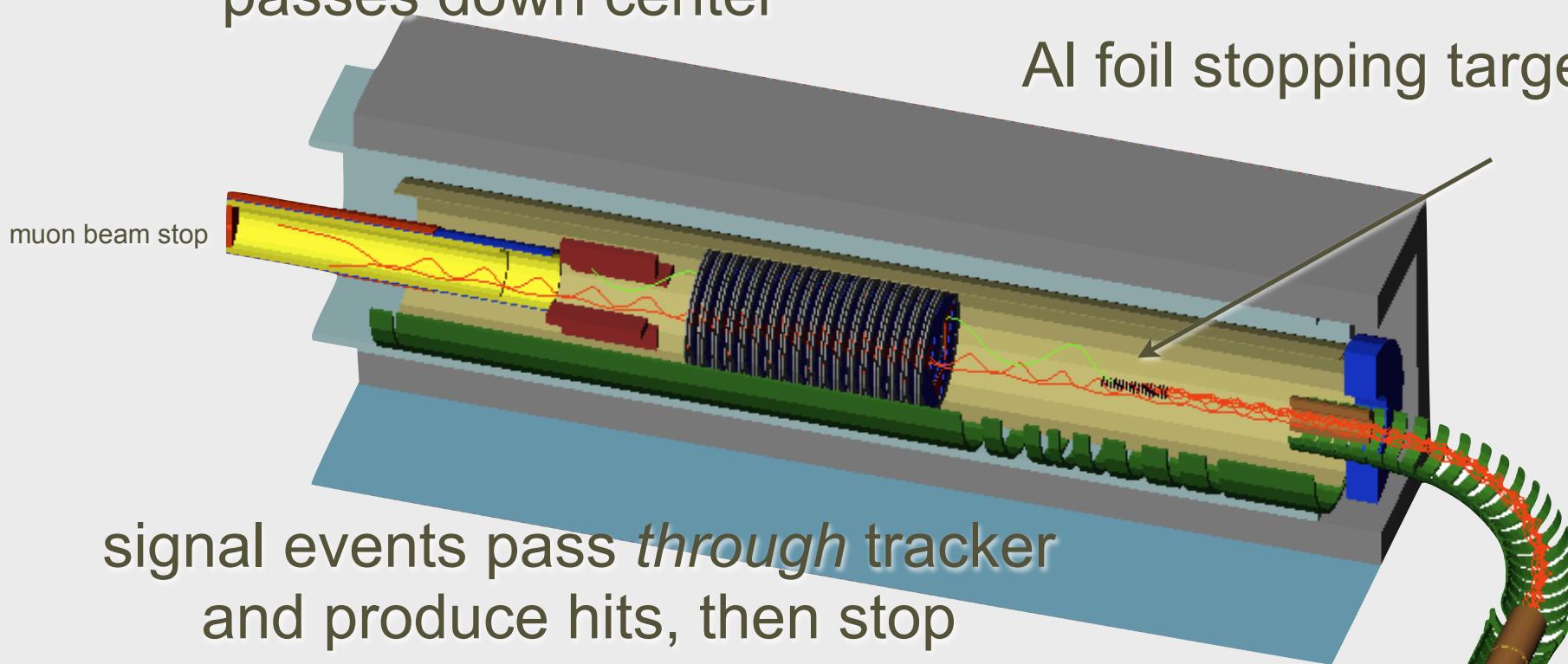
# Detector Solenoid

*octagonal tracker surrounding central region:  
radius of helix proportional to momentum,  
 $p=qBR$*

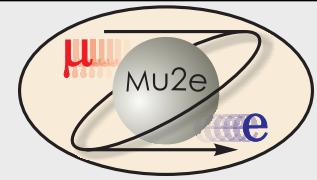
low momentum particles and  
almost all DIO background  
passes down center

10 m × 0.95 m

Al foil stopping target

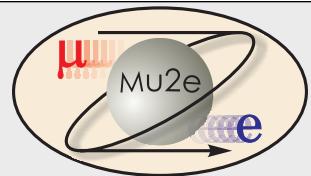


signal events pass *through* tracker  
and produce hits, then stop  
in calorimeter





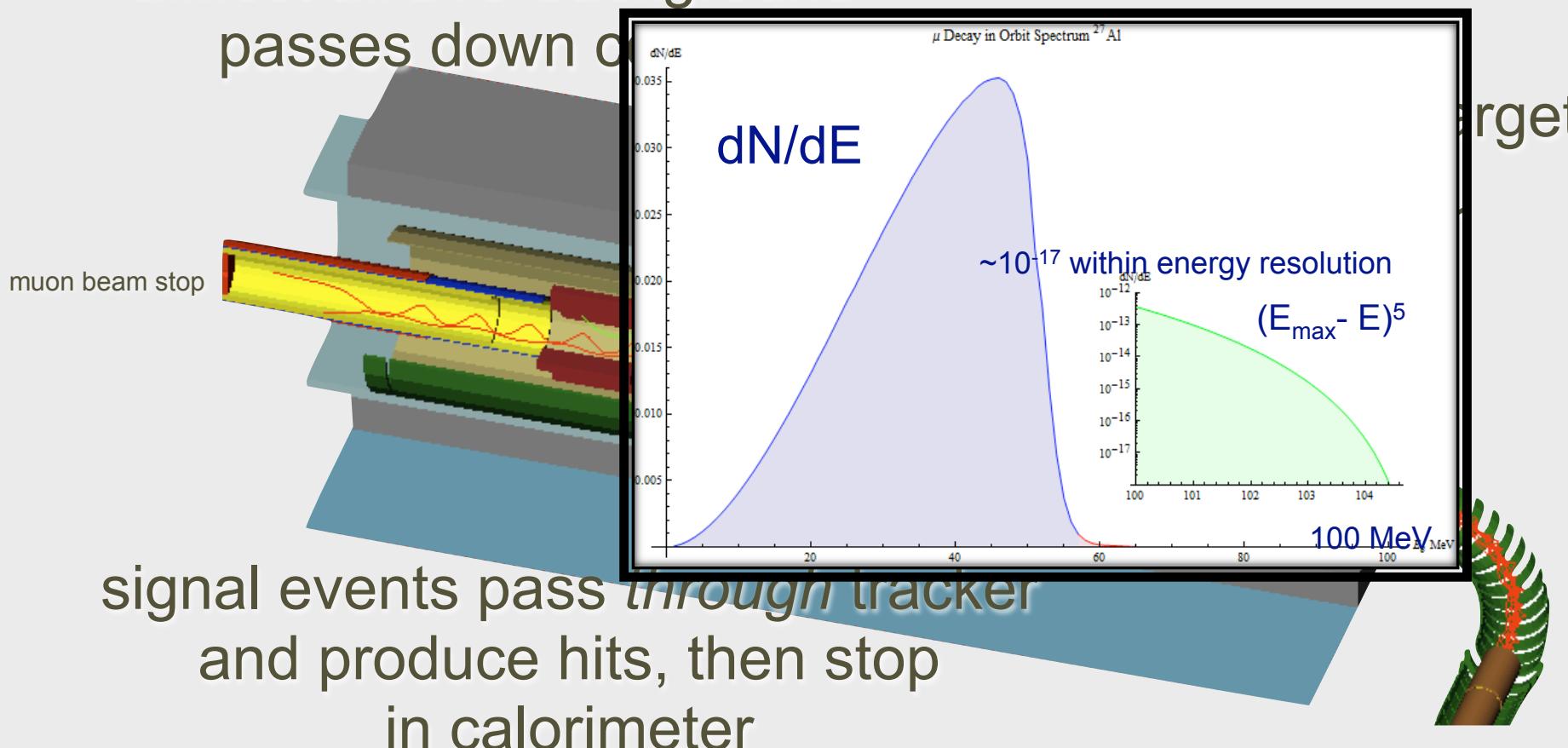
# Detector Solenoid



*octagonal tracker surrounding central region:  
radius of helix proportional to momentum,  
 $p=qBR$*

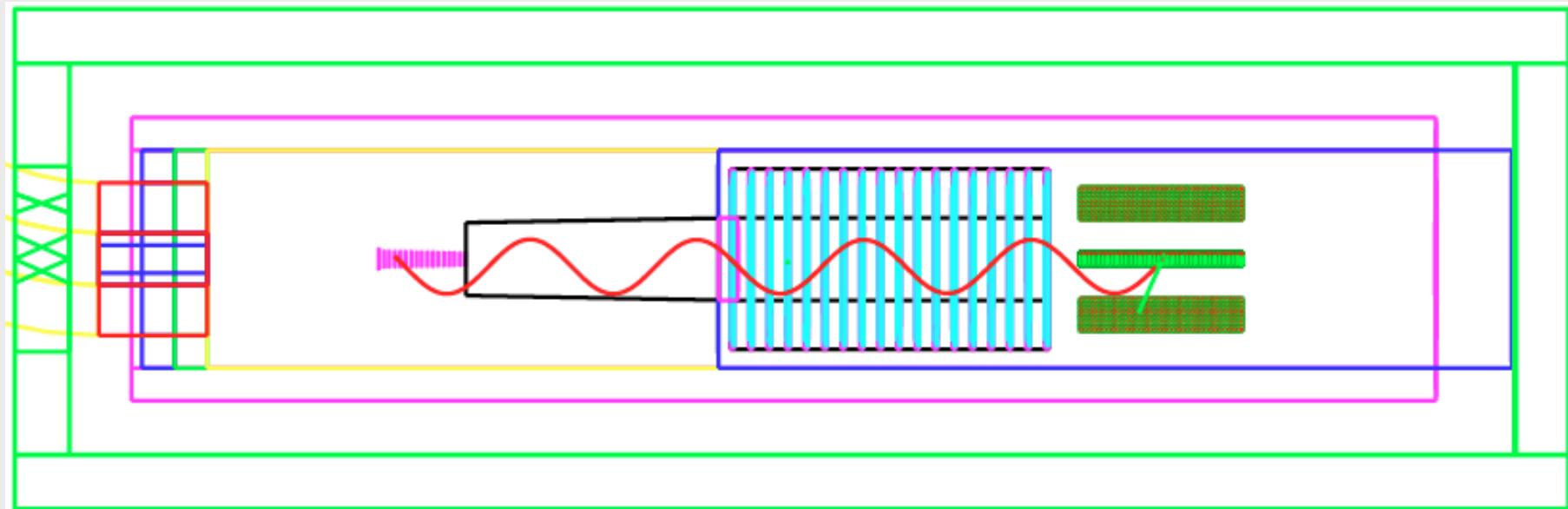
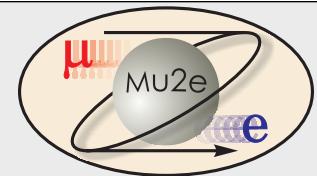
low momentum particles and  
almost all DIO background  
passes down on

10 m × 0.95 m





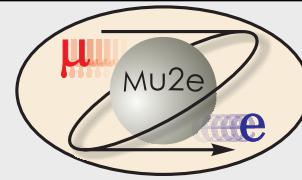
# Detector



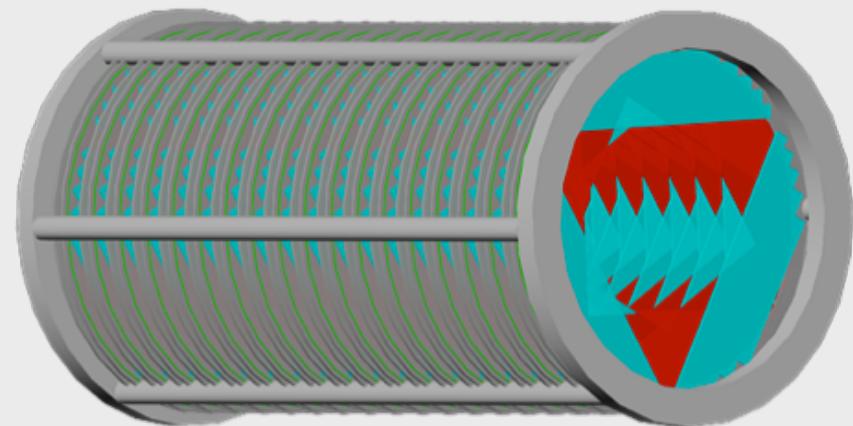
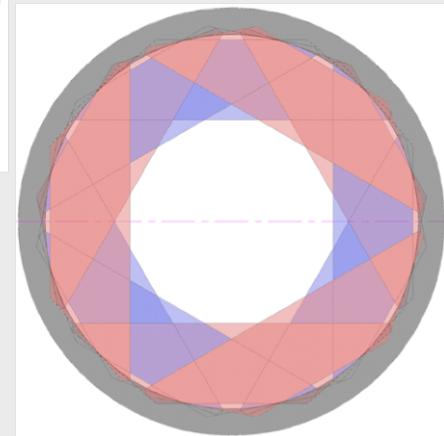
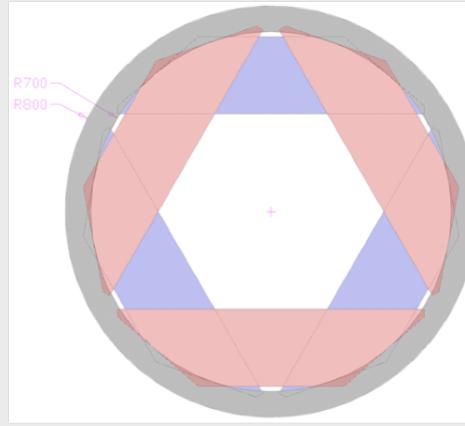
- Immersed in solenoidal field, so electrons follow near-helical path
- Conversion Electron born in Stopping Target
- Tracker followed by Calorimeter
- Tracker: (straw tubes with axes transverse to beam)
  - 21,600 straws
    - 18 stations of 5 mm diameter conducting straws
    - length from 33-118 cm
- Calorimeter:
  - 1024  $3.5 \times 3.5 \times 12$  cm PbWO<sub>4</sub> or LYSO
  - 4--5% resolution



# T-Tracker Details

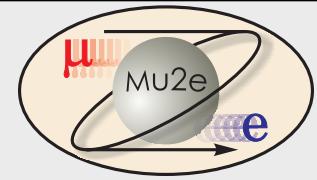


- This “windowframe” is repeated, rotated each time
- Just know which straw is hit
  - Makes PR complicated!
- Using time division  $\sim 8 \text{ cm } \sigma$  to give  $(r, \phi)$  and seed KF
  - very successful!



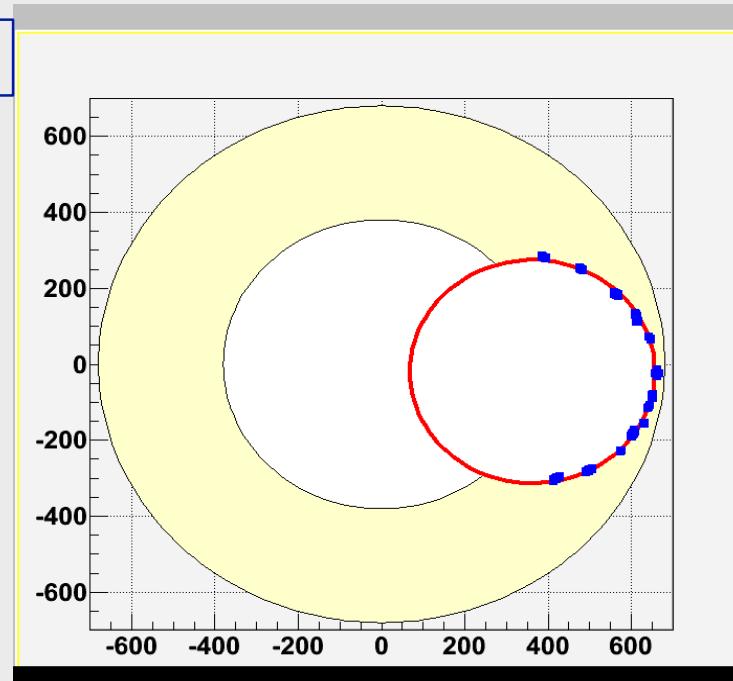


# Pattern Recognition and Tracking

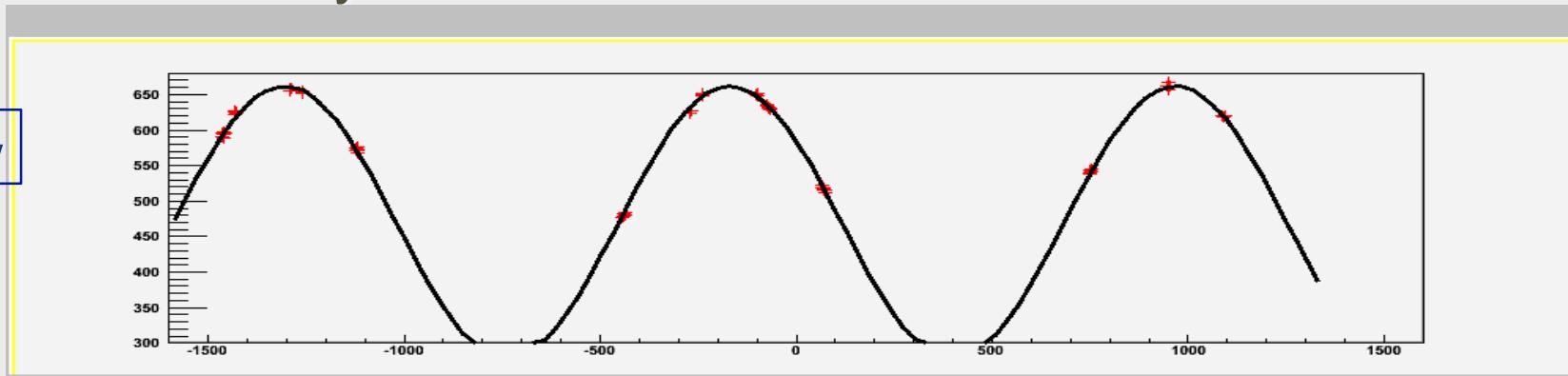


- Using Hough Transform for Pattern Recognition
- Have ported BaBar Kalman Filter for Track Fitting
  - robust against noise so far
  - extensive misreconstruction studies underway

X-Y View



R-Z view



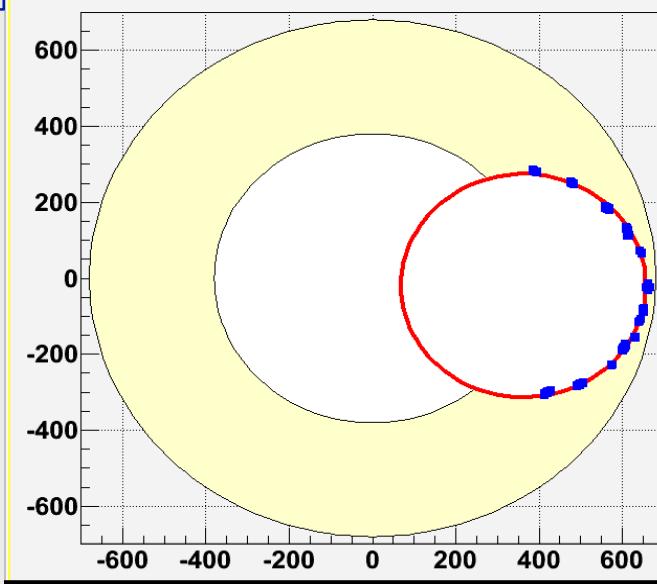
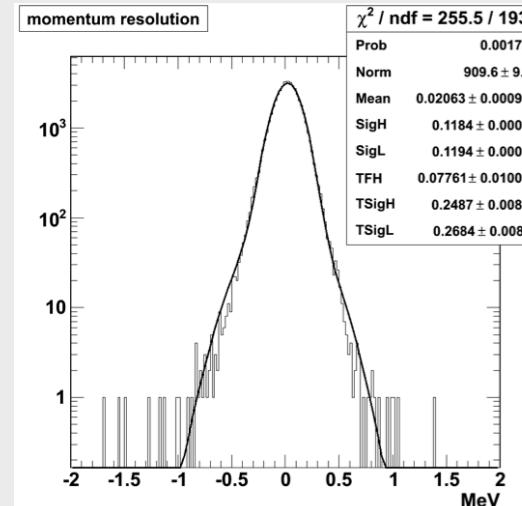


# Pattern Recognition and Tracking

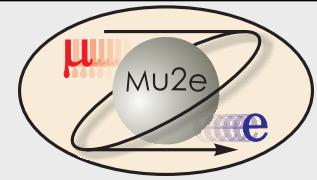
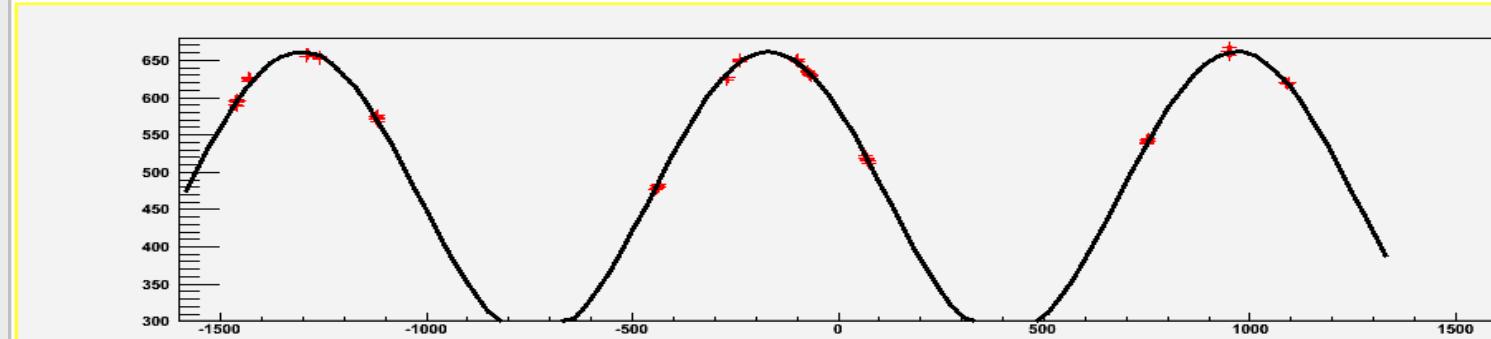
- Unified simulation from protons on target to end of detector in GEANT4
- Pattern Recognition Program joined to BaBar Kalman Filter
  - No sign of significant mis-reconstruction background with realistic rates from all known sources

work in progress! join here...

X-Y View

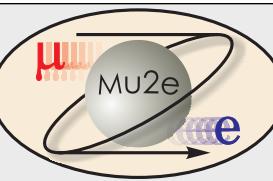


R-Z view

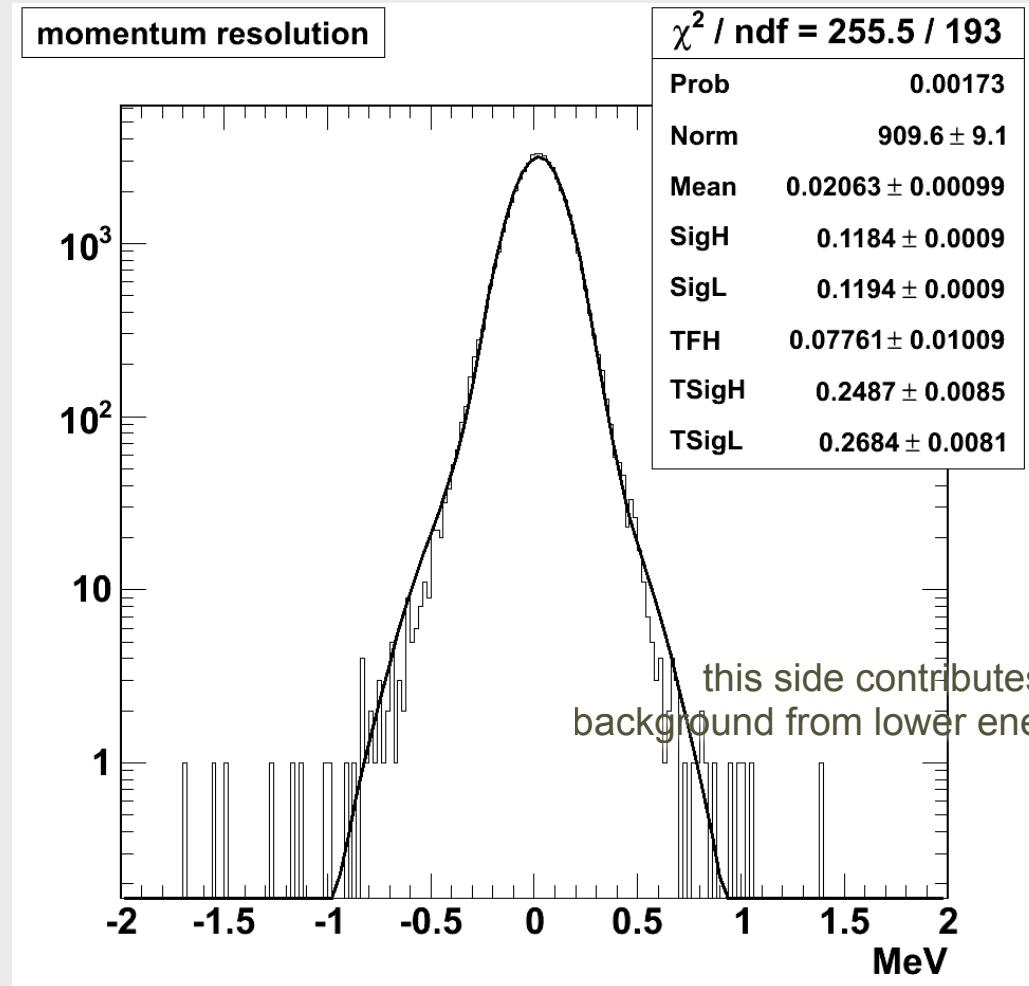


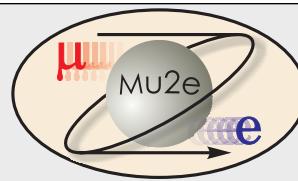


# Understanding Resolution

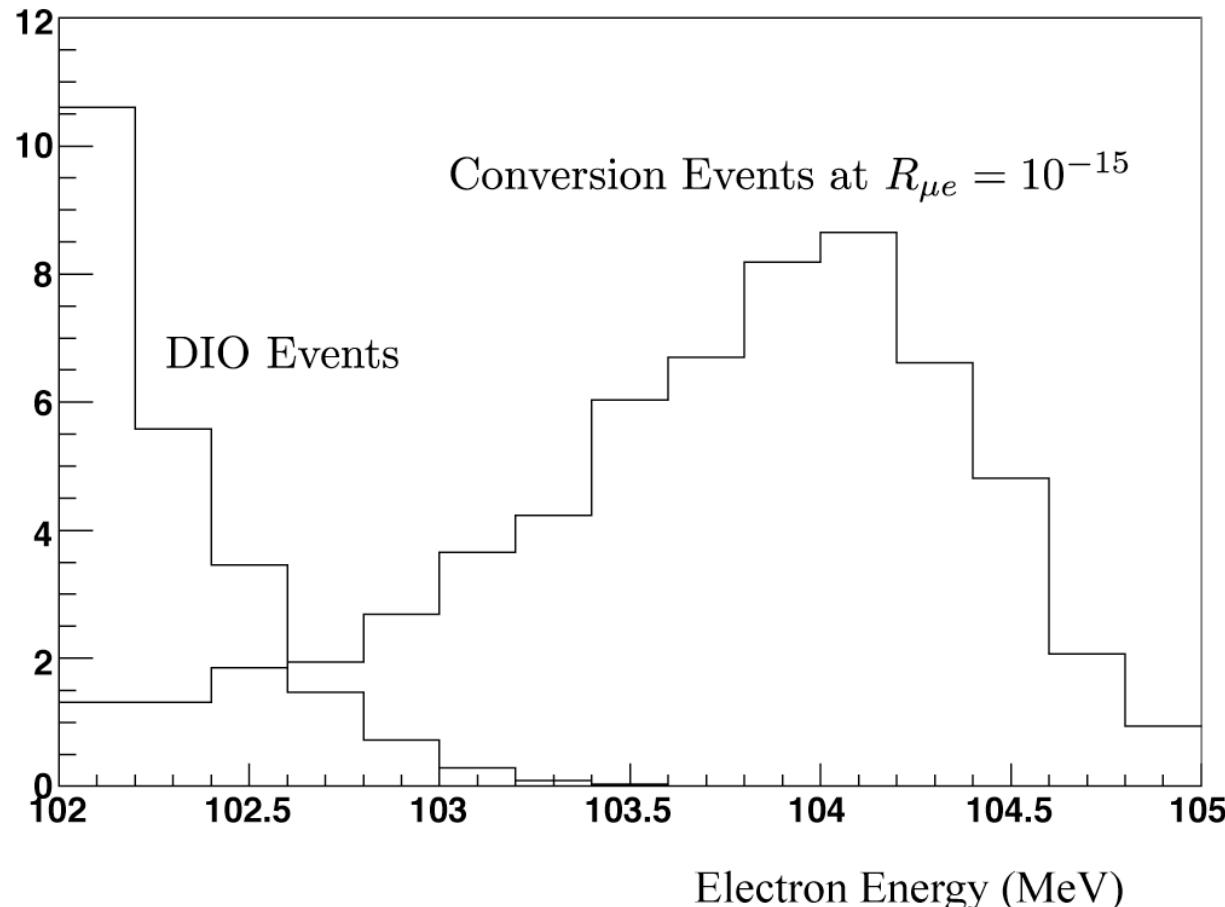


- Measure resolution/check acceptance:
  - special runs varying target foils, field, location of stopping target
  - Use  $\pi^+ \rightarrow e\nu$  decay: monochromatic line at  $\sim 70$  MeV
  - Gaussian part yield  $\sim .035$  DIO smeared upwards

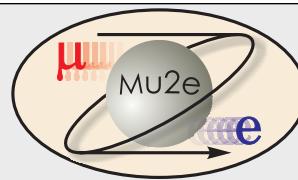




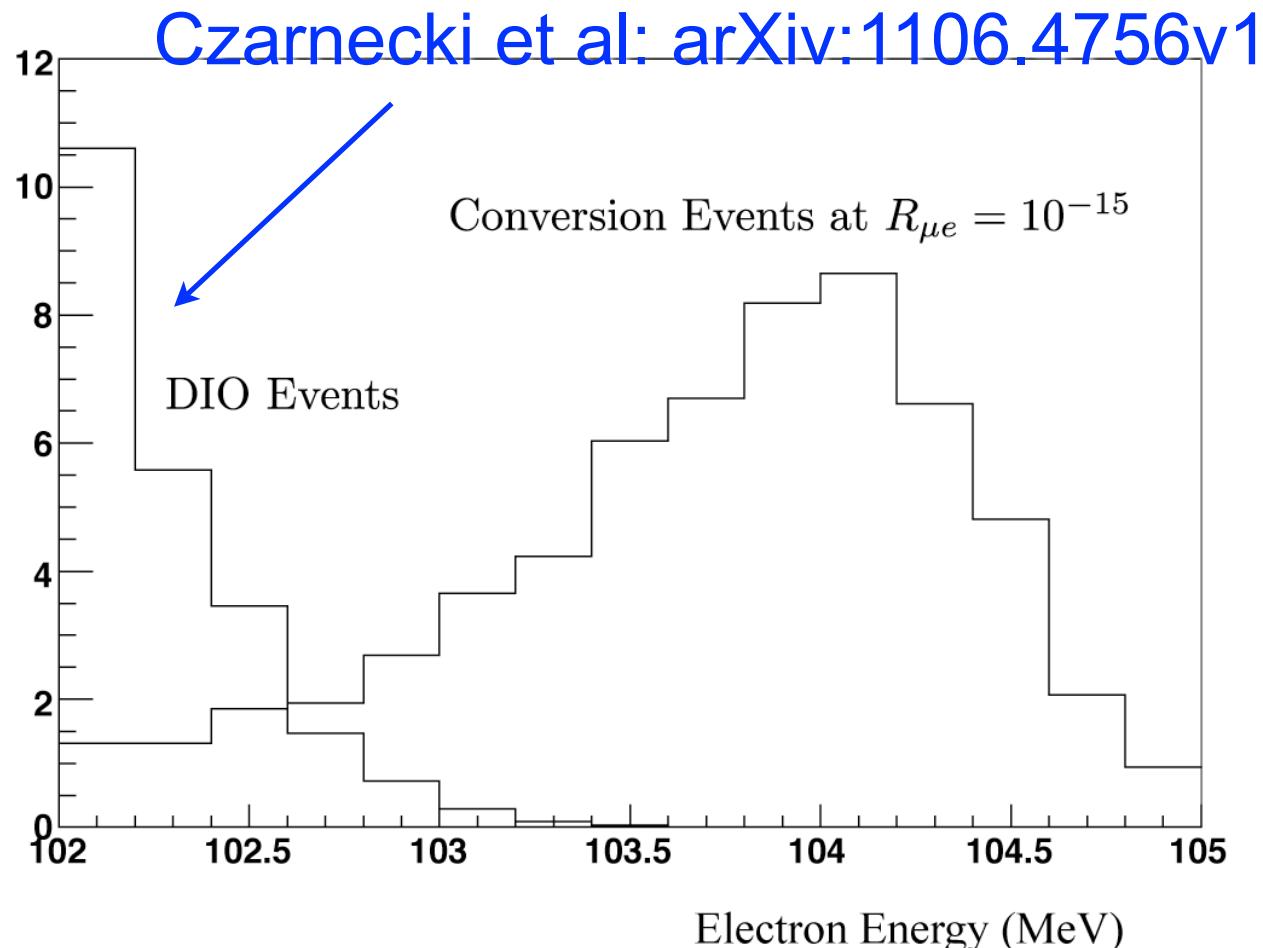
# Signal and Background



energy loss in stopping target and other material shifts  
electron down to  $\sim 104$  MeV



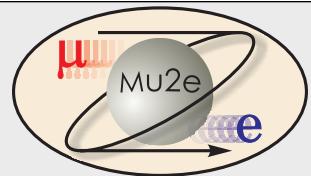
# Signal and Background



energy loss in stopping target and other material shifts  
electron down to  $\sim$ 104 MeV



# Final Backgrounds

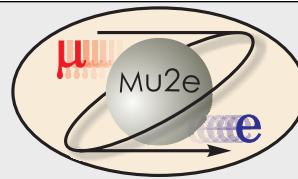


- For  $R_{\mu e} = 10^{-15}$   
~40 events / 0.18 bkg  
(LHC SUSY?)
- For  $R_{\mu e} = 10^{-16}$   
~4 events / 0.18 bkg

Source	Size	Error (not symmetric)
RPC ( $\bar{p}$ -induced)	0.06	$\pm 0.060$
RPC ( $\pi$ -induced)	0.04	$\pm 0.020$
CR	0.025	$\pm 0.025$
DIO	0.035	$\pm 0.018$
$\mu$ induced in-flight	0.01	$\pm 0.0005$
$\pi$ induced in-flight	0.003	$\pm 0.0015$
Scattered $e^-$	0.0006	$\pm 0.0003$
Radiative $\mu$ Capture	$< 2 \times 10^{-6}$	—
<b>TOTAL</b>	<b>0.18</b>	<b><math>\pm 0.08</math></b>



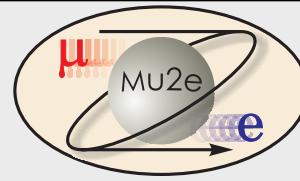
# Outline



- The search for muon-electron conversion
- Experimental Technique
- *Fermilab Accelerator*
- Project X Upgrades and Mu2e
- Cost and Schedule
- Conclusions



# FNAL Beam Delivery



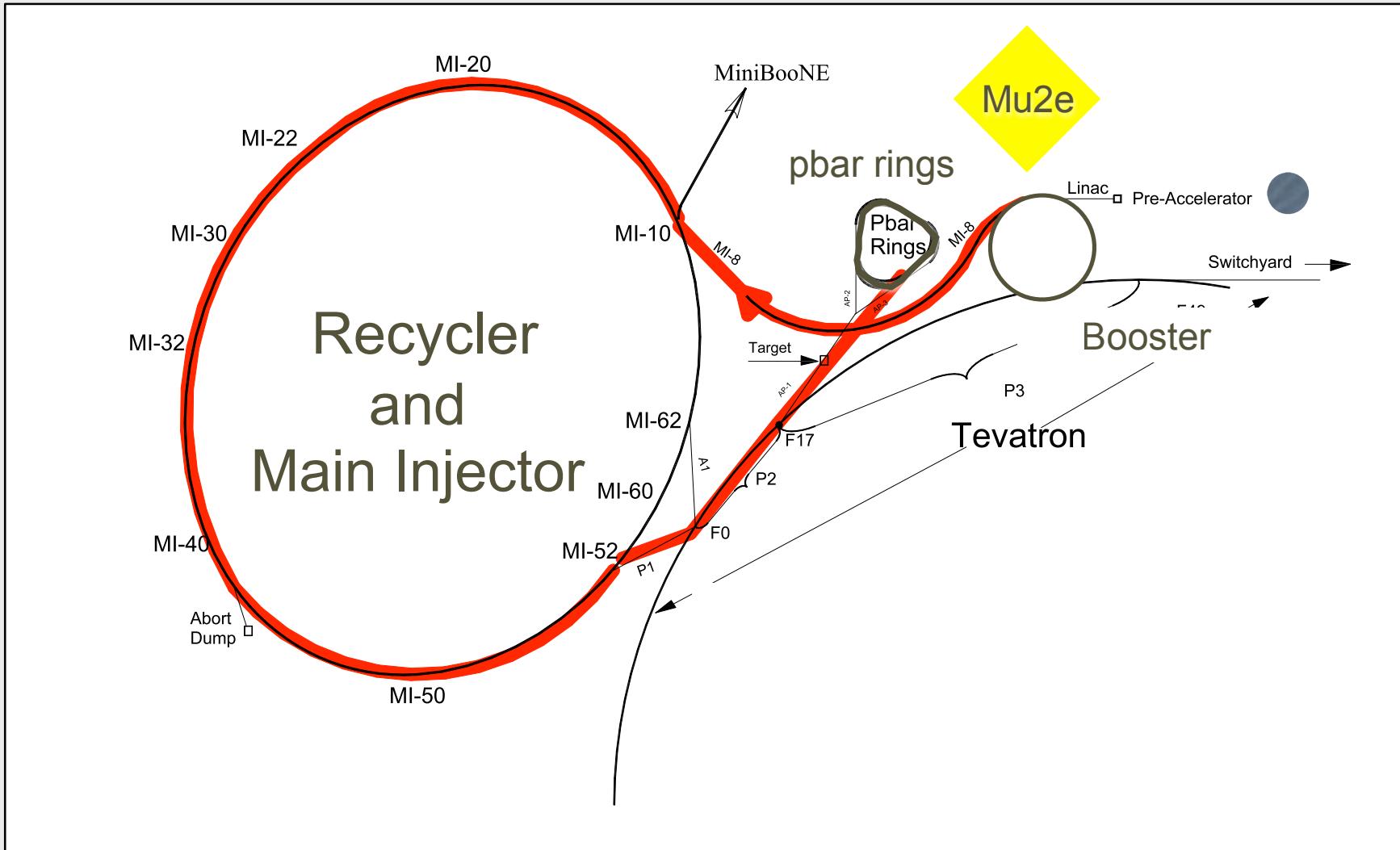
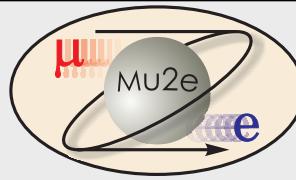
- FNAL has unique, major strength:

## Multiple Rings

- *no interference* with NOvA neutrino oscillation experiment
- reuse existing rings with only minor modifications

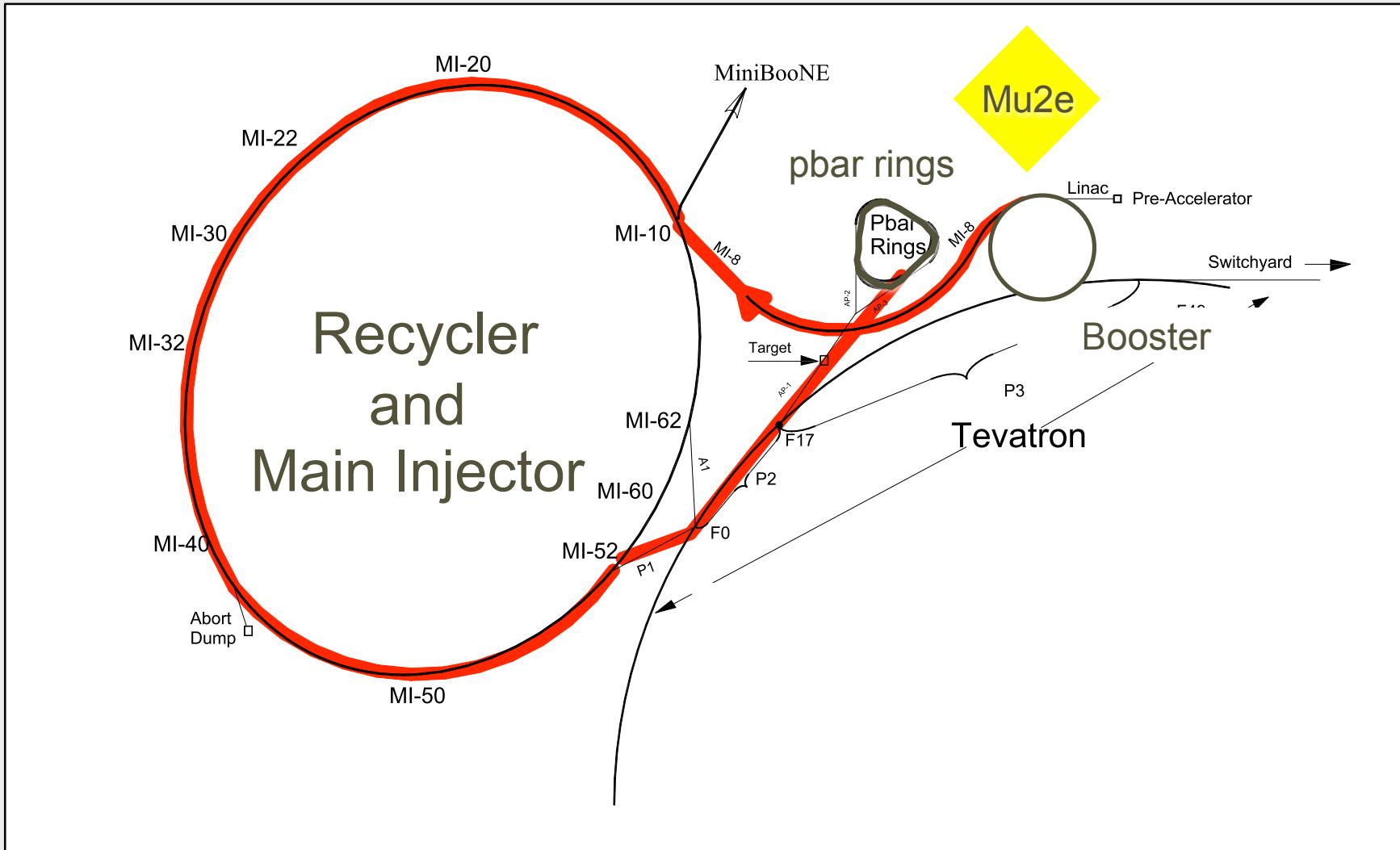
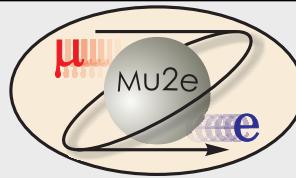


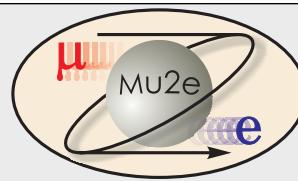
# “Boomerang” Scheme





# “Boomerang” Scheme





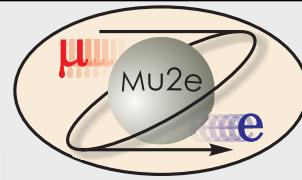
# And on the Civil Construction



Figure 2 –Site Photo Indicating Location of Proposed mu2e Conventional

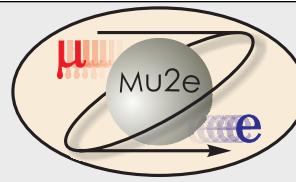


# And on the Civil Construction





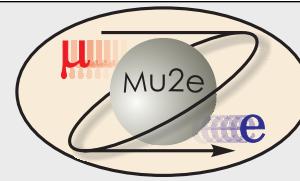
# Outline



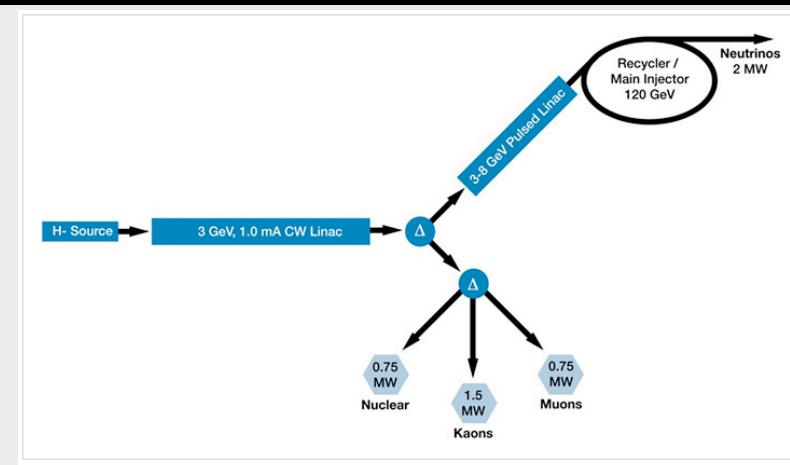
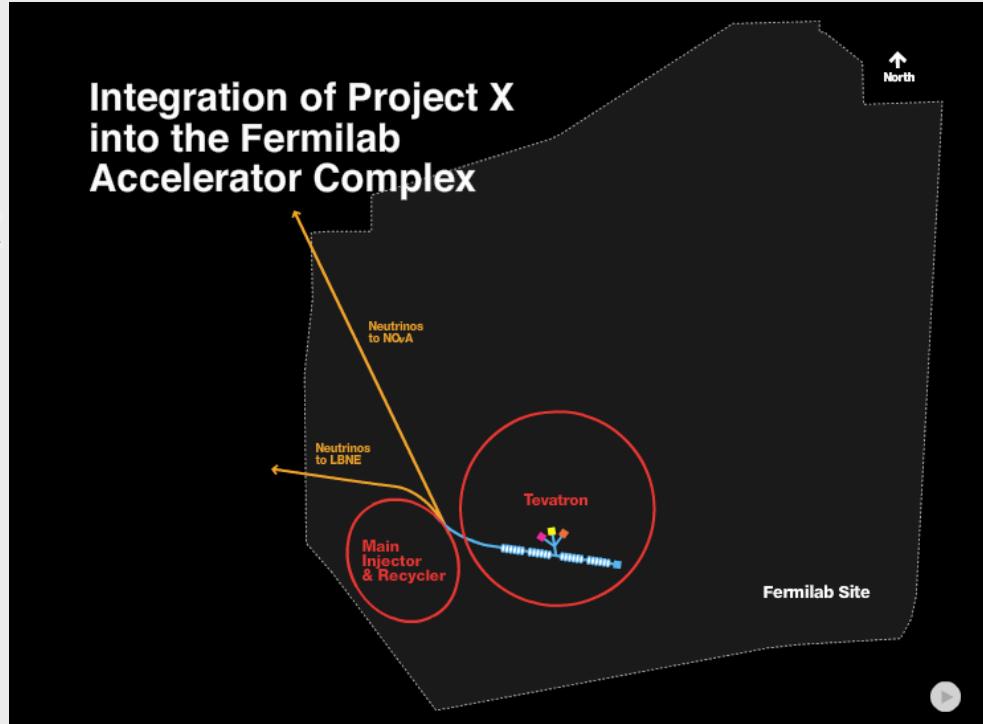
- The search for muon-electron conversion
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# Upgrades at Project X

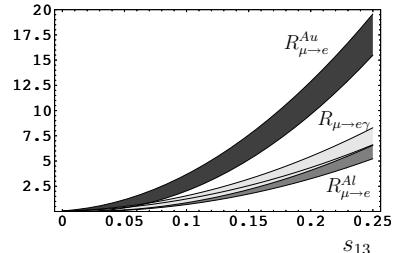
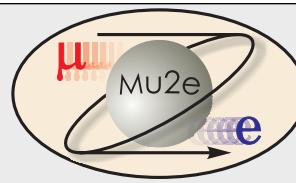


- Project X is a concept for an intense 8 GeV proton source that provides beam for the Fermilab Main Injector and an 8 GeV physics program.
- Can drive next generation experiments in intensity frontier physics: rare processes, neutrinos
- Potential to upgrade Mu2e by  $\times 100$ 
  - *study new physics*
  - *set stronger limit*





# Upgrade Plans...

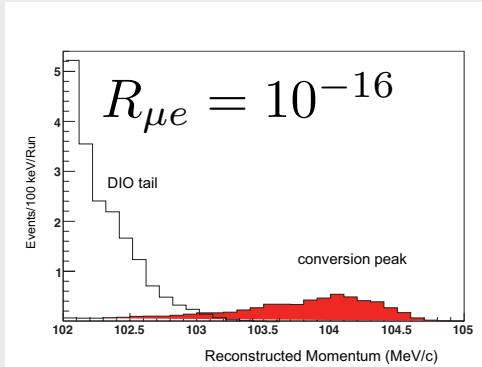


Yes

Signal?

No

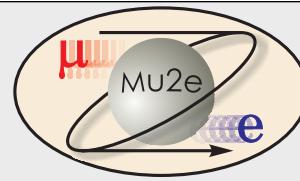
1. Change Z of Target to determine source of new physics
2. Prompt Rates will go up at higher Z, have to redesign detector and muon transport



1. Both Prompt and DIO backgrounds must drop to measure  $R_{\mu e} \sim 10^{-18}$
2. Detector, Muon Transport, Cosmic Ray Veto, Calorimeter



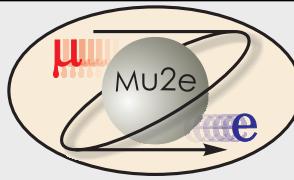
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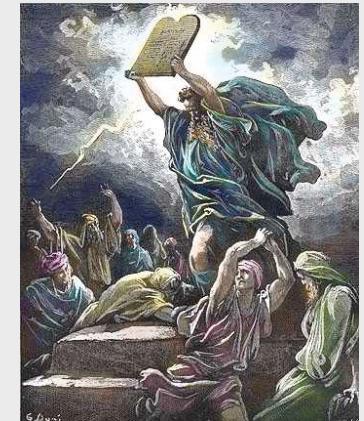
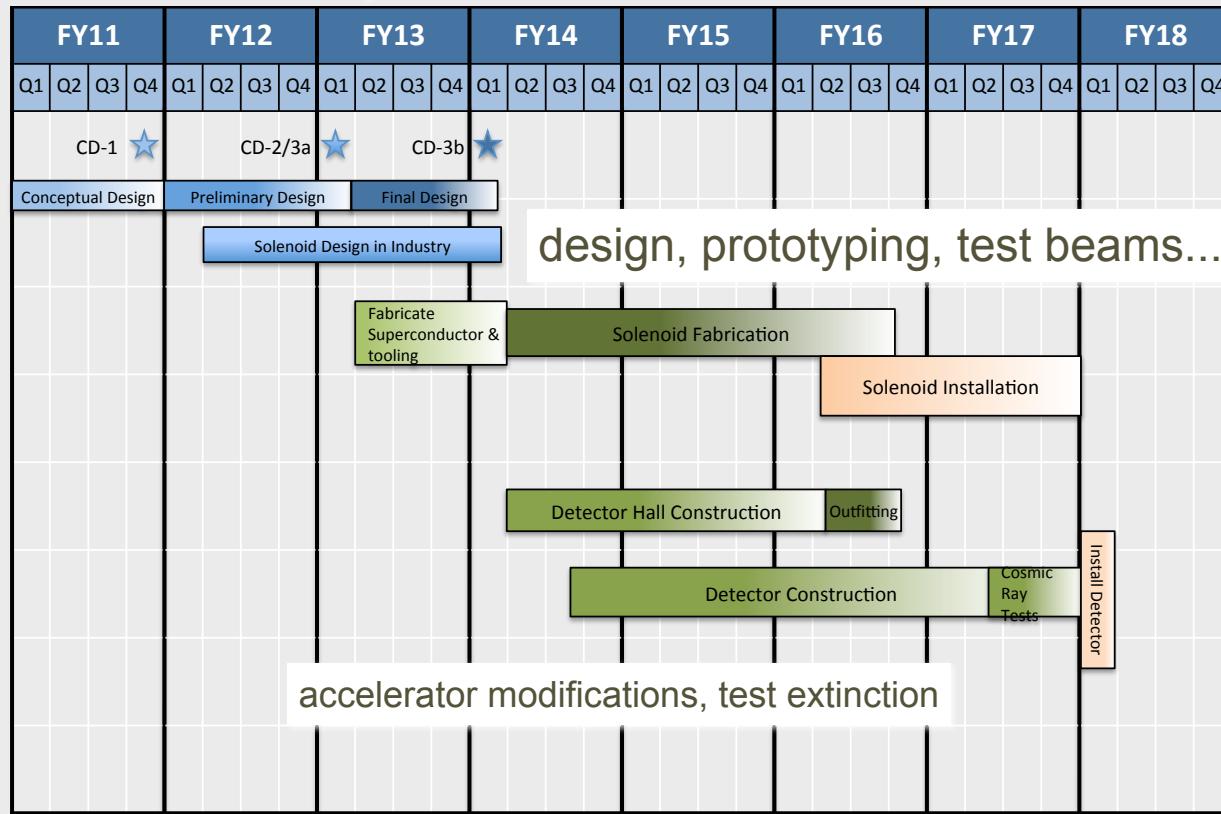
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# Cost and Schedule

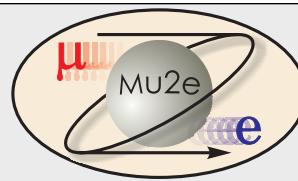


- *This is a technically limited schedule* data-taking 2018
- Critical Path is Superconducting Solenoids
- \$200M “fully-loaded” Total Cost at CD-0

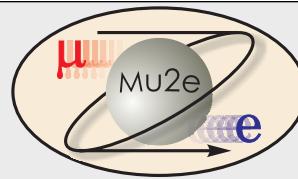




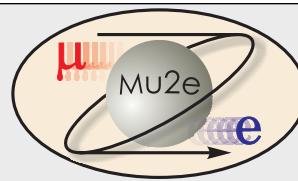
# Conclusions



- Mu2e will either:
  - *Reduce the limit for  $R_{\mu e}$  by more than four orders of magnitude ( $R_{\mu e} < 6 \times 10^{-17}$  @ 90% C.L.)*
  - *Discover unambiguous proof of Beyond Standard Model physics and*
  - *Provide important information either complementing LHC results or probing up to  $10^4$  TeV mass scales*
- With upgrades, we could extend the limit by up to two orders of magnitude or study the details of new physics



And Perhaps Answer Rabi's Question  
about the physics of flavor and generations



# And Perhaps Answer Rabi's Question about the physics of flavor and generations



Who ordered that?